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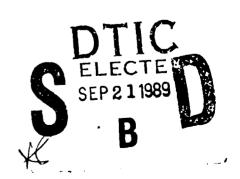
NRL Memorandum Report 6497

Catalog of Absolutely Calibrated, Range Normalized, Wideband, Electric Field Waveforms from Located Lightning Flashes in Florida: July 24 and August 14, 1985 Data

J.C. BAILEY AND J.C. WILLETT

Atmospheric Physics Branch Space Science Division

August 1, 1989



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The electric field and its time derivative from located intra-cloud and cloud to ground lightning discharges were measured in Florida over salt water paths with flat plate antennas mounted on the roof of a metallic trailer. Also recorded was the HF energy spectral density									
within	within a 1.45 µsec time window at either 5, 10, or 15 MHz averaged over a 1 MHz bandwidth.								
In this report we present range-normalized waveforms of all located lightning events in two storms occurring on 07/24/85 and 08/14/85. These events are classified into 15 categories									
according to previously identified electric field waveforms. Statistical and spectral analyses of these data will be presented in forthcoming papers. One result became evident									
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CATALOG OF ABSOLUTELY CALIBRATED, RANGE NORMALIZED, WIDEBAND, ELECTRIC FIELD WAVEFORM FROM LOCATED LIGHTNING FLASHES IN FLORIDA: JULY 24 AND AUGUST 14, 1985 DATA

I. <u>Introduction</u>

Understanding the electromagnetic signals produced by lightning is important both scientifically, to learn about the physics of the discharge, and practically, to define the electromagnetic hazard. The broad-band (i.e., time-domain) electric and/or magnetic field impulses tend to have characteristic and reproducible waveforms that can be used to identify different processes within the complex phenomenology of a cloud-to-ground (CG) or intra-cloud (IC) lightning flash (e.g.; Kitagawa and Brook, 1960; Lin et al., 1979; Weidman and Krider, 1978, 1979). The identification of these processes is invariably inexact, however, due to borderline cases and processes whose physical characteristics are unknown. Therefore, a compilation of many examples of each type of process is useful in developing a more consistent classification. This compilation should also lead to a better understanding of the composition of both CG and I freshes.

Field experiments were conducted at Kennedy Space Center (KSC),

Florida, during the summers of 1984, 1985, and 1987 to obtain calibrated,

wideband, radiation field waveforms and spectra of impulsive events within

the lightning flash. These events include first return strokes and various

types of leaders, subsequent return strokes, and cloud pulses. The 1984

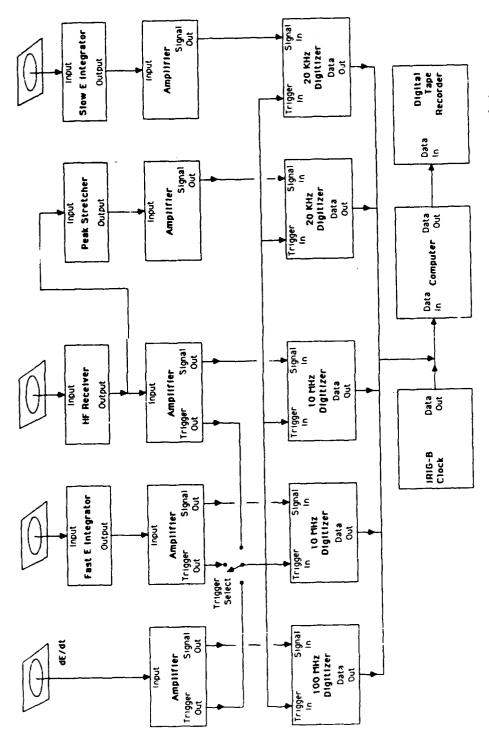
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experiment used relatively simple equipment and has been described in detail by Leteinturier et al. [1985]. They discussed the statistics of the peak electric field and its derivative, the energy spectral density (ESD), and the implied mean peak current derivative of various events within one Florida thunderstorm. A more complex, fully digital data-acquisition system was used to record vertical electric fields at ground level from lightning during the summers of 1985 and 1987. Bailey et al. [1988] have discussed the statistics and spectra for some of the 1985 data. The 1987 data will be analyzed in future reports.

In this report we present a collection of 15 different classes of CG and IC lightning events for the days July 24 and August 14, 1985. The event type is identified by the general characteristics of each waveform. Future reports, for additional storms, may be expanded to include additional classes as the analysis progresses. All lightning waveforms given herein except the "slow E" (see sect. III) have been range normalized to 100 Km (1/R dependence), which facilitates comparison of the radiation fields among different flashes. A simple procedure based on HF radiation is used to test whether "characteristic pulses" (defined later) precede all first return strokes as suggested by Beasley et al. [1982].

II. <u>Instrumentation</u>

A block diagram of the 1985 apparatus is shown in Fig. 1. This system was designed to record an entire lightning flash with moderate time resolution, while also capturing several of the impulsive events within that flash with much higher resolution. High resolution data were obtained from a group of digitizers sampling the "fast electric field" (fast E) and HF



dE/dt signal in parallel but at a different gain to give better dynamic range. The digital tape recorder was also replaced with a second computer to display the data in real time and record it to replaceable 20 MB disk The trigger selector was set to the output of the HF receiver for the data reported here. In 1987 a second 100 MHz digitizer recorded the FIG. 1 -- Block diagram of the electric field recording system used in cartridges. 1985.

ESD (fast HF) at 107 samples per second and the time derivative of electric field (dE/dt) at 10^8 samples per second. The E antenna system had a 10 to 90% risetime of about 160 ns and a decay time constant of 1 ms. The HF receiver measured, as a function of time, the ESD within a time window of effective width 1.45 μ sec, averaged over a 1 MHz bandwidth about the center frequency. The dE/dt antenna was similar to that described by Weidman and Krider [1980] and had a bandwidth of better than 30 MHz. For the data reported here, the system was triggered by the output of the HF receiver, which could be tuned to any center frequency between 3 and 18 MHz. data from these high resolution digitizers were read into a computer memory and the digitizers re-armed in about 100 ms. Up to seven such bursts of high resolution data, each time-stamped to an accuracy of ±2 ms, could be recorded during a single lightning flash. The duration of the "flash" was defined by two slow digitizers that sampled the "slow electric field" (slow E) and the peak HF ESD (slow HF) at 2 X 104 samples per second for a period of 1.075 s starting 403 ms before the first trigger. The slow E antenna had a time constant of 10 s.

In 1987 the recording system was improved by reducing the dead time between high resolution events to 13 ms and by increasing the maximum number of bursts to 24. A second 100 MHz digitizer recorded the dE/dt signal in parallel but at a different gain to give better dynamic range. The slow sampling rate was also increased to 5 X 10⁴ samples per second, and the duration of a "flash" raised to 1.31 s, without changing the basic configuration of the system.

Four 0.2 m^2 flat plate antennas were centered on the roof of a 5.5 m X 2.4 m X 2.4 m grounded metal trailer. In 1985, the trailer was connected

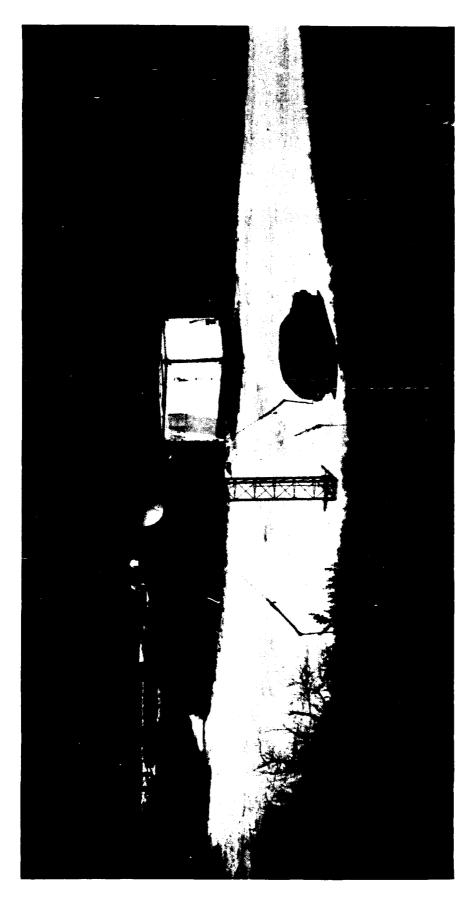
to a single ground rod with a DC ground resistance of 0.25 Ω . In 1987 six rods grounded the trailer, one at each corner and one in the center of each long side. The DC ground resistance of each rod was less than half an ohm.

During both 1985 and 1987, the instrumentation trailer was located just behind a row of low dunes on the ocean beach near KSC camera site 10. A photograph of the site is given in Fig. 2. The trailer was only 44 meters from mean high water, so that the propagation path for signals from lightning over the Atlantic Ocean was almost entirely over salt water.

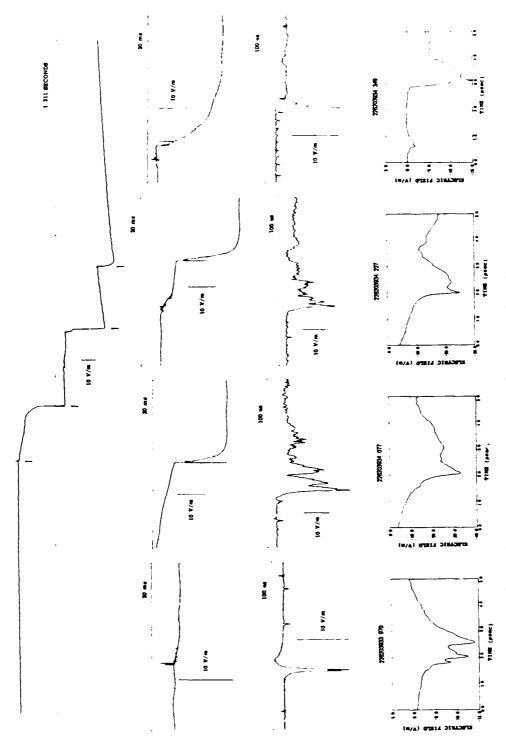
The field enhancement factor for the antennas on the roof was determined to be 2.25 (1985) and 2.15 (1987) by comparing their signals with those of an identical antenna on flat ground nearby. This comparison was accomplished at night, using waveforms with little spectral power above 50 KHz, produced by distant return strokes, because the enhancement factor was found to be independent of the azimuth of incidence at these frequencies. Angular anomalies observed at higher frequencies are believed to be due to coupling into the long cable between the ground antenna and the electronics inside the trailer. The overall absolute accuracy of the peak electric field measurements is estimated to be ± 6%.

A comparison of dE/dt spectra from the roof antennas with those from a similar antenna on flat ground, connected to the trailer by fiber optics, (Bailey and Willett, 1987) showed no serious resonances in the HF band for the 1987 configuration. The composite return-stroke spectra in 1985 and 1987 were very similar, which suggests that the previous grounding configuration did not produce resonances or contaminate the 1985 data either.

Lightning locations were obtained from a three-station network of gated magnetic, direction-finders (manufactured by Lightning Location and



the Atlantic Ocean for the 1987 experiment. A microwave transmitting tower is in the foreground. The 1985 location was nearly identical except that FIG. 2 -- Photograph showing the location of the trailer on the beach of the microwave tower was absent. The view is approximately ENE.



of 1.31 s. The second through fourth sets of waveforms correspond to slow E, fast E, and integrated dE/dt, respectively. They are expanded electric field signatures of the events indicated by the arrows in the first row on time scales of 20 ms, 100 μ sec, and 5 μ sec, respectively. This flash has on August 14, 1987. The top waveform is the slow E record for a duration FIG. 3 -- Electric field change data for a CG flash occurring at 20:39:34 not been range normalized.

Protection, Inc., Tucson, Arizona - Binford et al., 1983) that was operated for KSC by the Cape Canaveral Air Force station. This system was designed to locate only the ground-strike points of the first return strokes in normal CG flashes, i.e., those transferring negative charge to earth. For this study the raw data from the individual direction finders were reanalyzed with the application of appropriate site-error corrections (Krider, personal communication).

III. DATA AND PLOT FORMAT

Comparison of the radiation field magnitudes from lightning discharges requires knowledge of the discharge location. This requirement limits the data to either CG flashes or IC flasher occurring in a localized storm so that the range can be estimated. Therefore most of the data in this report is collected from CG flashes.

The CG flash (Uman, 1987, pp. 10-17) usually includes preliminary variations ending with a sequence of "characteristic cloud pulses" (Beasley et al., 1982) followed by a "stepped leader" and "first return stroke". "Subsequent return strokes" of several types may then follow. For example, the CG flash (Fig. 3) recorded at 20:39:34 on August 14, 1987, contains a characteristic pulse followed by two stepped leader and first return stroke pairs. They are followed by a subsequent return stroke preceded by a "dart-stepped leader". Preceding the initial first return stroke in the slow E record is a good example of the electrostatic field change due to transfer of charge by the relatively slowly (22 X 105 m/s) propagating stepped leader. Subsequent return strokes exhibit much briefer leader field changes because the propagation speeds of both dart (23 X 106 m/s)

and dart-stepped ($\simeq 1 - 2 \times 10^6$ m/s, V.P. Idone, personal communication) leaders are faster.

Multiple stroke flashes have more than one channel to ground approximately 44 % of the time (Thompson et al., 1984 - Fig. 2) suggesting that more than one first return stroke, as identified by the radiated field waveform, can occur within a single flash. Therefore, location of one stroke within a flash does not necessarily imply that all the other events have the same location. In order to obtain the most precise locations possible for range-normalization, we rejected events which might have occurred in channels to ground different from the one located by the direction finding system.

A stroke was considered located primarily by time coincidence with the magnetic-direction finders to an accuracy of approximately ± 5 ms. Other events in the flash were taken to have the same location as this stroke unless there was evidence that another first return stroke occurred in the flash. This criterion most often limited the number of located first return strokes in a flash to one. It also frequently forced us to throw out subsequent strokes and other events in flashes containing more than one first stroke. One case (see flash at time 17:31:04.283 on 08/14/85) with more than one first return stroke has been included, however. For this case, all three direction finders counted 11 or 12 strokes in the flash (within the 2.5° azimuth window for subsequent strokes). Because the slow E field change exhibited at least 11 steps, the ranges of two first return strokes within this flash were deemed known to sufficient accuracy.

Data were collected from two storms over the Atlantic Ocean, which occurred on July 24 (Fig. 4) north of the trailer and August 14, 1985

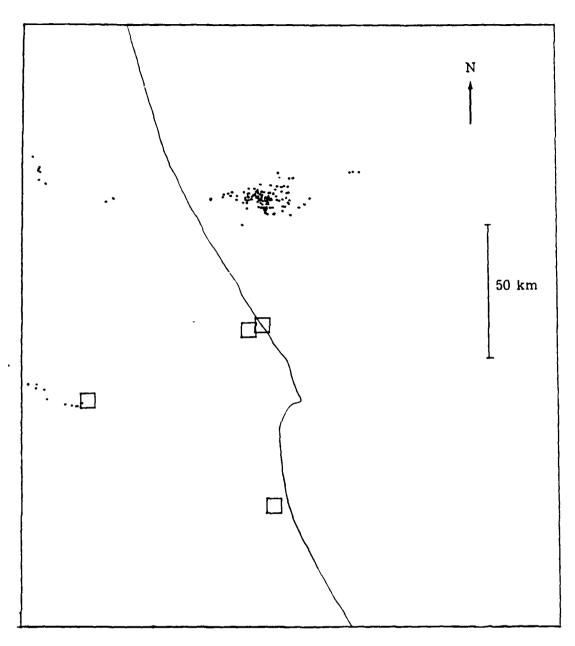


FIG. 4 -- Map showing the ground-strike points of CG flashes located by the magnetic direction-finding network during the period 20:25 - 21:00 UT on 7/24/85. The north-eastern-most square represents the location of our instrumentation trailer on the east coast of the Florida peninsula, and the other squares represent the three direction finders in the network.

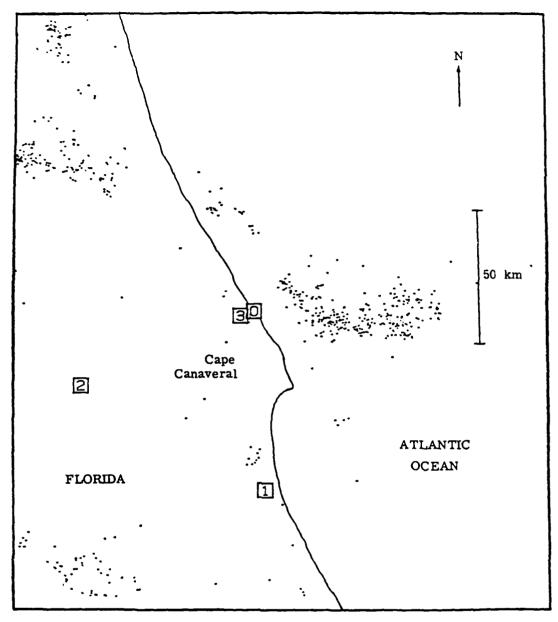


FIG. 5 -- Map showing the ground-strike points of CG flashes located by the magnetic direction-finding network during the period 17:22 - 18:42 UT on 8/14/85. The squares have the same representation as in FIG. 4.

(Fig. 5) east of the trailer. These storms produced lightning at ranges from 40 to 60 km and 15 to 50 km, respectively. The ground-strike locations, discussed elsewhere for the 7/24/85 storm by Krider [1988], are believed to be accurate to ± 3 km or better. The propagation path for signals from both storms was almost entirely over the ocean, largely eliminating the severe high frequency attenuation that can occur because of signal propagation over poorly conducting land (Wait and Walters, 1963).

The HF receiver frequencies and the trigger thresholds, in $(V/m/Hz)^2$ at the trailer location, for both data sets are given in Table 1. Each table entry is correct up till the time of the next entry. The thresholds must be range normalized to 100 Km by multiplying by $(Range/100)^2$ for comparison with the data plots.

Table 1 -- HF center frequencies and trigger thresholds

Date	Time	Freq. (MHz)	Trigger Level (V/m/Hz) ²
07/24/85	20:06:00	5	1.87*10-14
, ,	20:26:00	10	4.68*10-15
	20:27:15	10	7.42*10-16
	20:47:00	15	3.30*10-16
	20:47:15	15	6.58*10 ⁻¹⁷
08/14/85	17:20:50	5	5.92*10 ⁻¹⁴
	18:20:00	5	2.97*10-14

The plots to follow are identified by event times rather than figure numbers. Every flash is plotted using identical formats. High resolution data for each located trigger event within a flash is represented by a "fast page" containing four plots. There is one "slow page" per flash containing two plots per page. More specific descriptions of the fast and slow pages are discussed below. The magnitude of the linear vertical scales among plots can vary and is chosen to give the best display of the

data. When digitizers are saturated, the saturation levels are indicated on the plots by dashed lines. The sign convention is that of conventional physics: a positive electric field at the ground is produced by deposition of negative charge overhead. A normal return stroke that lowers negative charge thus produces a negative field change. This sign convention is opposite to that used in much of the published literature.

Each fast page has the following format. On the left side, the fast E (V/m) waveform is plotted above the fast HF $(V/m/Hz)^2$ waveform, both at a resolution of 100 ns per sample for a total time of .4096 msec. On the right side, the portion of the fast E waveform coincident with dE/dt is expanded and centered above the dE/dt $(V/m/\mu sec)$ waveform which is sampled at 10 ns for a total time of 40.96 μsec . The fast page title contains a date/time group followed by the event definition and range in Km (in parentheses). The date/time group is the time of the trigger. For example, the date/time group 226181216.300 can be decoded as day 226, hour 18, minute 12, and second 16.300. Universal time (UT) is used throughout. Subtract four hours to convert to Florida daylight savings time.

The fast pages for a given flash are followed by a slow page showing the context of each event within the entire flash. The slow E (V/m) waveform is plotted above the slow HF $(V/m/Hz)^2$ waveform, both at a resolution of 50 μ sec per sample for a total time of 1.075 sec. The trigger times for the fast events are indicated by tick marks beneath the slow HF record. The slow title contains the start time of the record followed by the date and the range (in parentheses). During the July 24 storm, 60 Hz pickup dominates the slow E signal. This is caused by the increased gain needed for the low signal levels from the relatively distant lightning discharges.

Because of the nature of the fast and slow E integrators, there is no absolute electric field zero. For this reason, and to remove any electronic zero drifts from the data, zeros were computed numerically from each record. The fast and slow HF zero levels were set to the minimum value of the respective HF record. The slow E zero is equated to the average of the maximum and minimum values. The other two zero levels were determined by dividing the full records into four quadrants of identical length. The fast E zero level was equated to the average of the first quadrant whereas the dE/dt zero level was taken as the average of the first and fourth quadrants.

Assuming that the waveforms are primarily radiation field and hence retain their fundamental shape as a function of distance, all plots except the slow electric field (mostly electrostatic) have been range normalized to 100 Km. For some of the closest flashes, this assumption may be invalid. The original data can be reconstructed by multiplying the fast E and dE/dt waveforms by 100/Range, multiplying fast and slow HF by (100/Range)², and leaving the slow E alone.

IV. <u>DEFINITION OF EVENT TYPES</u>

A. General definitions

The content of the data analyzed so far has required the definition of several categories of return strokes, leaders, and cloud pulses. Future reports may add other definitions. Most of categories are based on observations of earlier investigators. The 15 categories, with the abbreviations used in the plots in parenthesis, are as follows:

First return strokes (1st RS)

Dart-stepped subsequent return strokes (DS subs RS)

Normal subsequent return strokes (Norm subs RS)

Chaotic subsequent return strokes (Chaotic subs RS)

Anomalous #1 return strokes (Anomalous #1 RS) - jagged then smooth Anomalous #2 return strokes (Anomalous #2 RS) - smooth then jagged Multiple ground return strokes (Multiple gnd RS)

Stepped leaders (Stepped leader)
Dart-stepped leaders (DS leader)
Dart leaders (Dart leader)
Chaotic leaders (Chaotic leader)

Characteristic cloud pulses (Char cloud pulse)
Narrow positive bipolar pulses (NPBP)
Narrow negative bipolar pulses (NNBP)
Unknown cloud pulses (Unknown cloud)

Examples of most of the return stroke and leader events can be found in either Weidman and Krider [1978] or Weidman [1982]. These authors have defined first, dart-stepped subsequent, normal subsequent, and chaotic subsequent return strokes. (To be precise, subsequent strokes are preceded by either a dart-stepped, dart, or chaotic leader. In this report, subsequent return strokes are named for the type of leader preceding them.) A particular event is defined as a return stroke if the signature of the fast-rising portion of the return stroke waveform is visible in the 41 µsec dE/dt time window. If only the leader's signature is visible in the dE/dt window, then the event is classified as a leader. In addition to those identified by Weidman and Krider [1978] or Weidman [1982], three other types of return strokes - called anomalous \$1\$, anomalous \$2\$, and multiple ground return strokes - have been defined as described below.

Wideband recording systems have identified several types of radiation fields associated with cloud processes. Uman [1987 - sect. 4.3 and 13.6] has a good summary of these events. They include trains of unipolar pulses (Krider et al., 1975), bipolar pulses with structure on the rise to peak

(Weidman and Krider, 1979 and Beasley et al., 1982), and bipolar pulses with smooth rise to peak (Le Vine, 1980; Cooray and Lundquist, 1985; and Willett et al., 1989). So far, located trains of unipolar pulses have not occurred in our data. Bipolar pulses with structure on the rise to peak, if they have negative initial polarity and precede first return strokes, are called characteristic pulses. Bipolar pulses with smooth rise to peak are called NPBP's if they have initial positive polarity and NNBP's if they have initial negative polarity, as long as they are sufficiently narrow and have the characteristic dE/dt signature defined by Willett et al. [1989]. Our last category, unknown cloud pulses, extends to all cloud pulses that appear not to fall distinctly into well defined categories. Unknown cloud pulses do not include any of the types listed above.

B. Specific Descriptions

Now that the general definitions have been discussed, this section gives more detailed descriptions. All of the following descriptions are based on the fast plots. The fast electric field serves as the main source of identification for each event, with the fast HF and dE/dt playing a supporting role. The descriptions will, in general, discuss the stepping structure before the peak, how rapidly the peak is attained, and the structure after the peak for each fast waveform.

A first return stroke preceded by a stepped leader (see time 226180000.072 for an example) has an electric field signature characterized by a slowly rising initial portion (front) before a fast transition to peak followed by a jagged section a few hundred microseconds in duration. This jagged section is believed to be produced by the superposition of radiation from the multiple branches evident in most photographs of first return

strokes (Weidman and Krider, 1978). Both the HF and the dE/dt waveforms exhibit spikes, corresponding to the leader steps, preceding a sharp peak followed by a relatively noisy section.

<u>Dart-stepped subsequent return strokes</u> (226174528.924) have shorter front durations before the fast transition to peak than first return strokes. The relatively smooth section of waveform after the peak, presumably due to an absence of branches, characterizes all types of subsequent return strokes. The <u>dart-stepped leaders</u> before these strokes tend to have more regular and shorter time intervals between steps than stepped leaders (Krider et al., 1977). The fast HF and dE/dt signatures exhibit spikes similar to first return strokes, corresponding to leader steps, before the sharp peak followed, in general, by a very quiet section.

Normal (226181138.648) and chaotic (226180544.448) subsequent return strokes are similar to dart-stepped subsequent return strokes except for the leader sections. Normal subsequent return strokes exhibit no indication of spikes before the abrupt transition to peak on any of the fast waveforms, indicative of a smoothly propagating dart leader. The leader activity on the fast E record prior to chaotic subsequent return strokes is irregular with closely spaced pulses of erratic amplitude behavior. There is also evidence of this type of activity on both the fast HF and dE/dt waveforms of chaotic leaders. When the leader amplitudes are near the system noise level, the difference between dart and chaotic leaders is difficult to ascertain. The distinction between them is thus somewhat subjective, leaving open the question of whether they are really different physical phenomena.

The other three types of return strokes defined here are of a hybrid nature. They occur fairly rarely with respect to the types described

above. The waveforms of Anomalous #1 (226180000.175) return strokes start out like first return strokes but quickly become smooth, whereas Anomalous #2 (226175829.133) return strokes start out like a subsequent stroke but end up jagged. Decay of the first return stroke's channel during the interval between strokes suggests a plausible explanation for Anomalous #1 return strokes. The lower section of the channel has more time to decay than the upper section. Presumably the dart-stepped leader follows the same channel as the stepped leader at higher altitudes, without branching, but forms a new channel at lower altitudes, with branching. Because the return stroke propagates from the ground toward the cloud, this scenario would suggest a waveform similar to that seen in Anomalous #1 return strokes. The waveform of Anomalous #2 return strokes suggests branches at higher altitudes and the absence of branches at lower altitudes. The explanation for this type of waveform is not known. A multiple ground (205200917.206) return stroke exhibits two sharp transitions to peak and looks like two first return stroke waveforms superimposed. This type is assumed to originate from two almost simultaneous return strokes initiated by different branches of the same stepped leader (Guo and Krider, 1982).

Unlike return strokes, which have the same general shape among categories, cloud pulses are more varied. Beasley et al. [1982] believe that characteristic pulses (226174950.465) probably occur before all stepped leaders associated with first return strokes. They state that these bipolar pulses with negative initial polarity (our sign convention) occur in groups during a transition period of a few milliseconds between when the "preliminary variations" end and stepped leader begins. Our recordings also indicate that characteristic pulses usually occur in groups. Each

individual pulse (see the expanded fast E record) tends to exhibit several narrow spikes on a slow rise to peak followed by a smooth overshoot. NPBPs (205203928.628) are most easily distinguished by their electric field derivative, which contains high frequency oscillations superposed upon a slower variation of initially positive then negative polarity. The magnitude of the positive and negative peak is usually nearly the same. The electric field structure typically has a fast positive rise and decay followed by a small negative overshoot. NPBPs appear to be isolated events; seldom is any other slow HF peak found during the same "flash". A range of 45 Km was assumed for them because they were recorded during the isolated and localized storm on 07/24/85. NNBPs (205203429.863) are nearly identical to NPBPs except the polarity is reversed (Willett et al., 1989). Unknown cloud pulses (226173216.949) are included because they appear to be associated with the cloud to ground flash.

V. OBSERVATIONS AND CONCLUSIONS

Fast (10 and 100 ns/sample) and slow (50 µsec/sample) time domain waveforms are plotted for all the located flashes which occurred on July 24 and August 14, 1985. Each event within the lightning flash has been classified into one of 15 categories (see section IV). Three categories have been added to those discussed by earlier investigators. Anomalous #1 and Anomalous #2 return strokes appear to be hybrid strokes that contain some features of both first and dart-stepped subsequent return strokes. Unknown cloud pulses are events that we could not identify. Their waveforms, to the author's knowledge, have not been discussed in the literature.

Subsequent reports will follow and may include more than the 15 categories defined herein.

So far we have identified 65 first, 23 dart-stepped subsequent, 8 normal subsequent, 8 chaotic subsequent, 4 anomalous \$1, 1 anomalous \$2, and 5 multiple ground return strokes. These return strokes all have leader processes preceding them. For leaders that did not exhibit a peak in the dE/dt record corresponding to the return stroke fast transition, an additional 9 stepped, 1 dart-stepped, 0 dart, and 1 chaotic leaders were identified. In the cloud pulse categories, 17 characteristic, 18 narrow positive bipolar (NPBP), 2 narrow negative bipolar (NNBP), and 2 unknown cloud pulses were identified.

One simple analysis was performed in an attempt to determine whether characteristic pulses precede all first return strokes as suggested by Beasley et al. [1982]. These authors state that preliminary variations start, on the average, 118 msec before the return stroke peak for an average duration of 90 ms. Characteristic pulses that occur in the transition period between when preliminary variations end and stepped leaders begin would therefore start, on the average, approximately 28 ms before the first return stroke peak.

First return strokes can be identified either directly from the fast waveforms or by inspection of the electrostatic field change seen in the slow E record caused by charge transfer by the stepped leader. The occurrence of a peak in the slow HF record before the first return stroke peak might be indicative of characteristic pulses. In 1985, the dead time between high resolution events was too great to trigger on both the characteristic pulse and the first return stroke; therefore, the identification

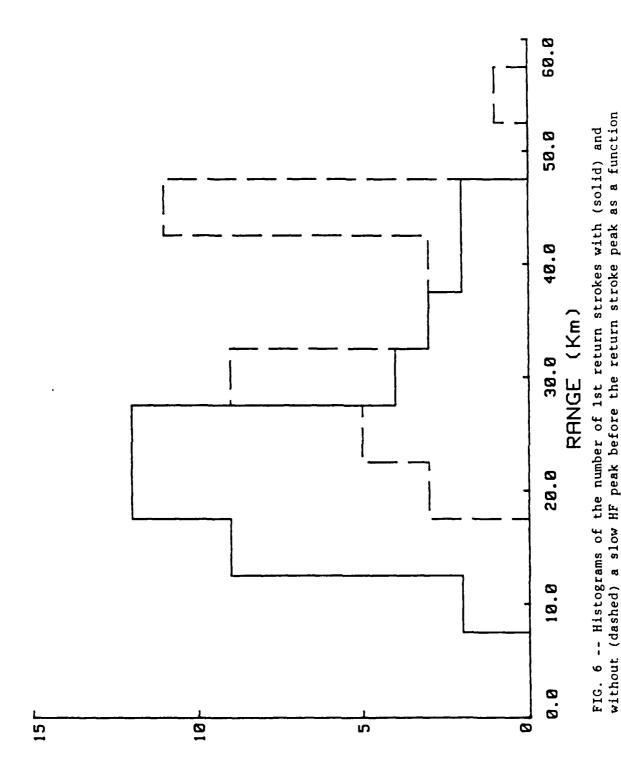
of the preceding characteristic pulse is often uncertain. Given these uncertainties, characteristic pulses were assumed present before the first return stroke as long as there was evidence of both stepped leader activity on the slow E and a slow HF peak near the same time as the start of the leader activity. If a slow HF peak did not occur near the start of the stepped leader, then this case was classified as not having a characteristic pulse. Because of the reduction in the dead time, the 1987 data is better but has not been analyzed yet.

The results are inconclusive. Out of 81 possible reliable cases of first return strokes, 46 of them contained the required HF radiation; however, there appeared to be a strong range bias. Figure 6 shows histograms for the number of cases with and without HF radiation indicative of characteristic pulses prior to the return stroke as a function of range. There is a clear bias toward shorter ranges for the cases with HF radiation. Any attenuation due to propagation effects would be expected to be greater for the return stroke (lower source altitude) than for the characteristic pulse, in contradiction to the observed result. The explanation for this bias is not readily apparent. As a result, it is unclear whether the lack of HF radiation on the remaining 35 cases is due to low signal level or indicates that characteristic pulses do not always precede first return strokes. The characteristic pulses preceded the first return stroke peak radiation from 8 to 60 ms with an average of 21 ± 12 ms in agreement with the results of Beasley et al. [1982].

One flash is worthy of mentioning because of special circumstances.

The "flash", as defined by the slow digitizers, at time 18:34:51.311 on

August 14, 1985 is actually two separate flashes located at different



of range. The presence of the slow HF peak could be an indication of a characteristic pulse (see text).

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ranges. Each event within the flash has been range normalized to its appropriate range, whereas the slow HF has been range normalized to the later event's range.

VI. ACKNOWLEDGEMENTS

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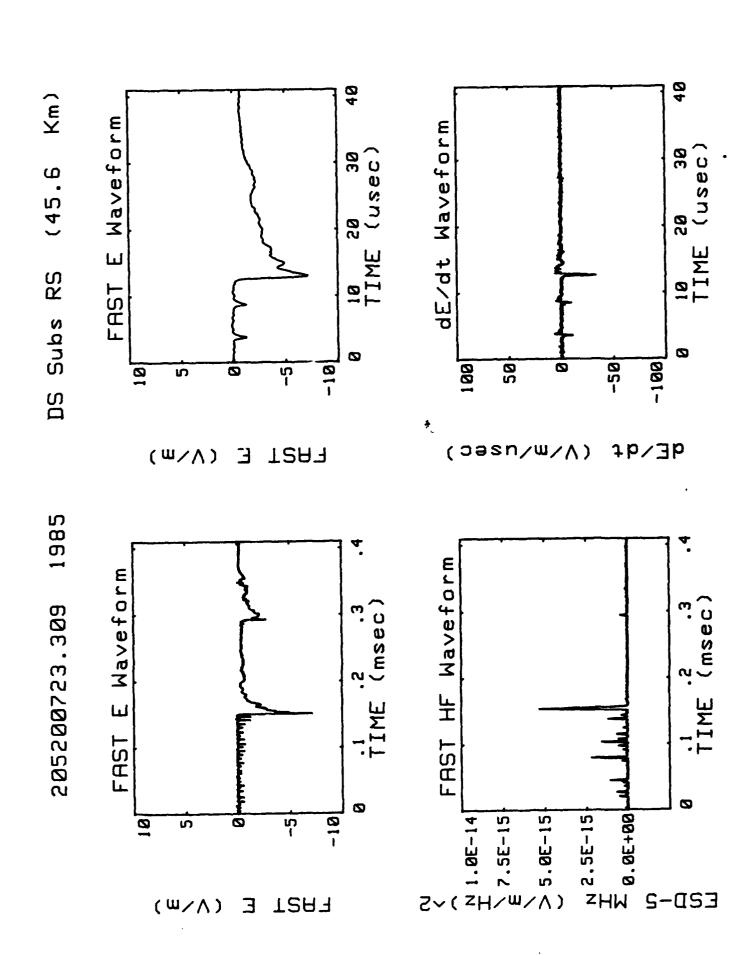
VII. REFERENCES

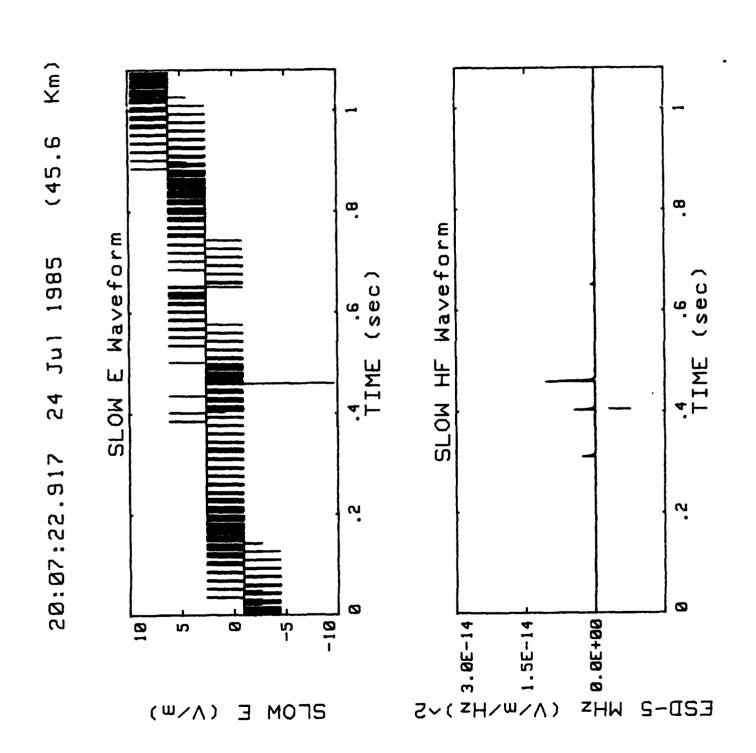
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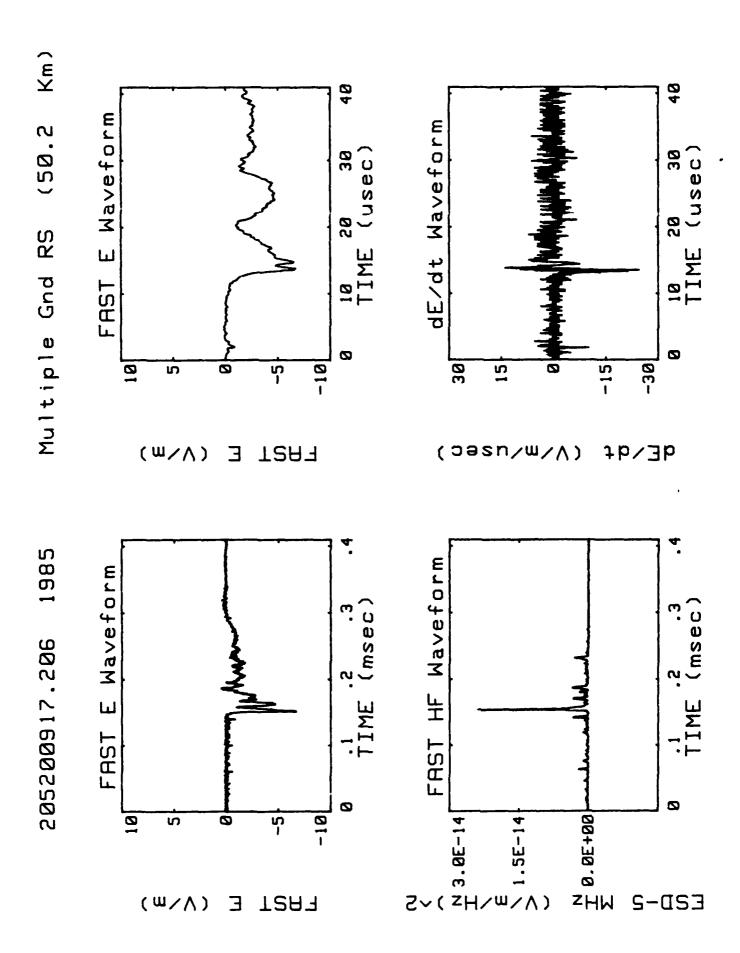
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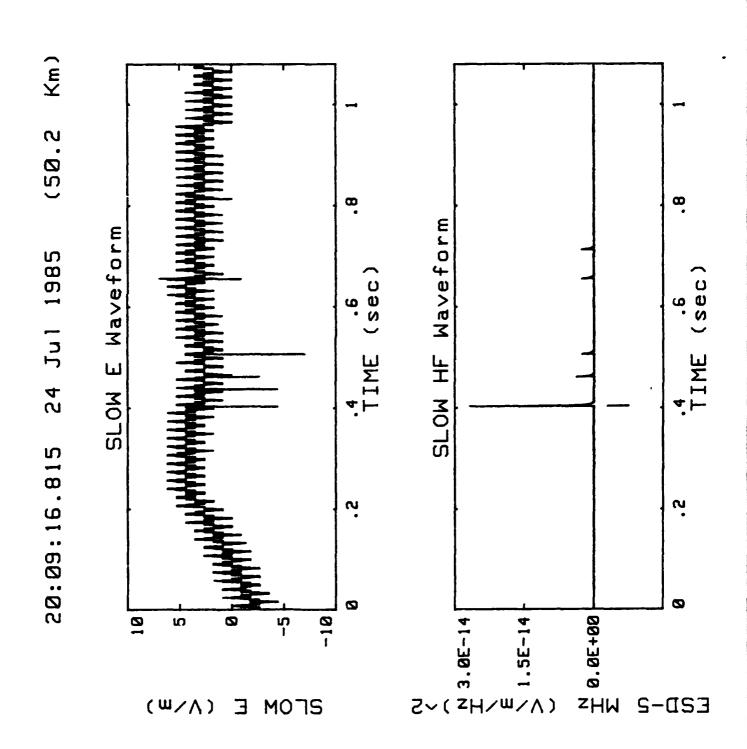
APPENDIX A

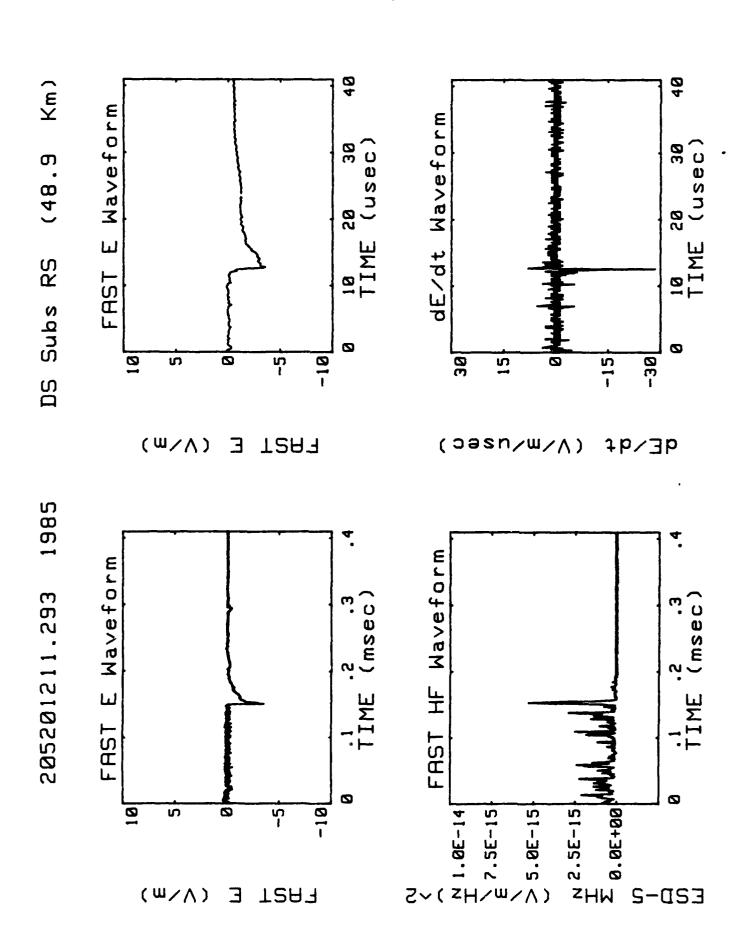
July 24 Data Plots

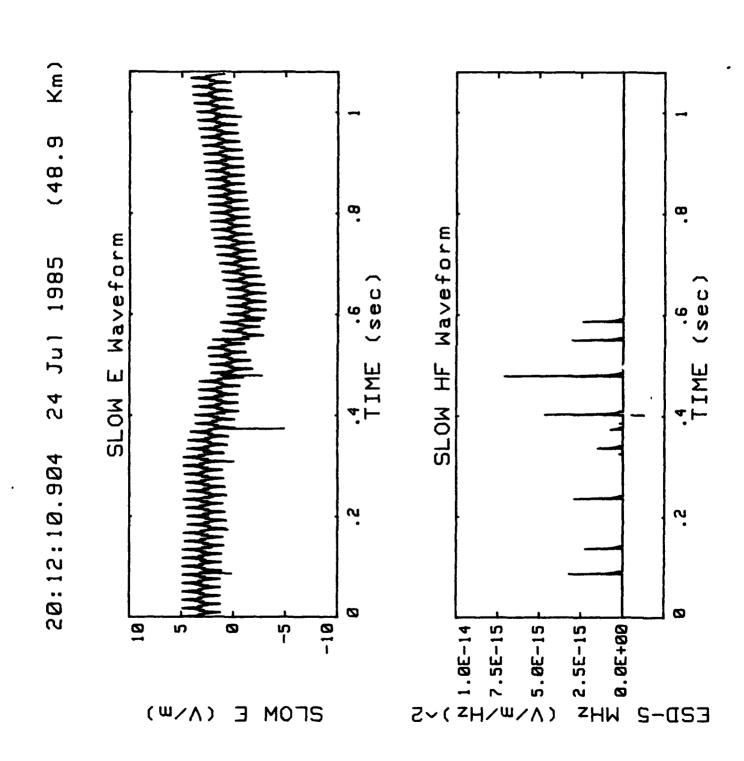


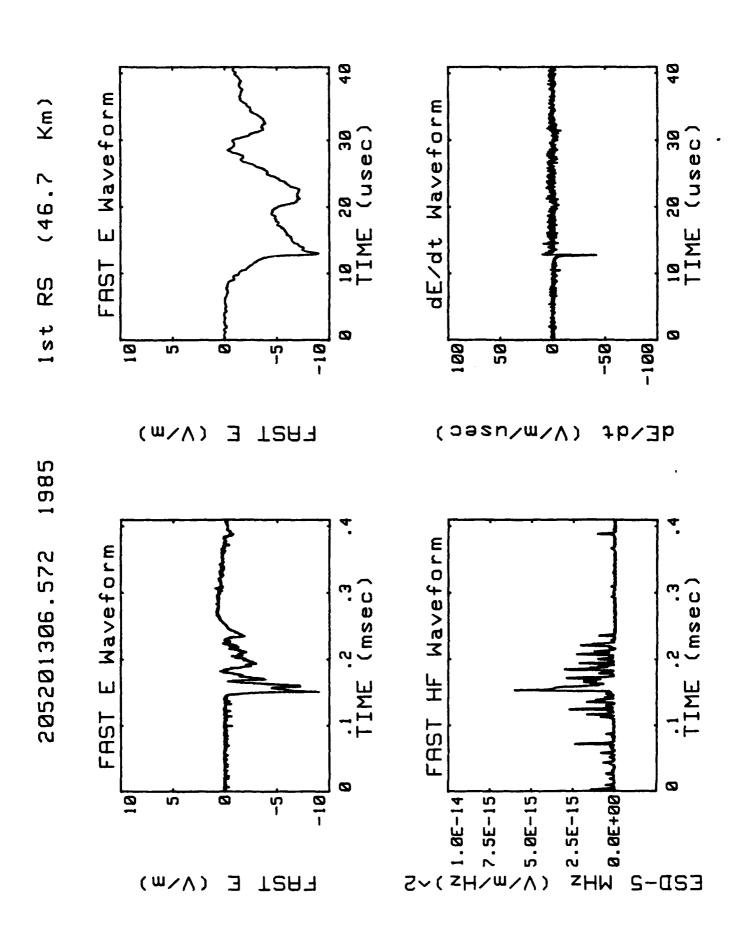


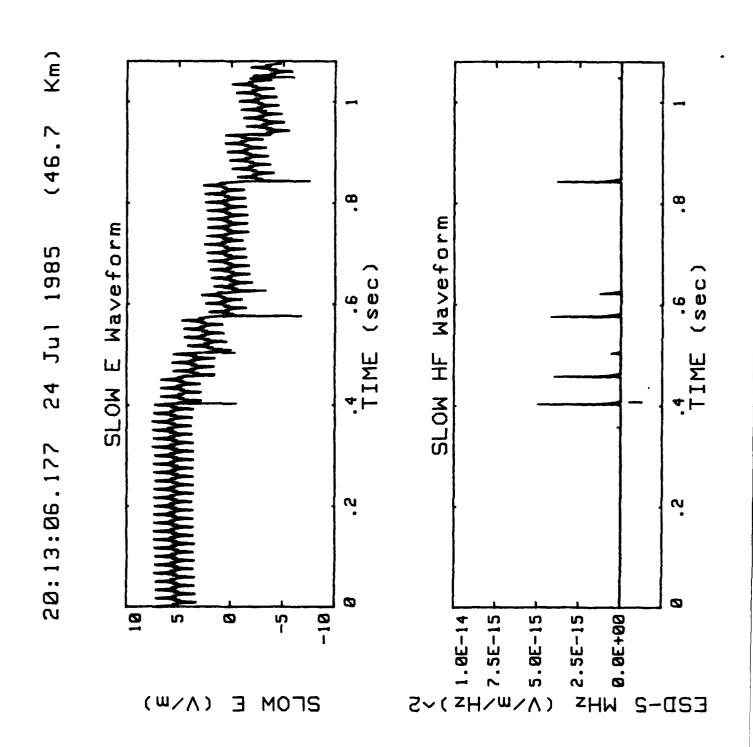


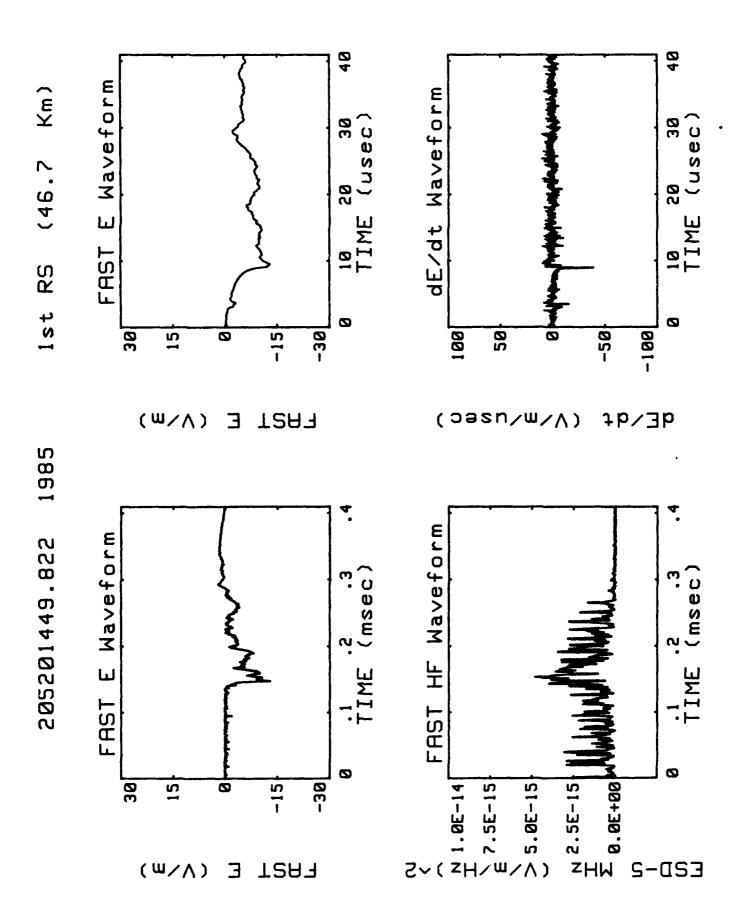


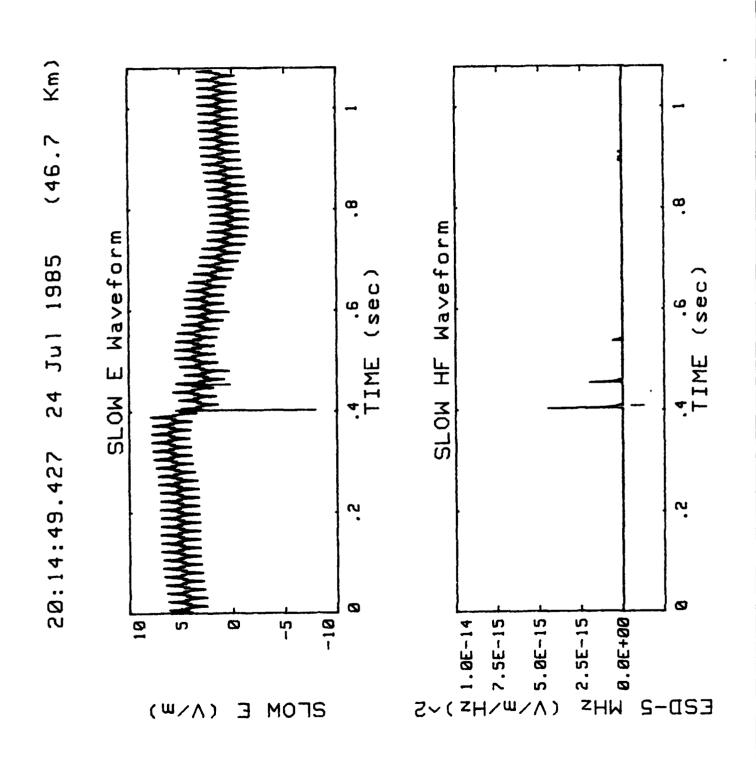


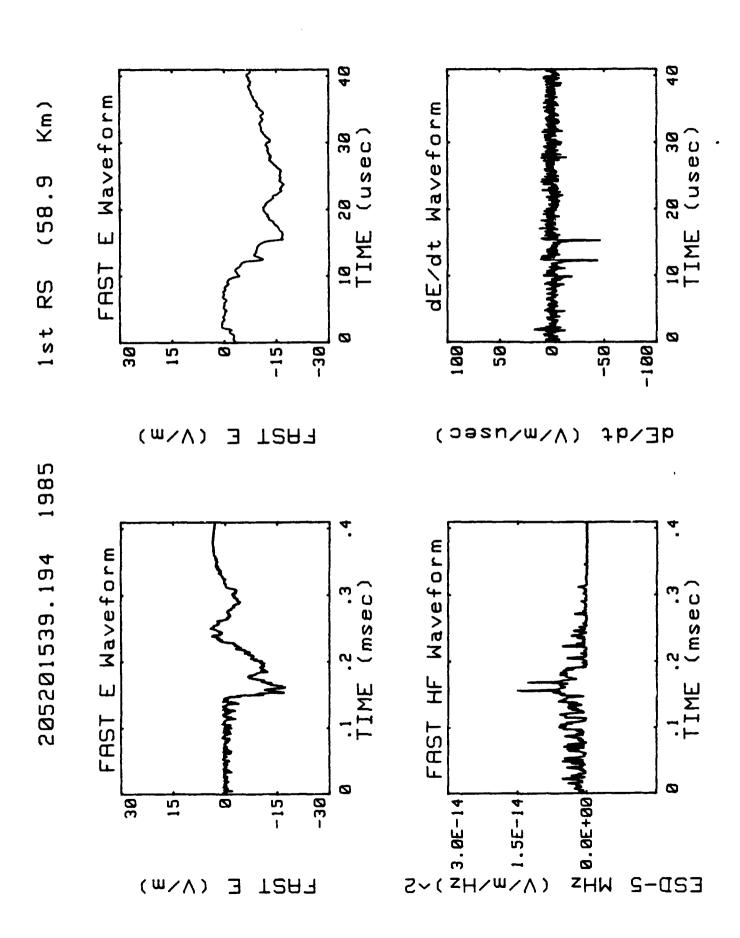


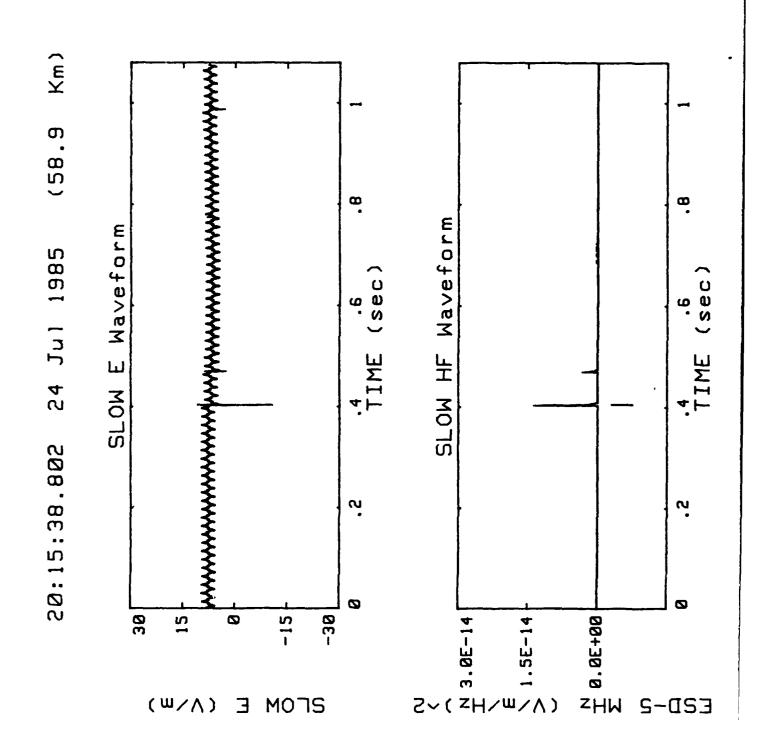


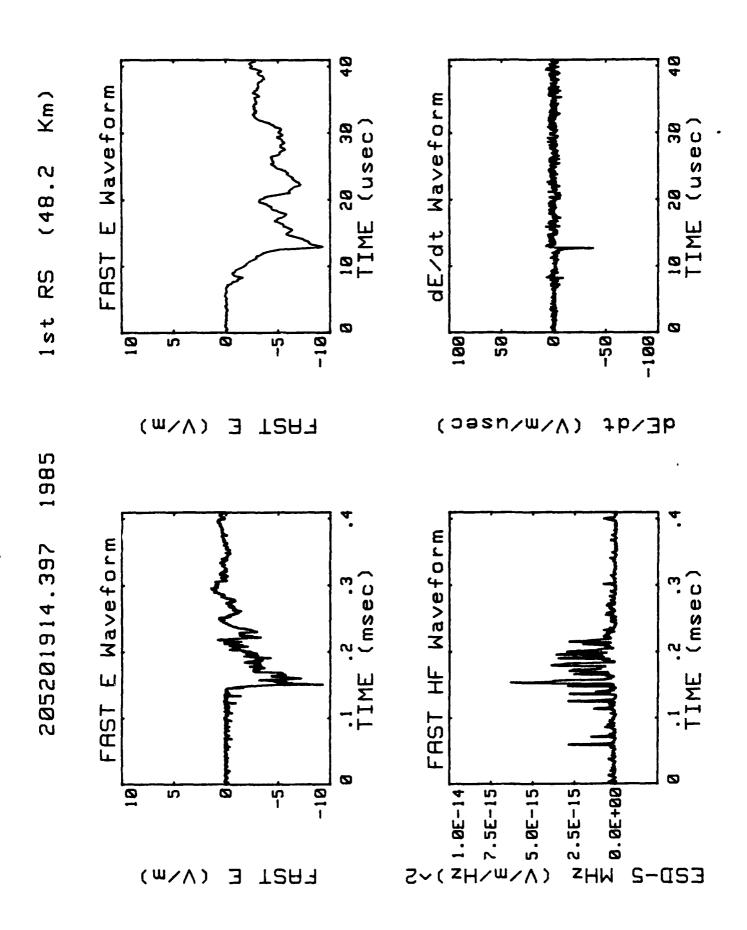


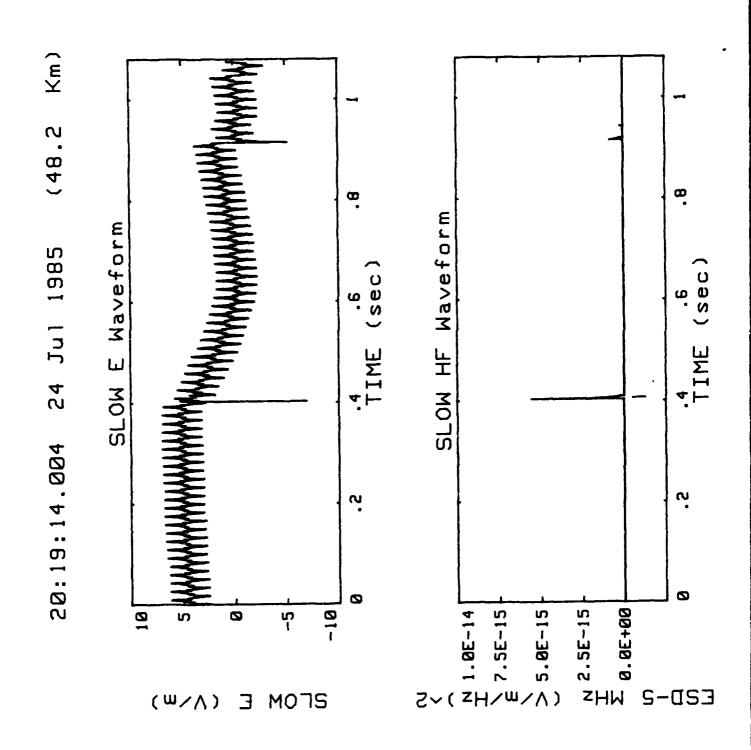


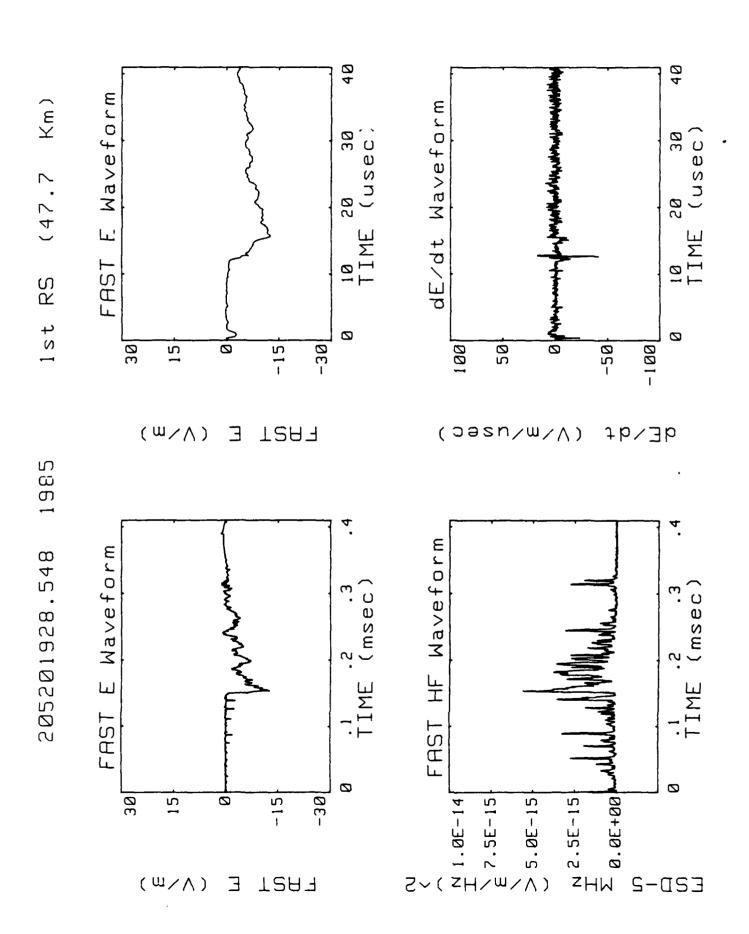


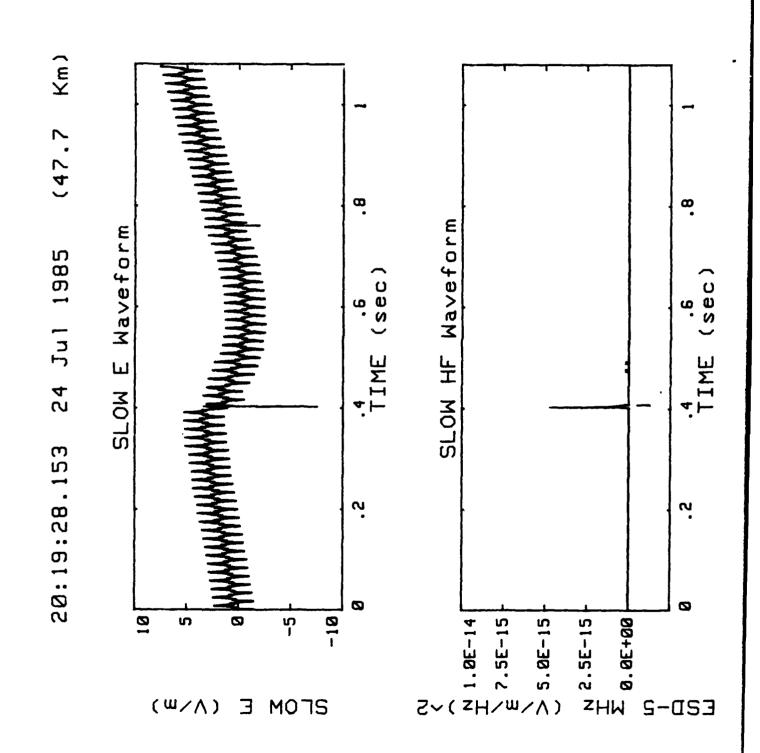


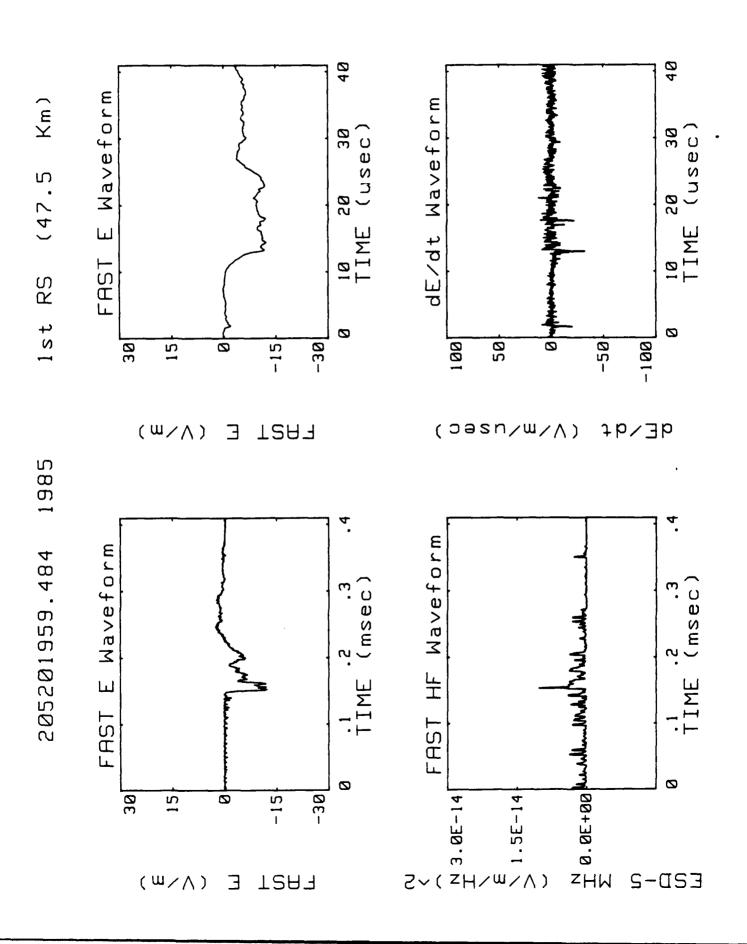


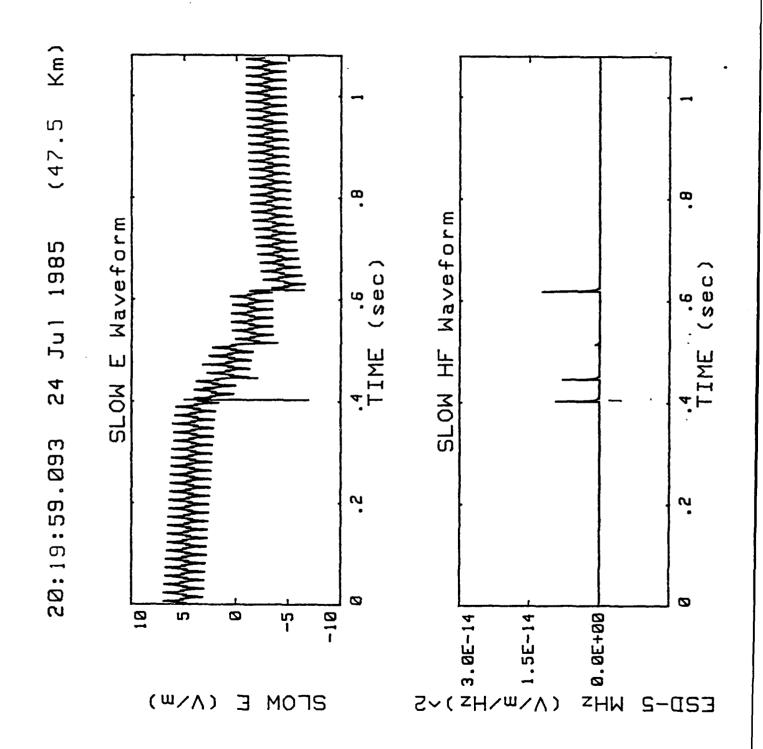


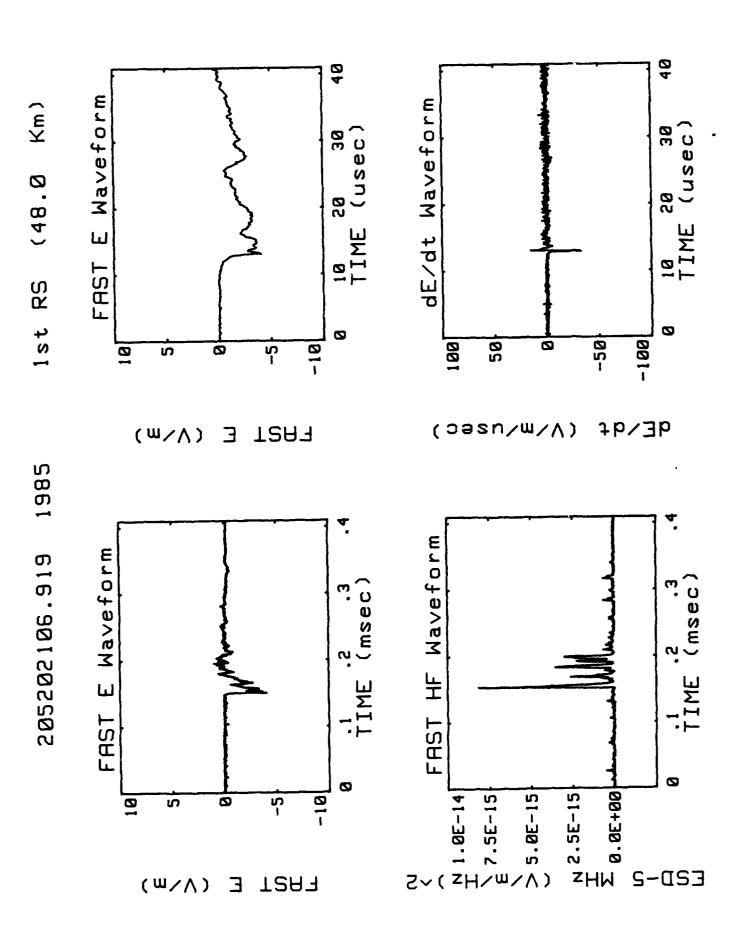


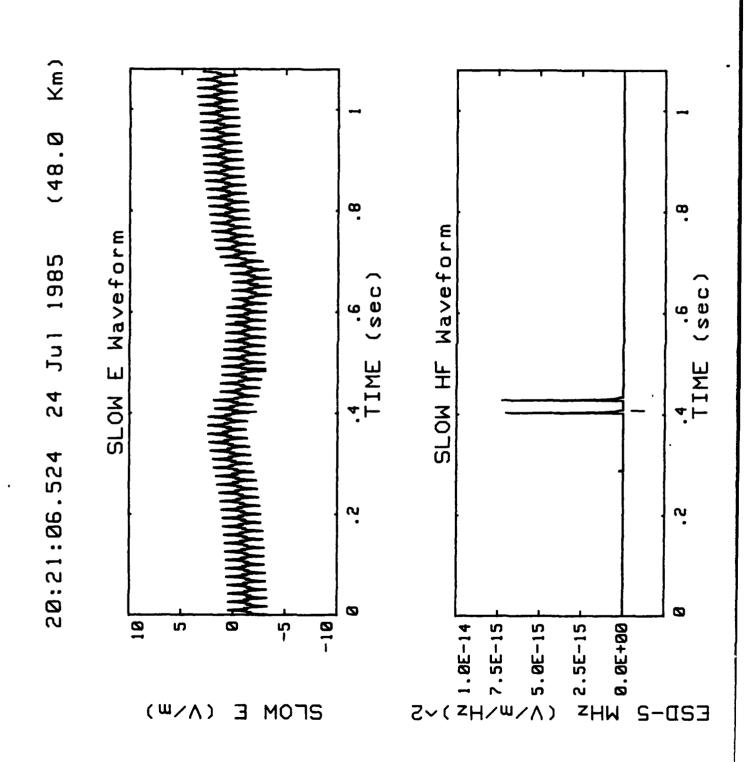


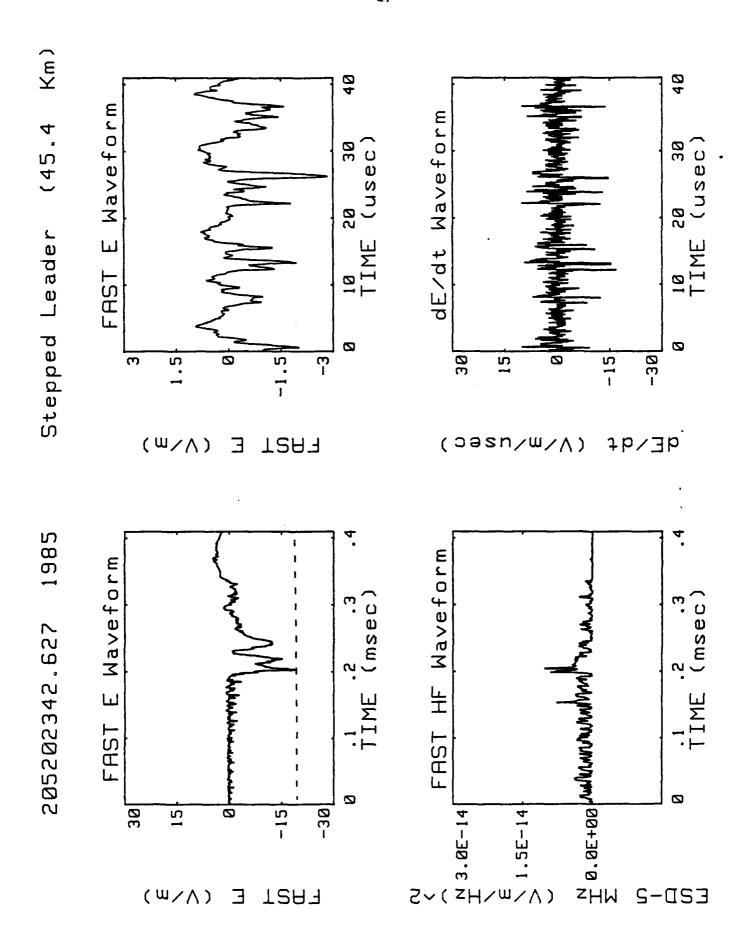


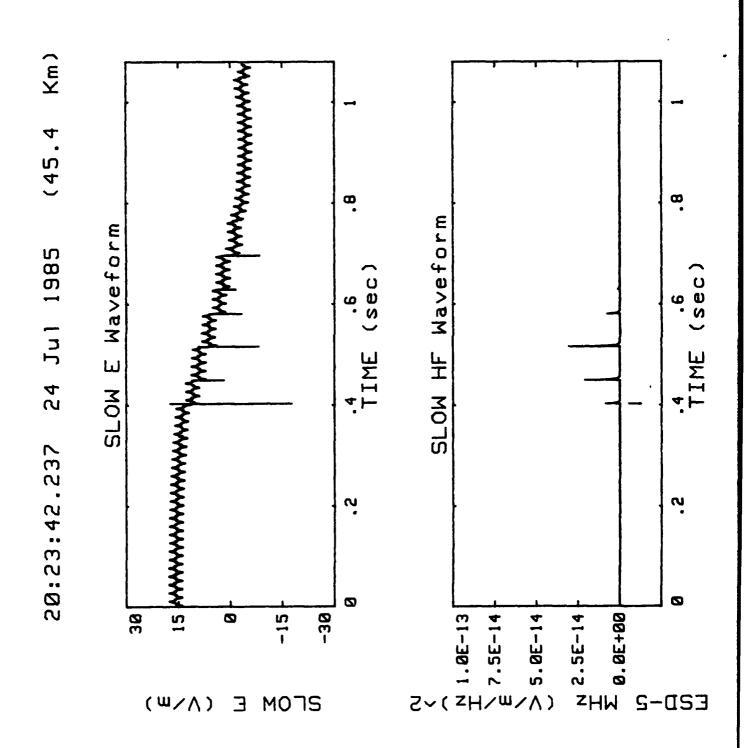


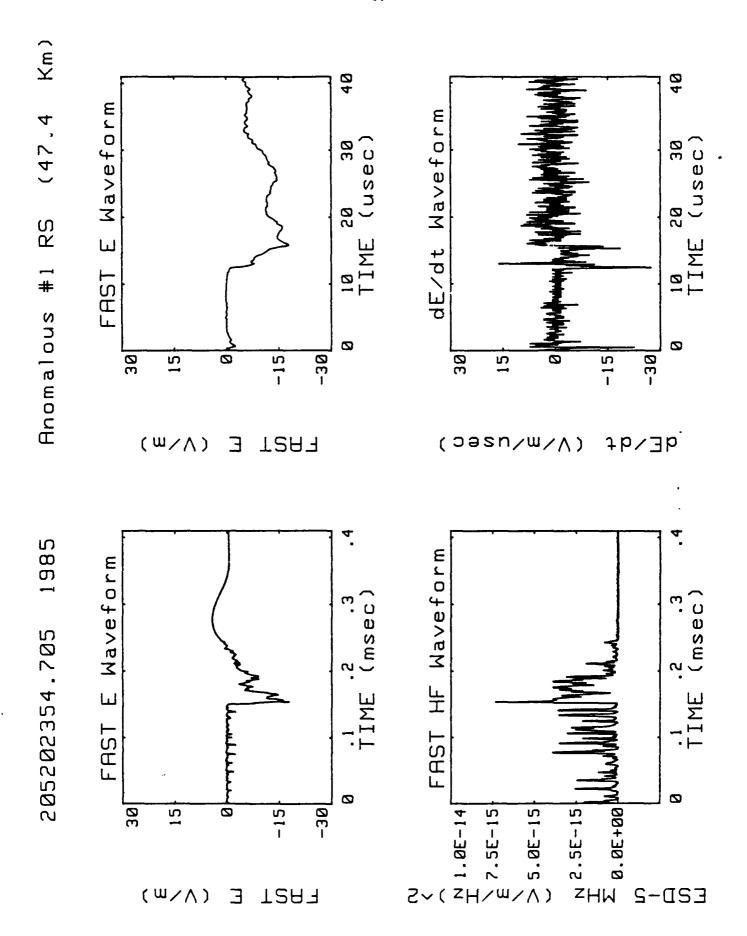


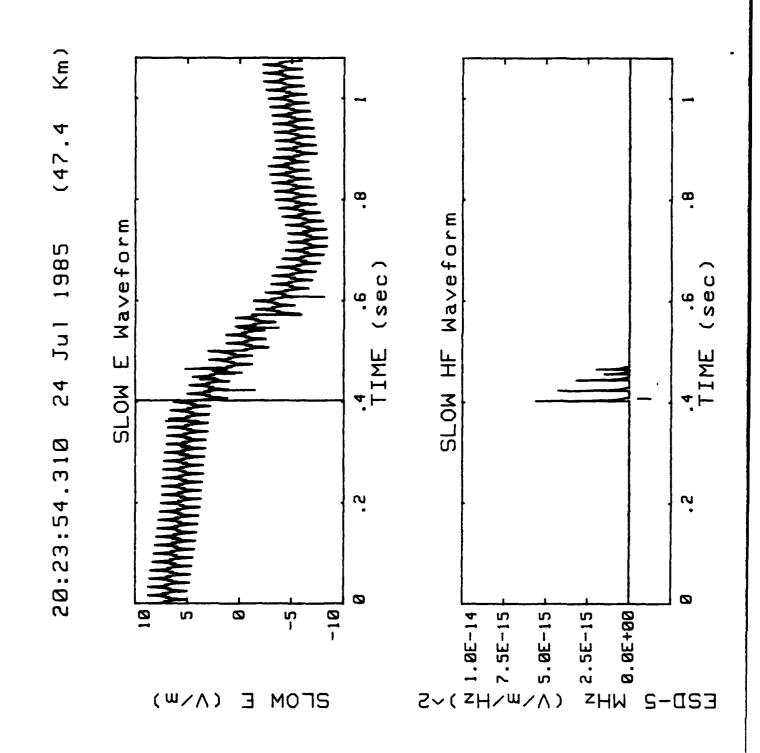


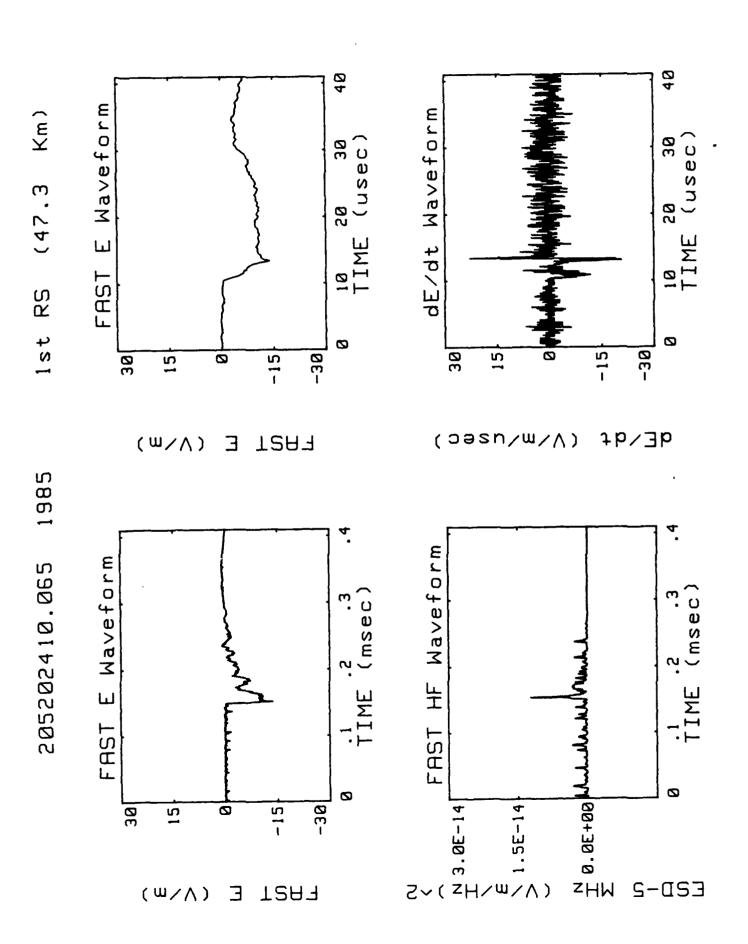


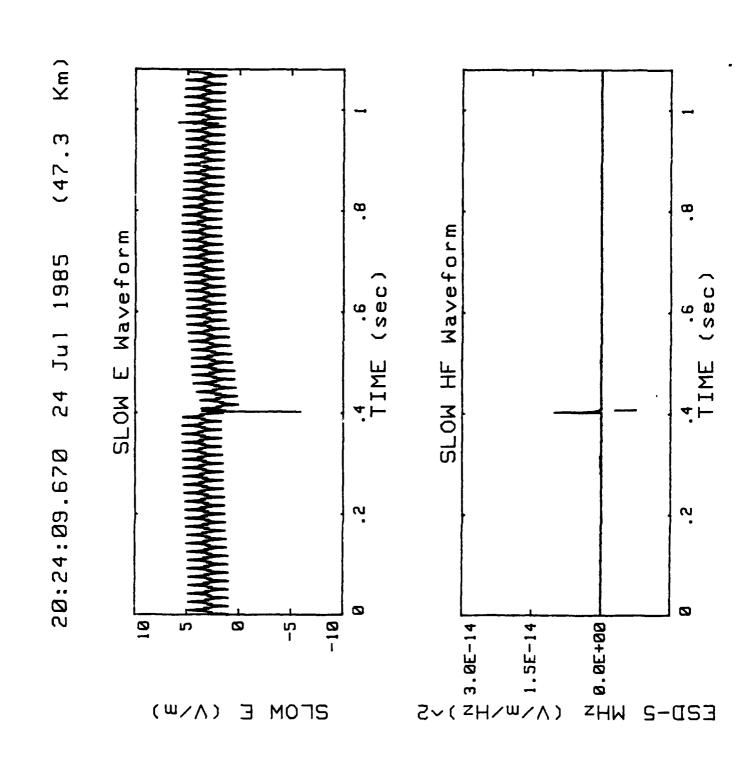


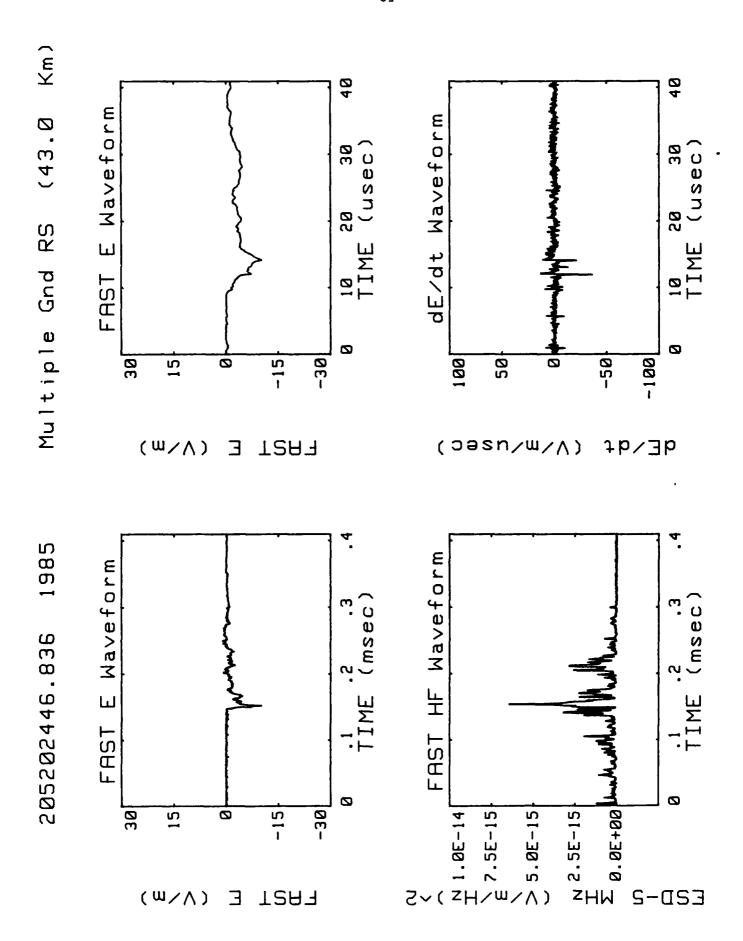


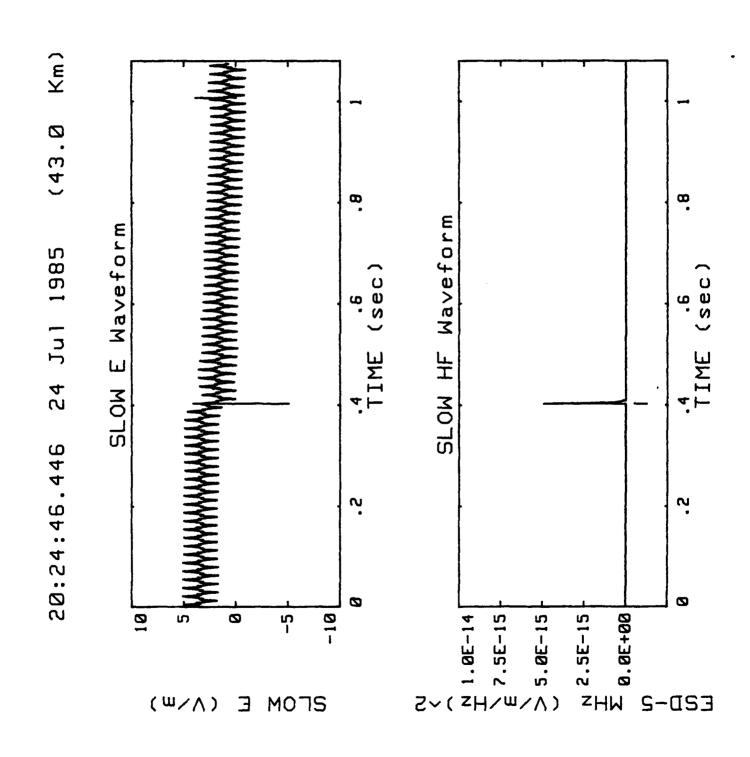


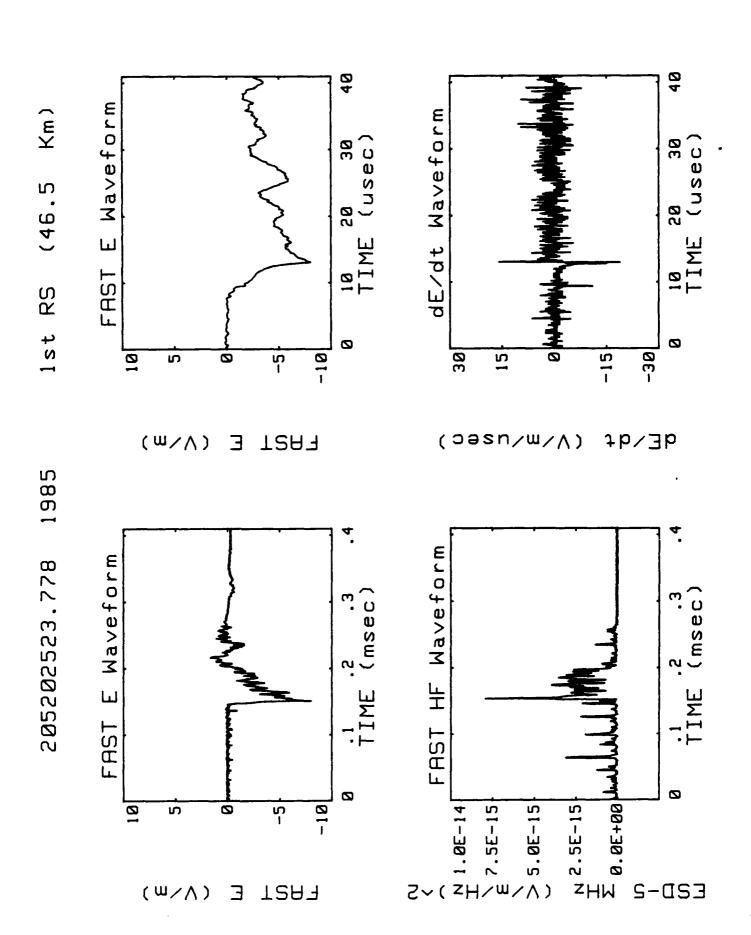


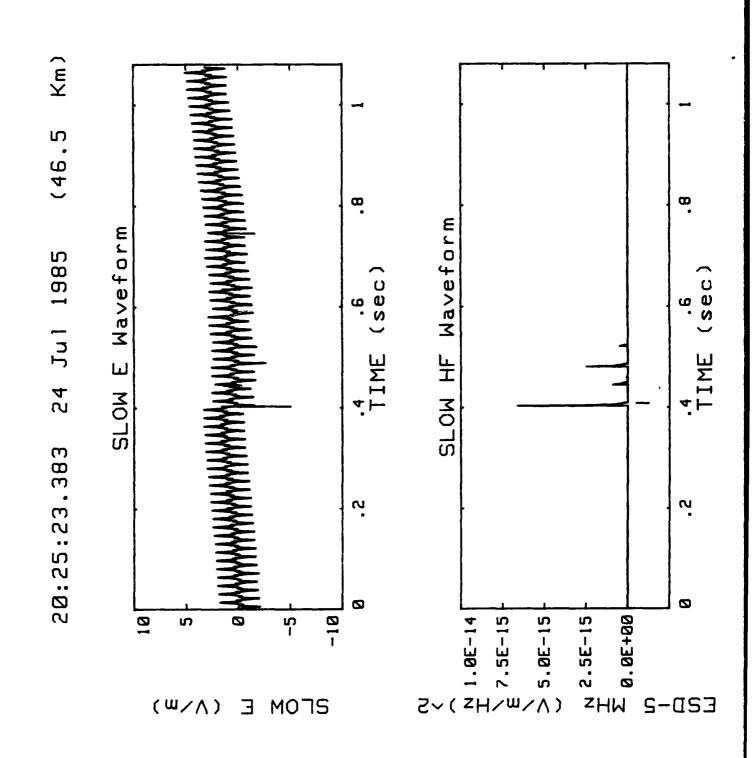


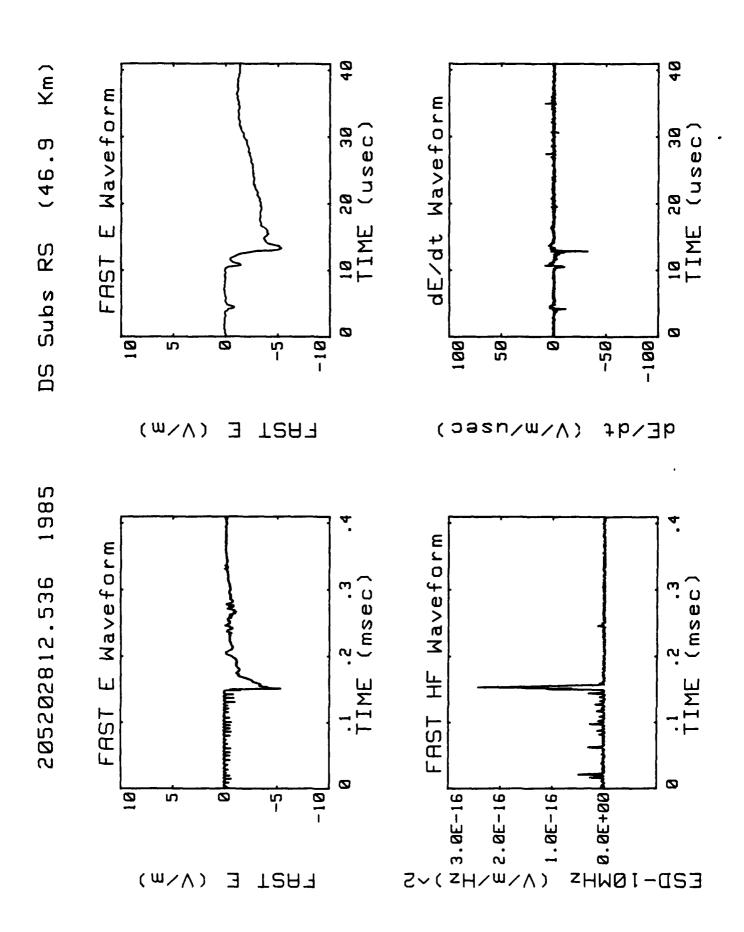


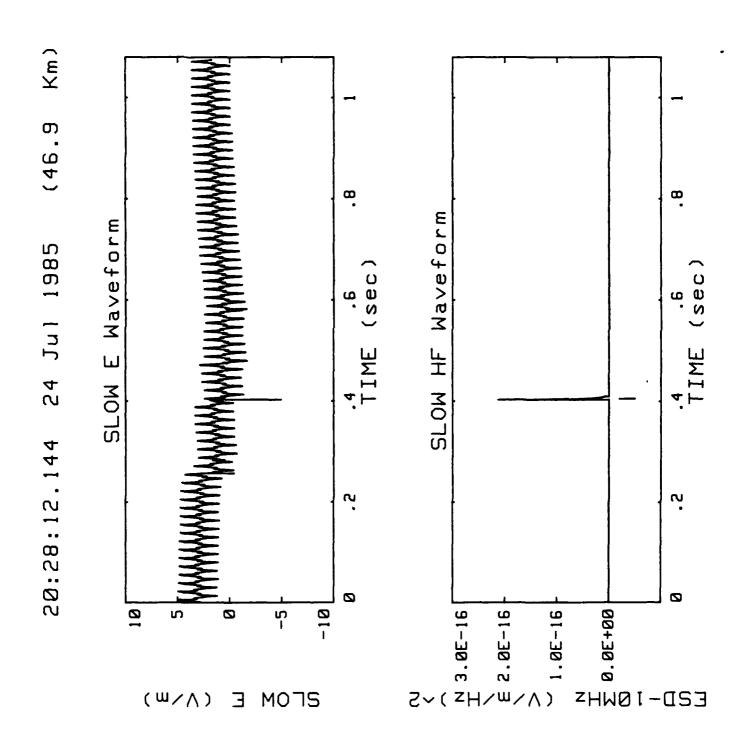


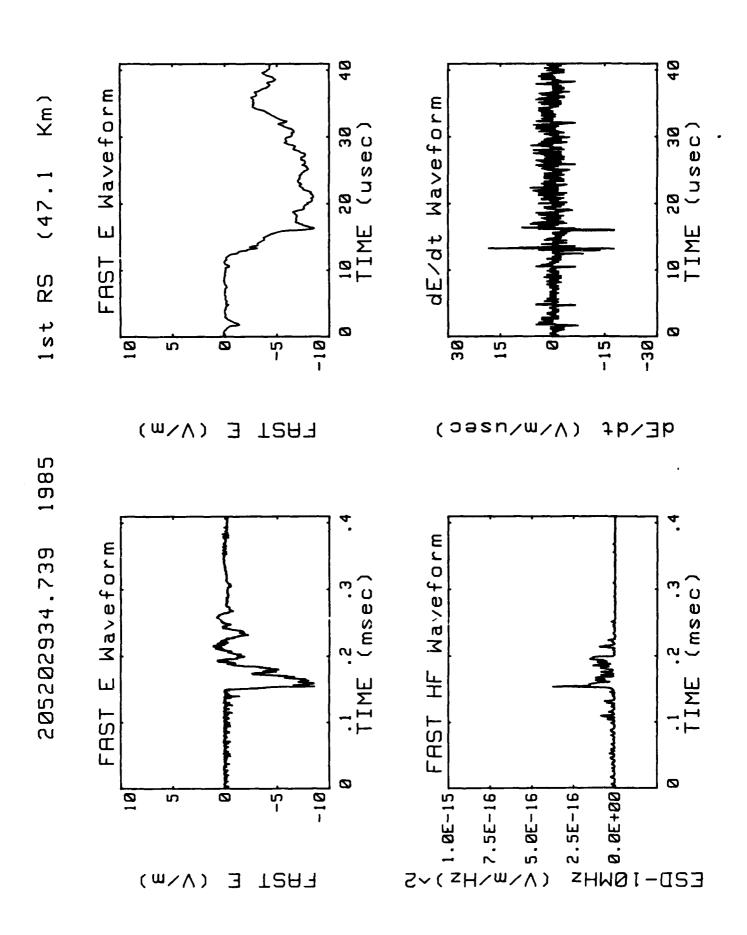


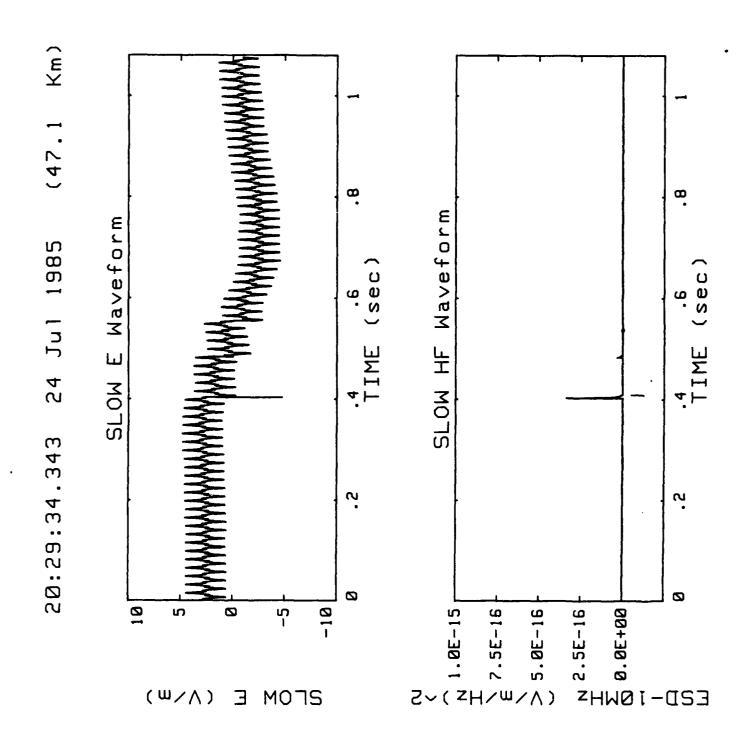


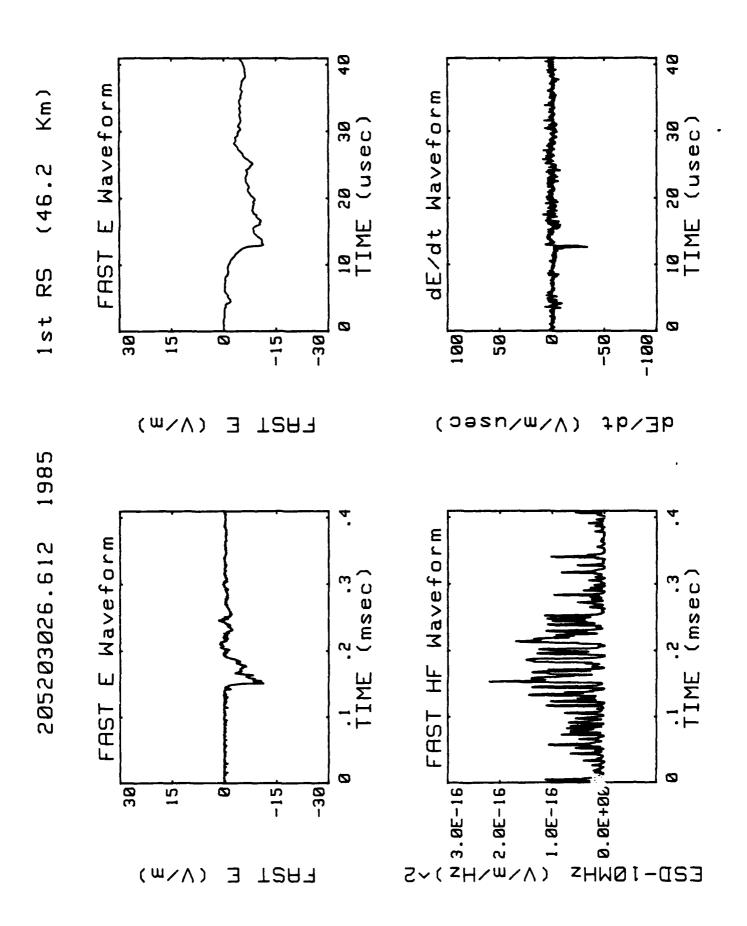


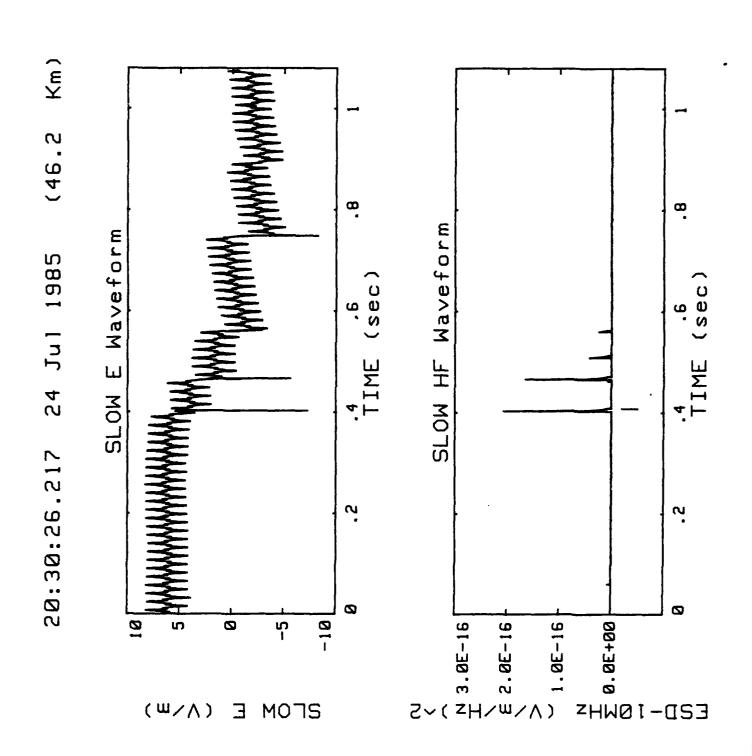


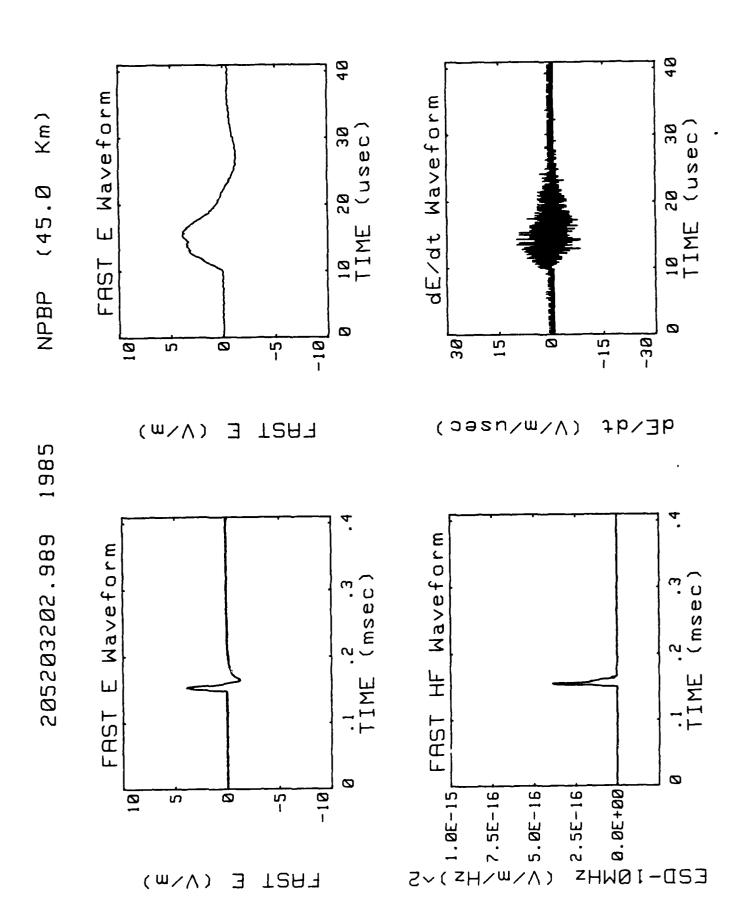


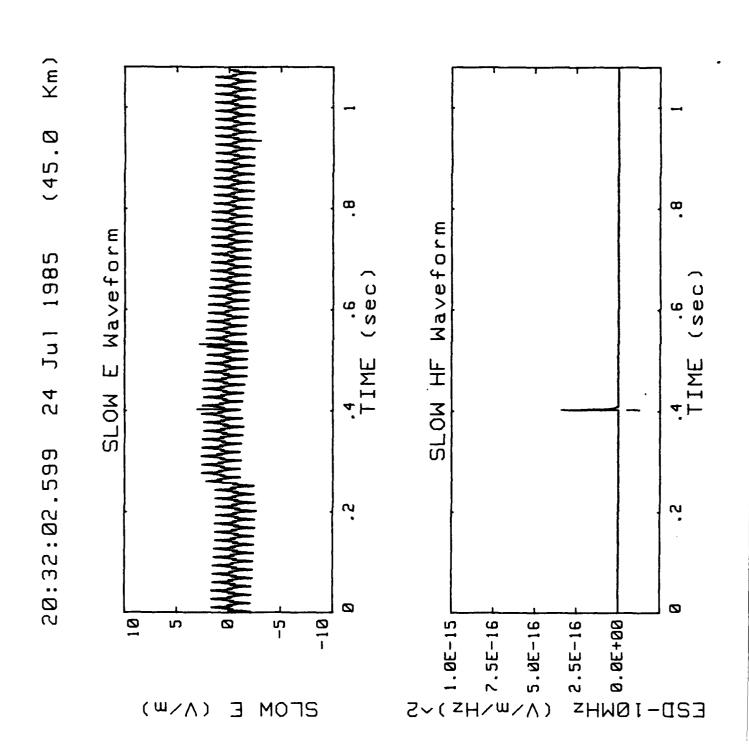


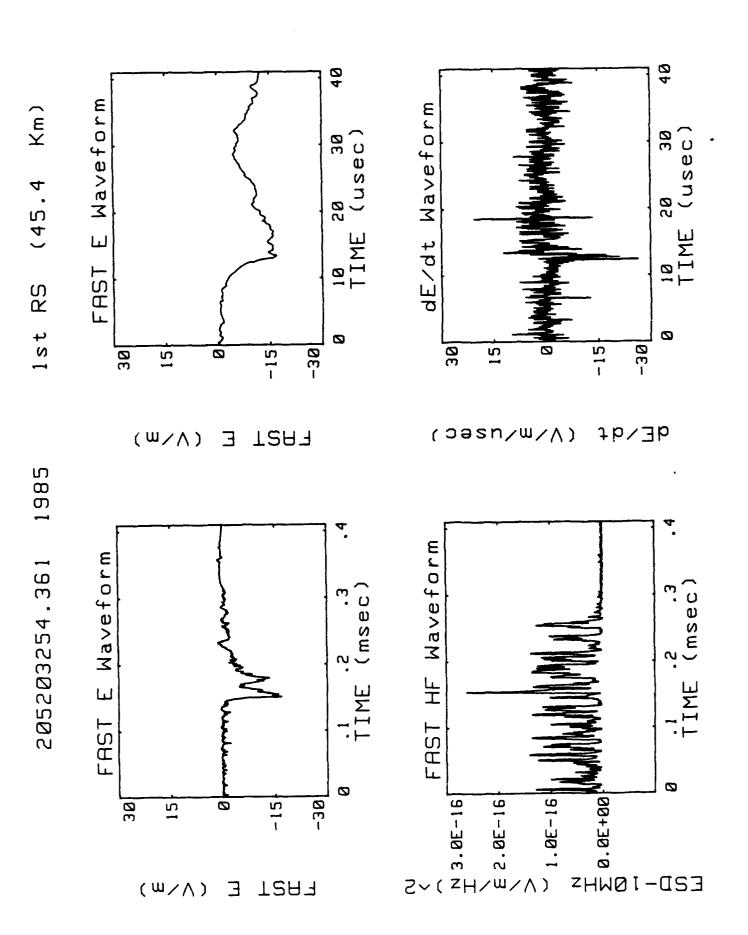


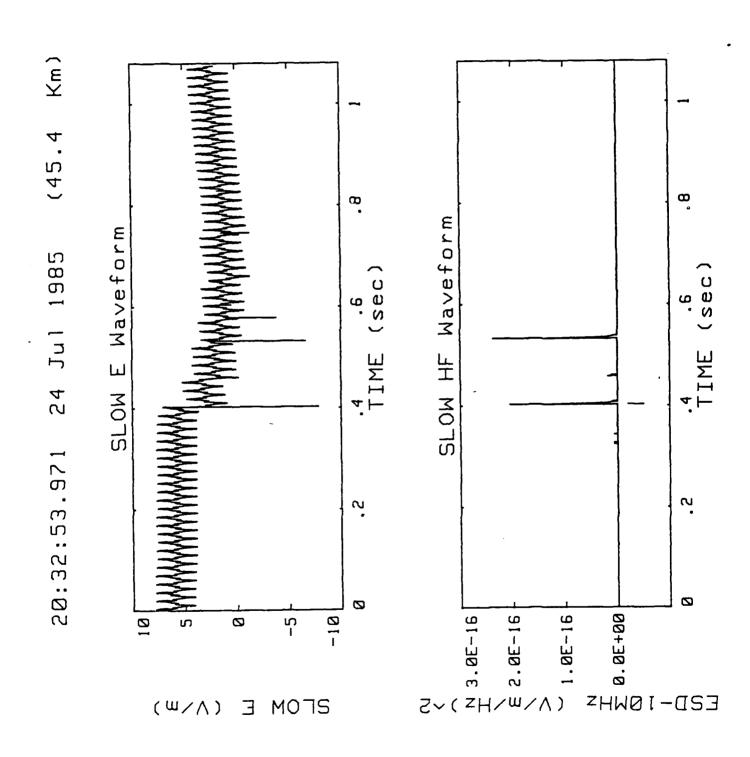


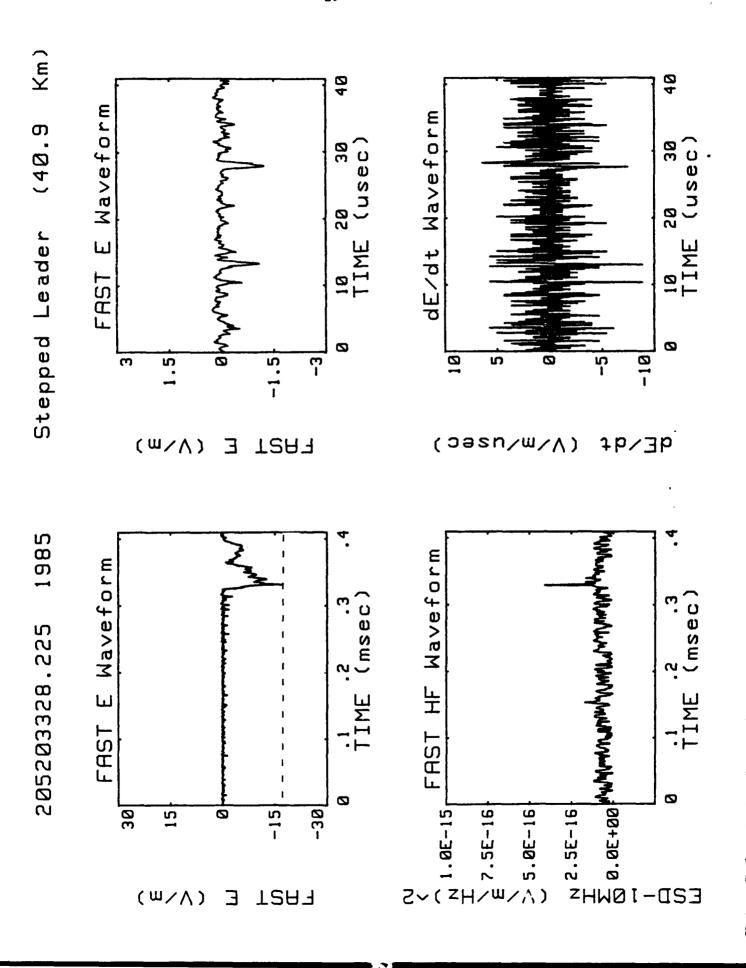


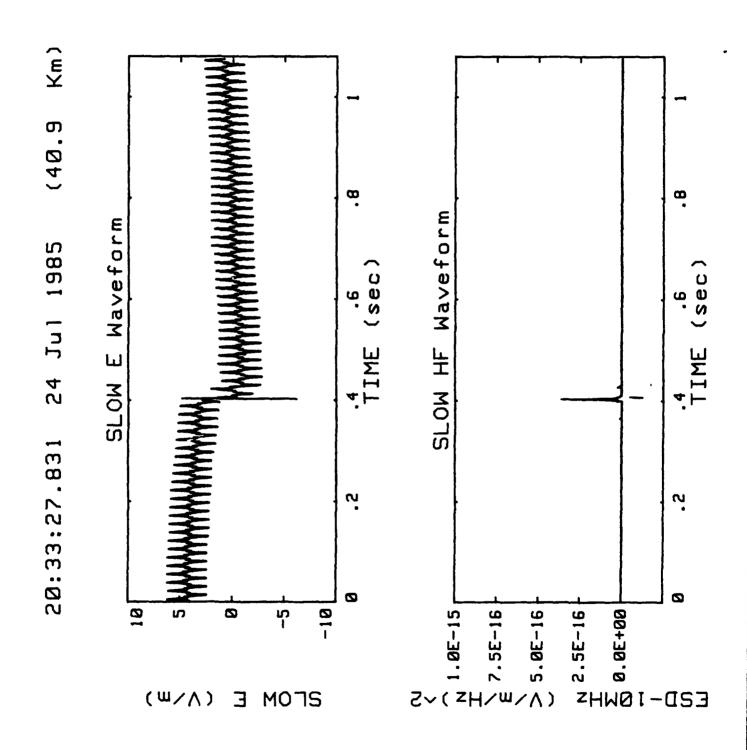


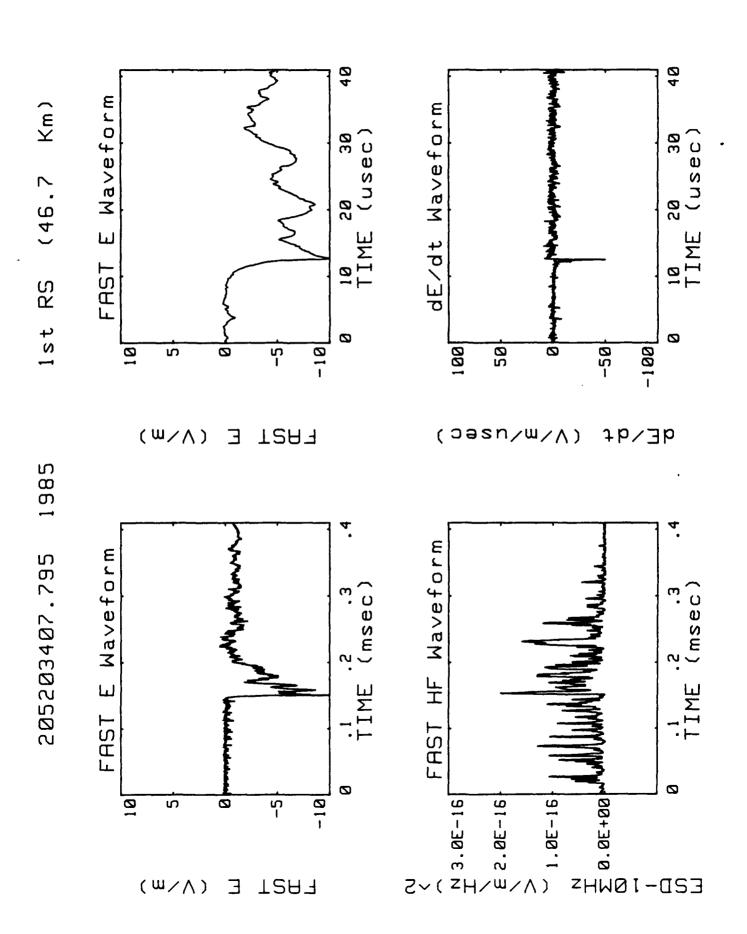


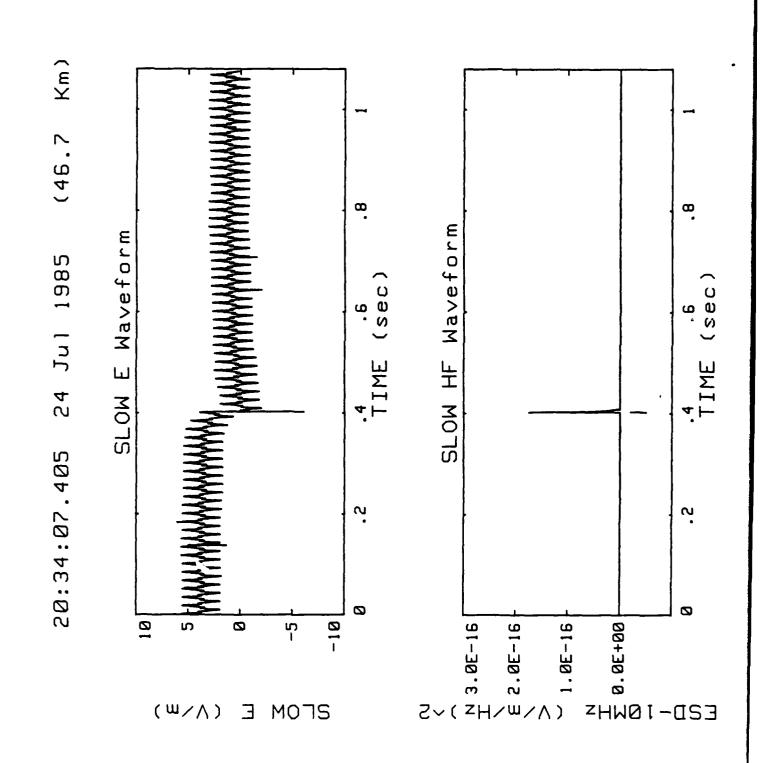


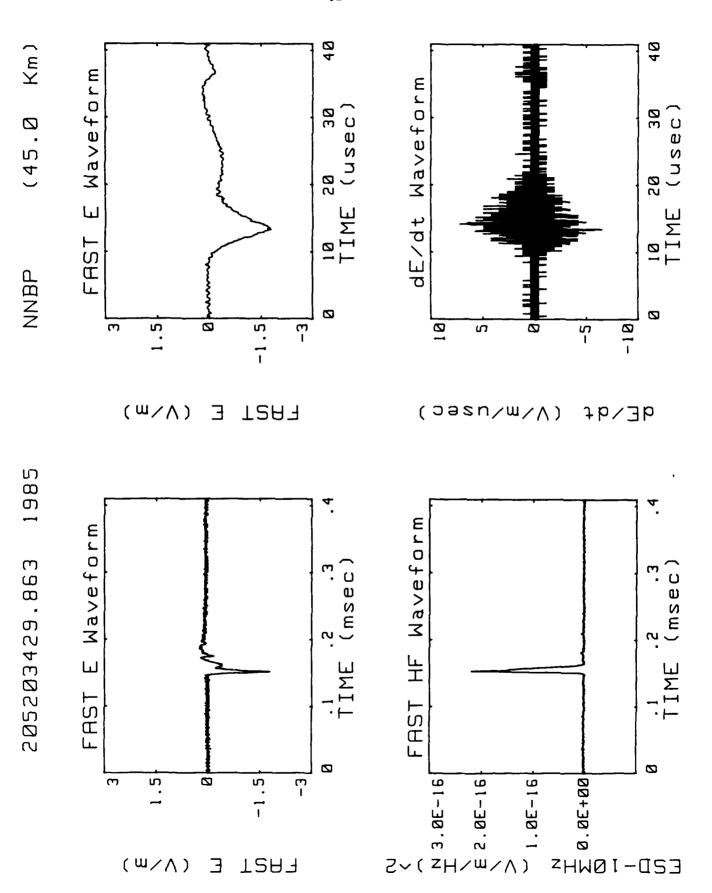


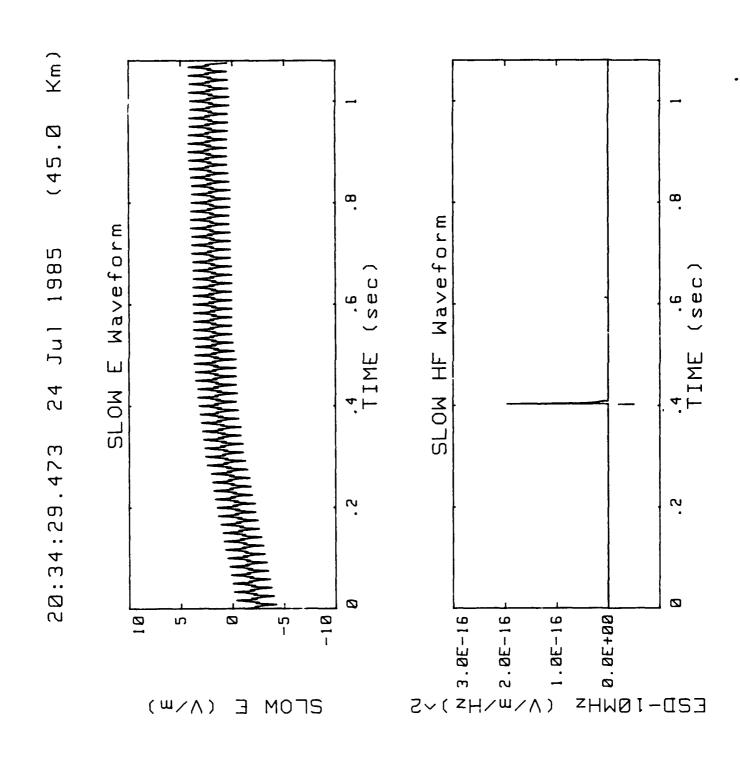


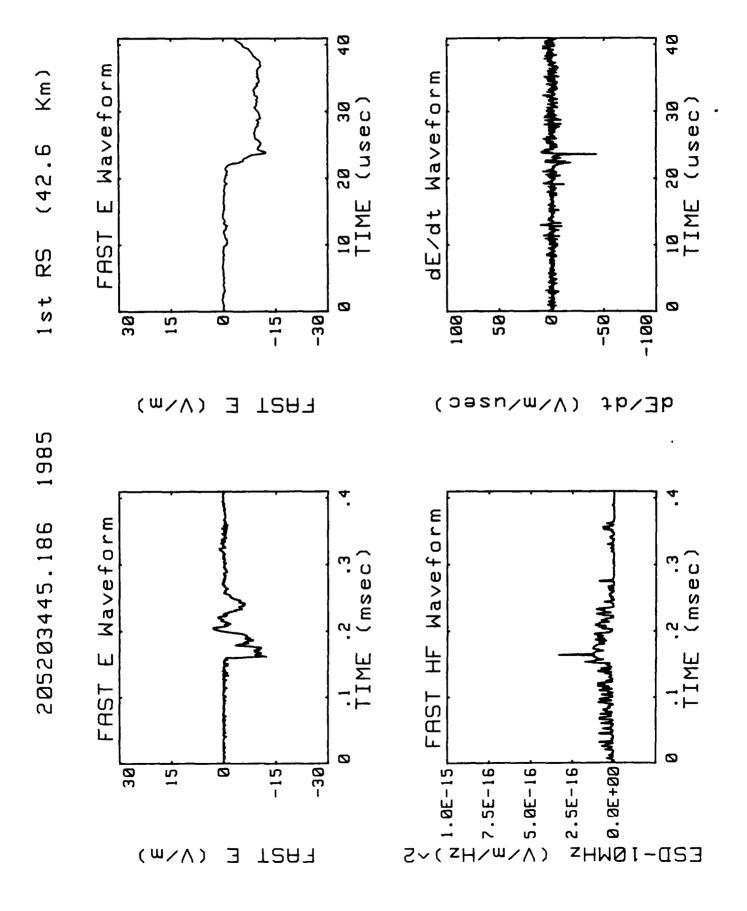


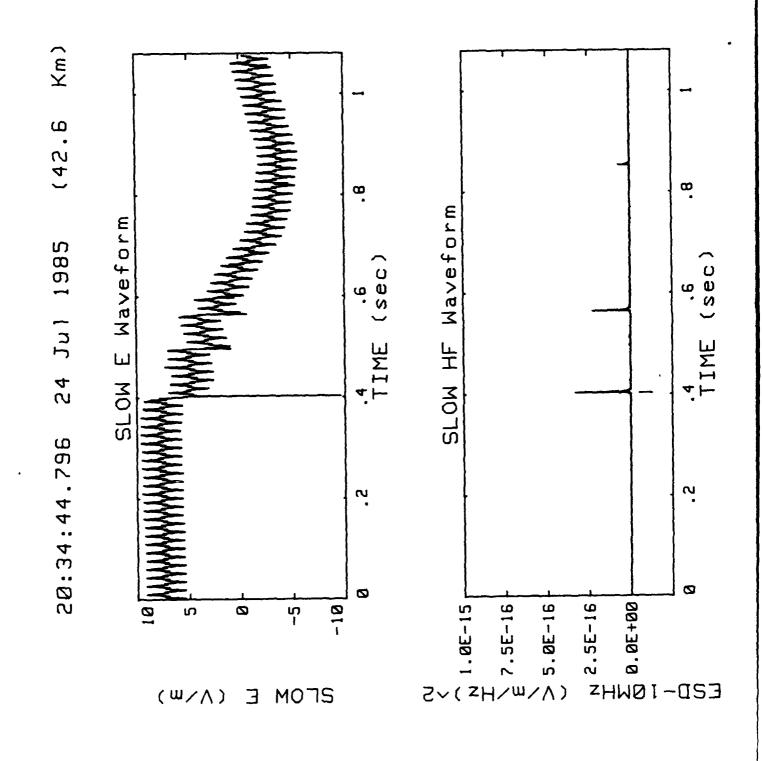


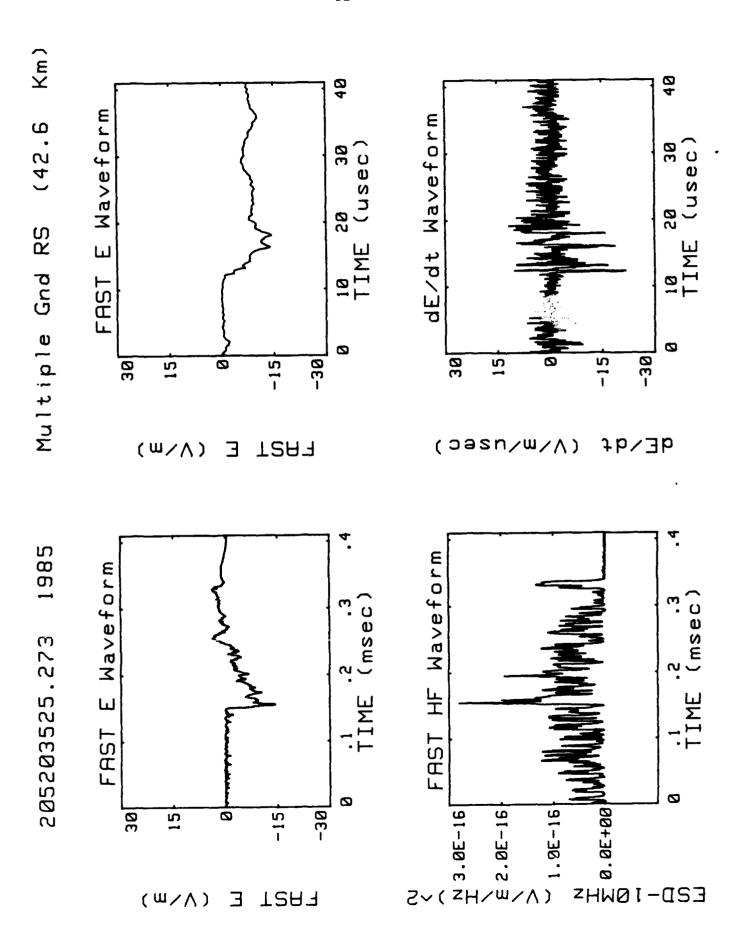


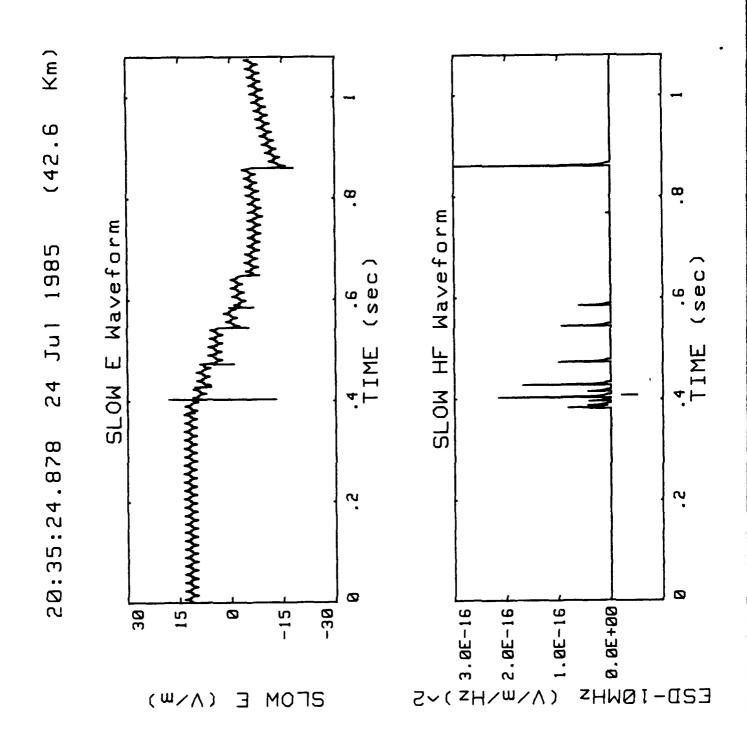


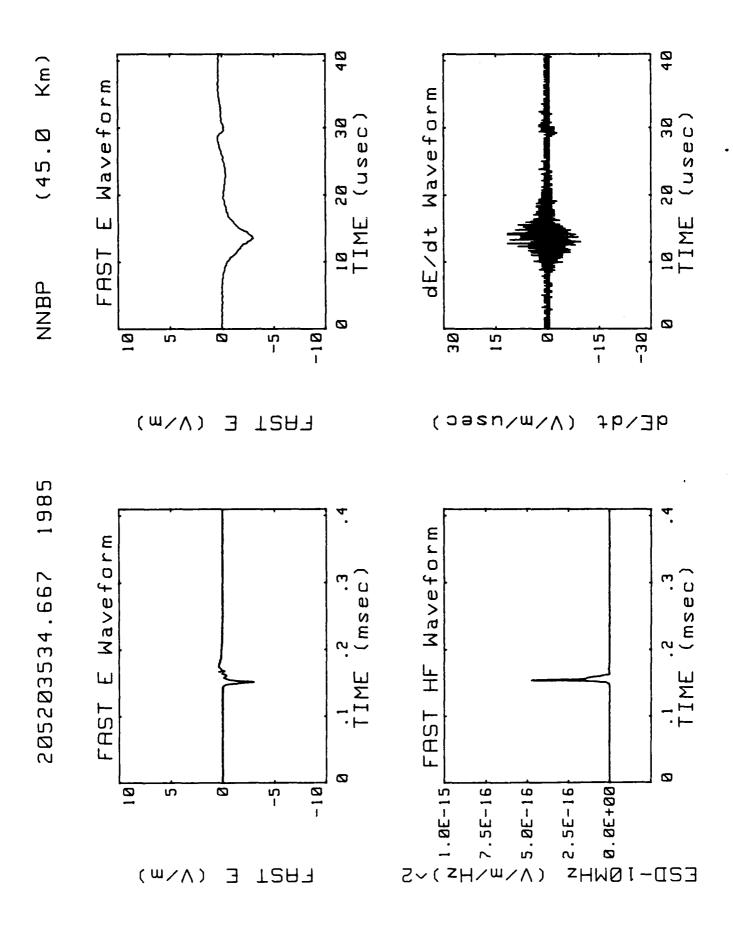


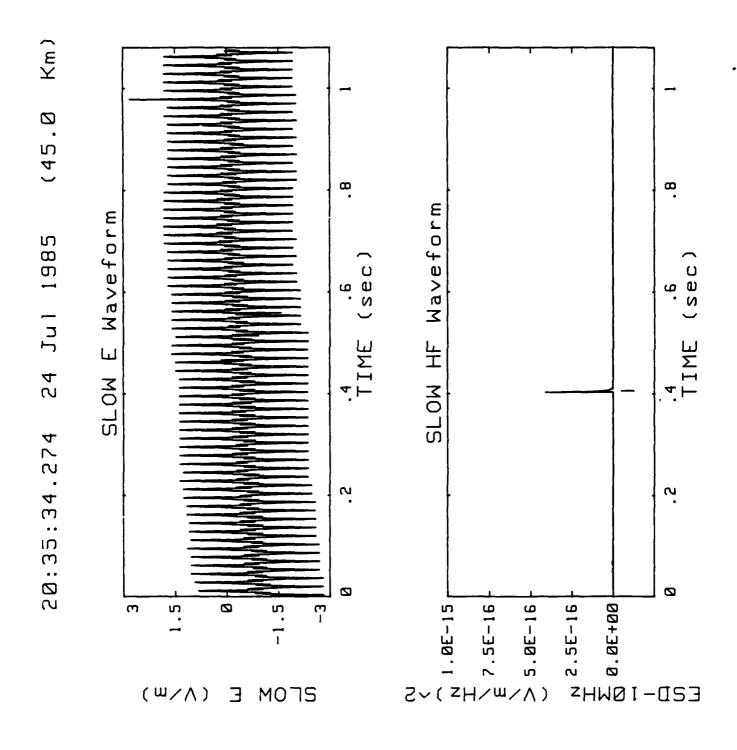


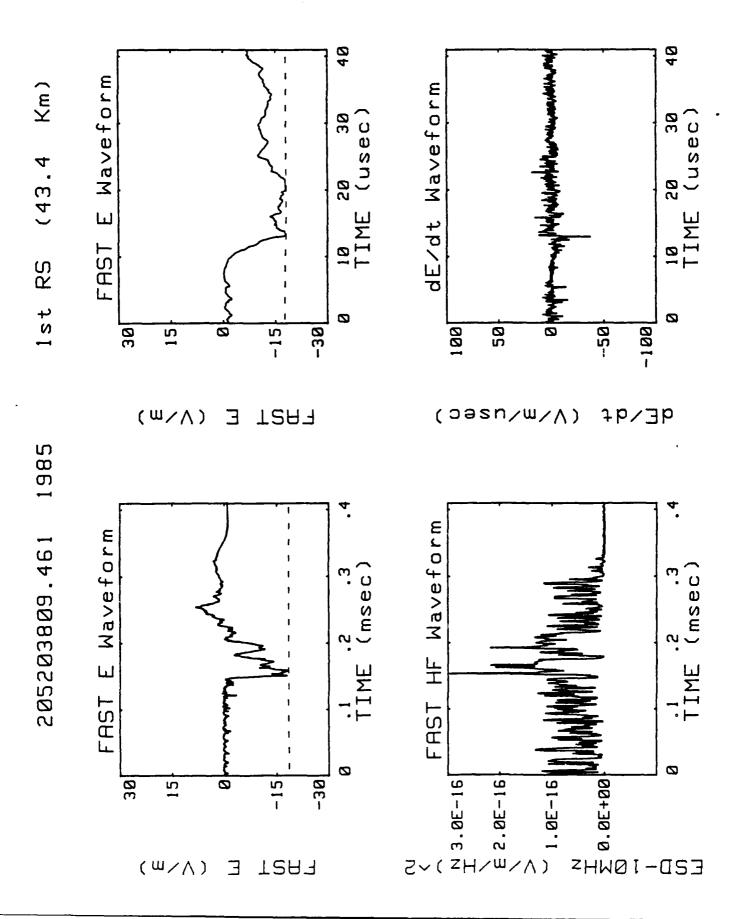


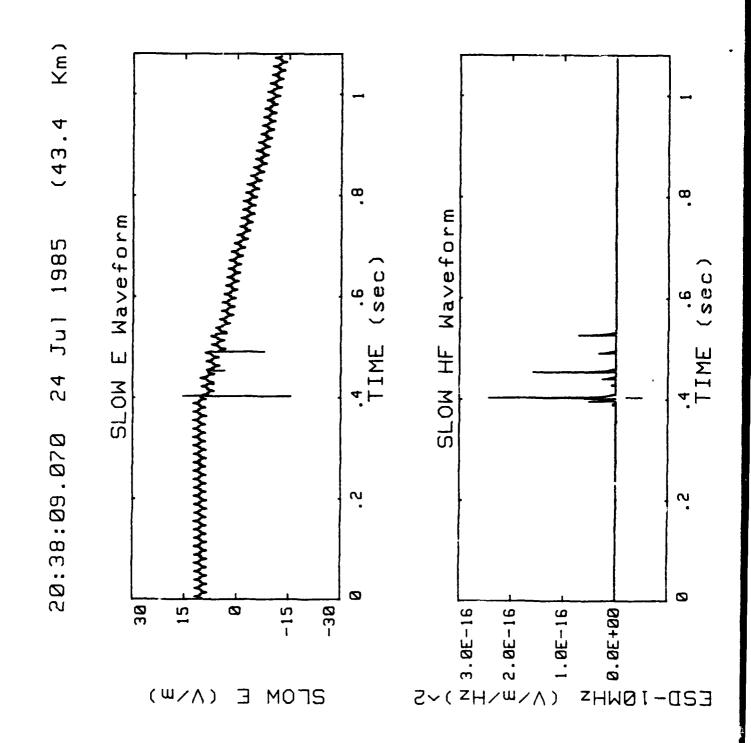


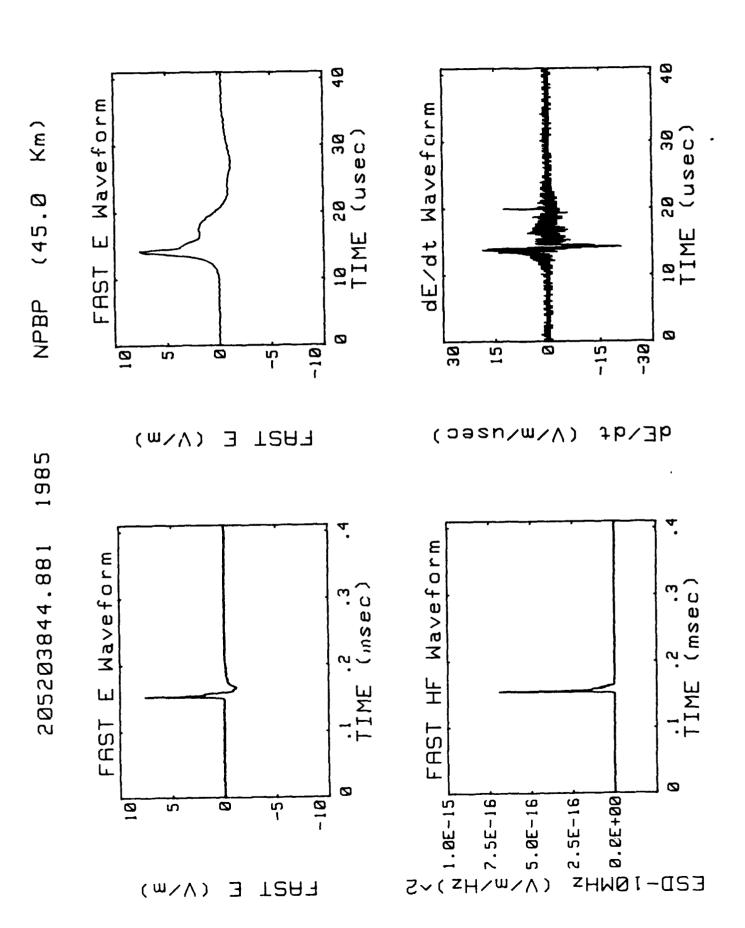


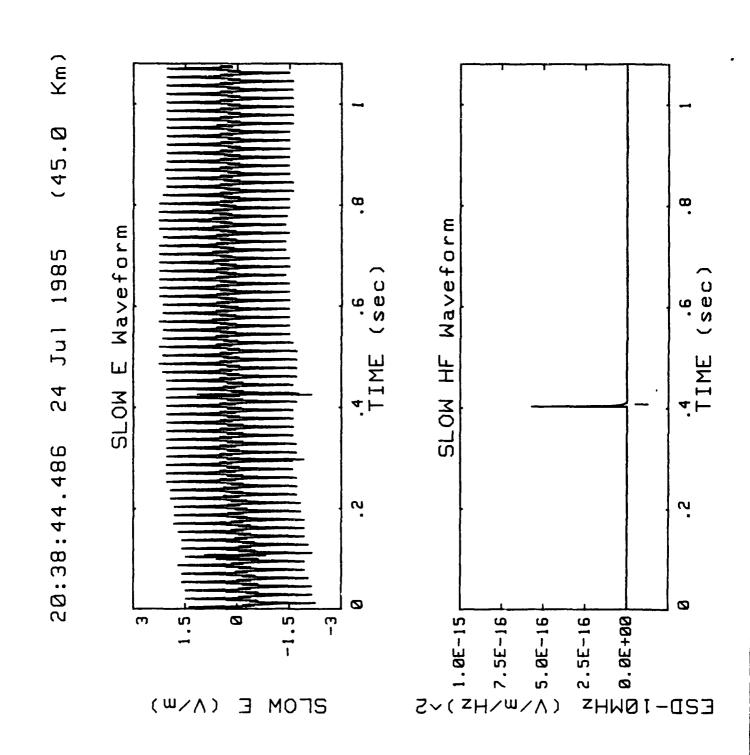


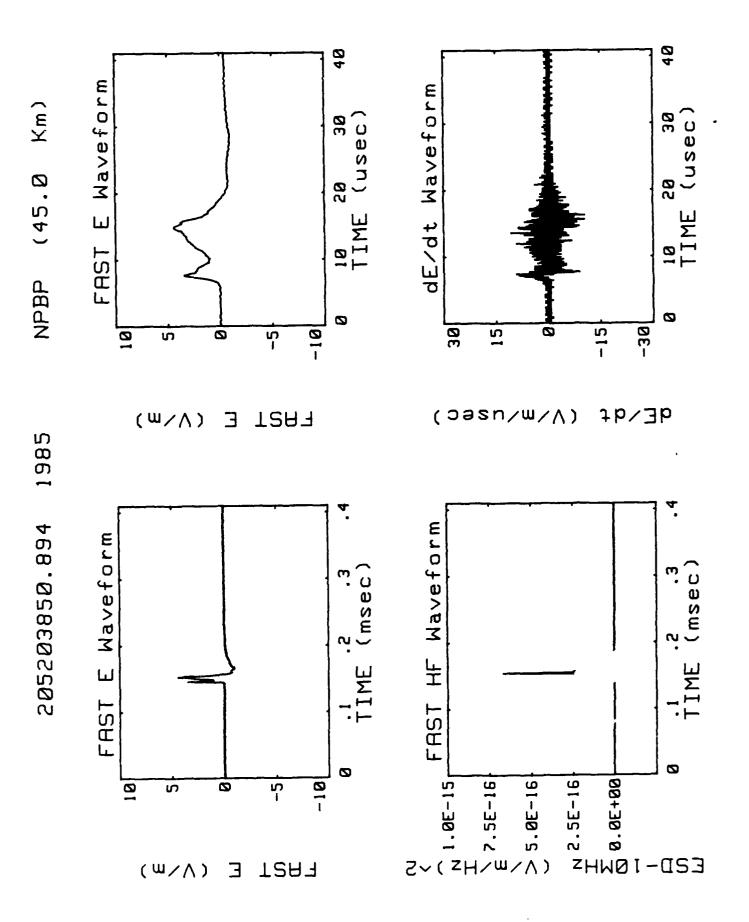


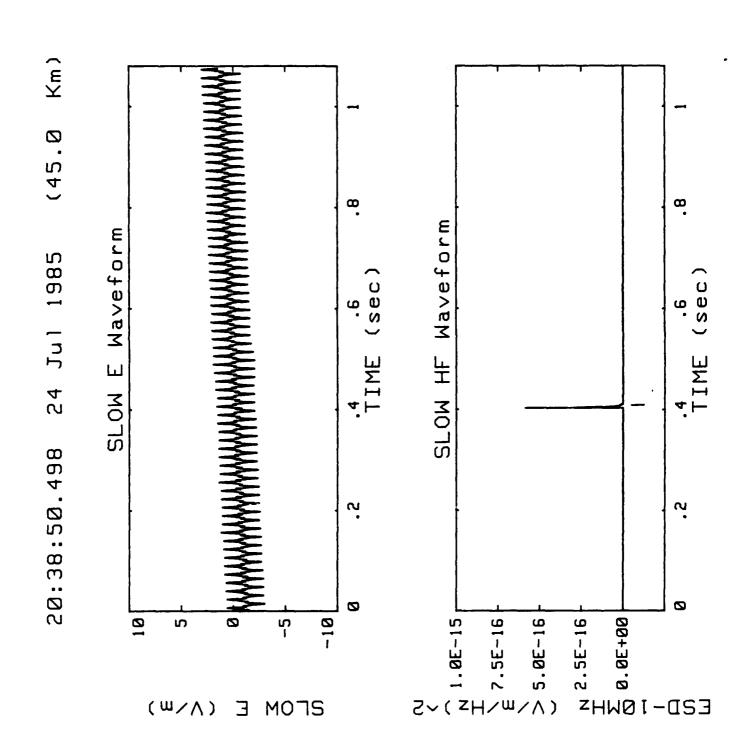


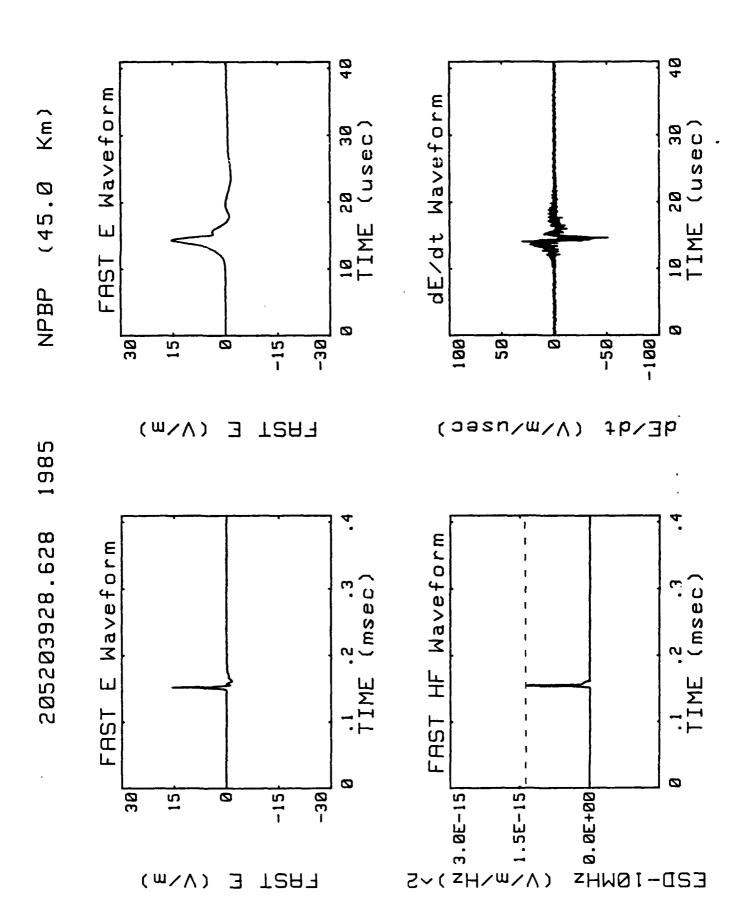


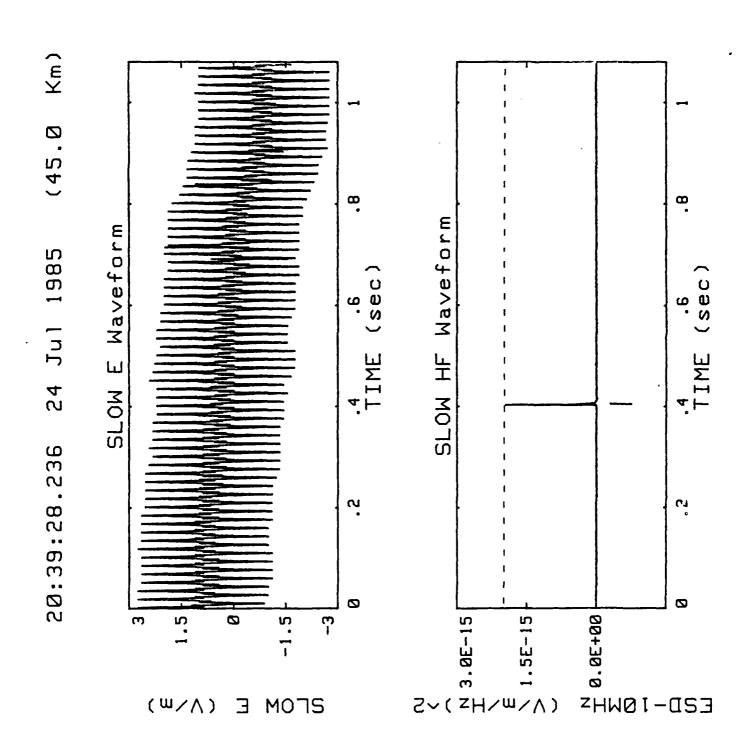


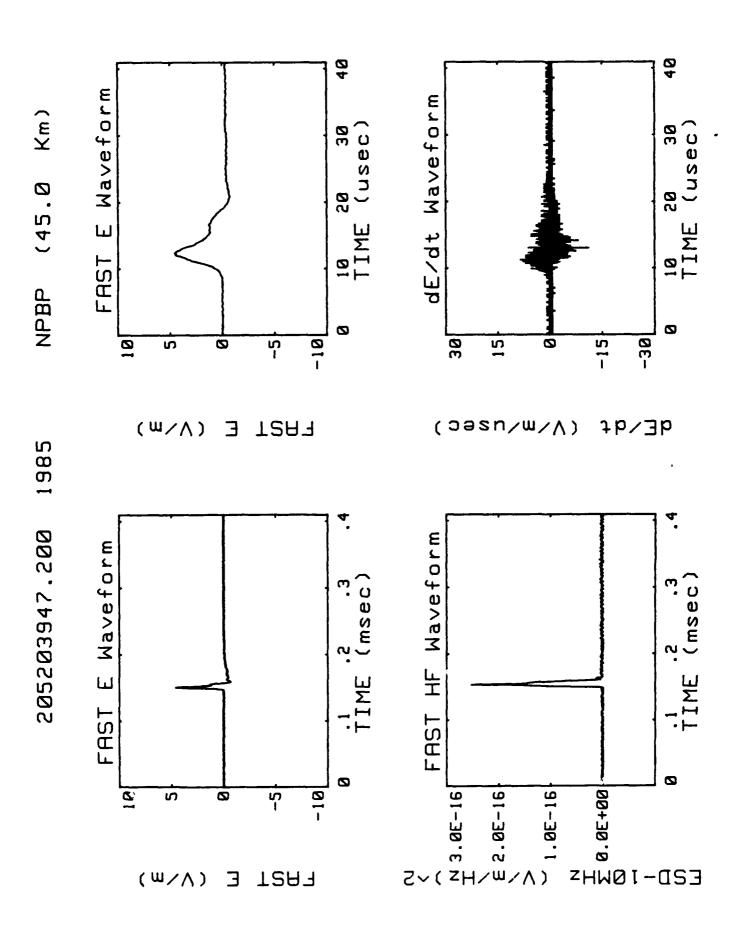


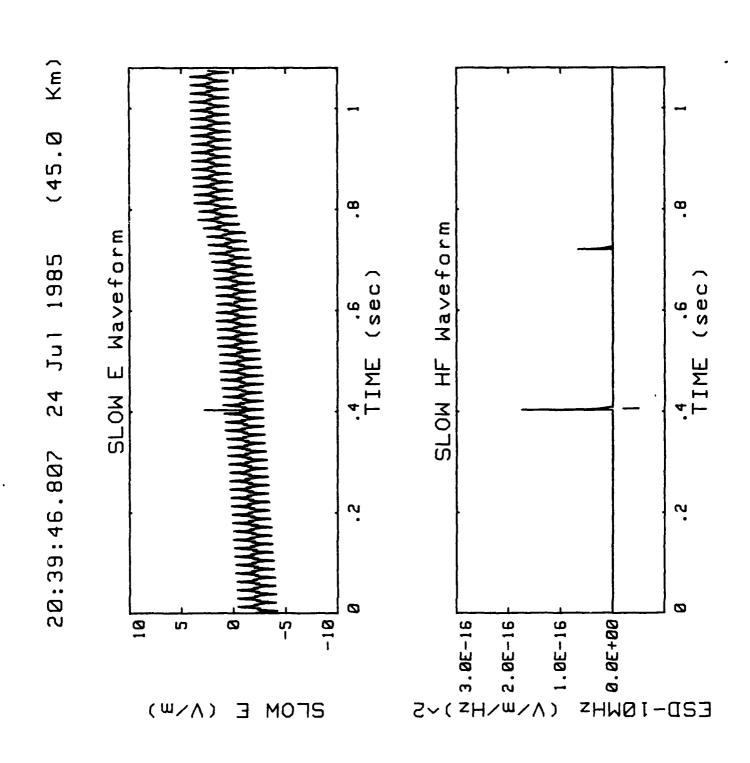


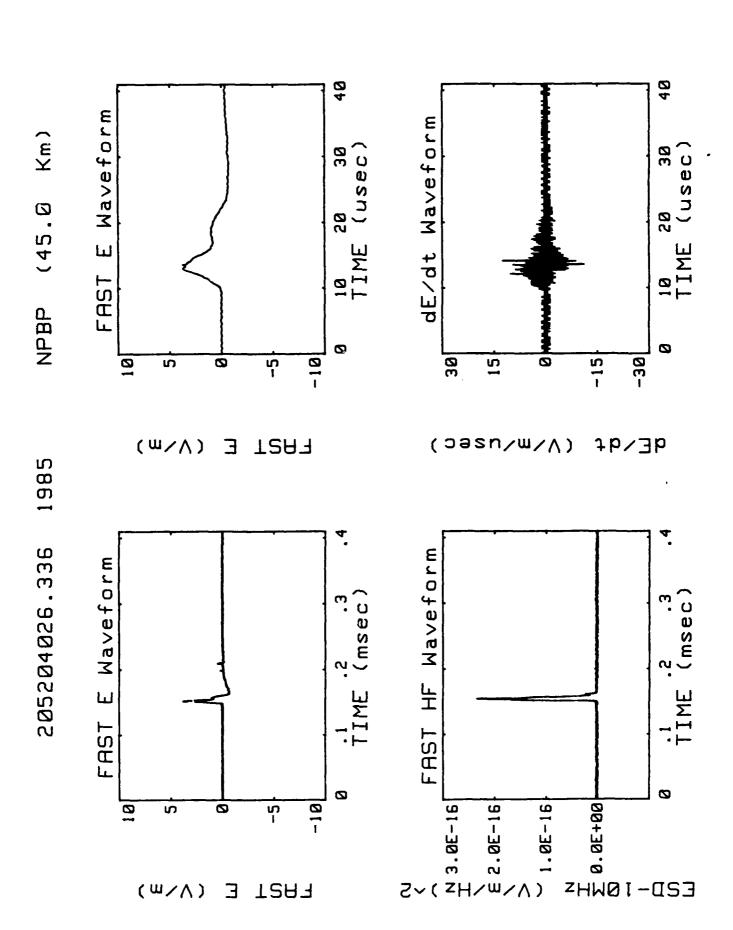


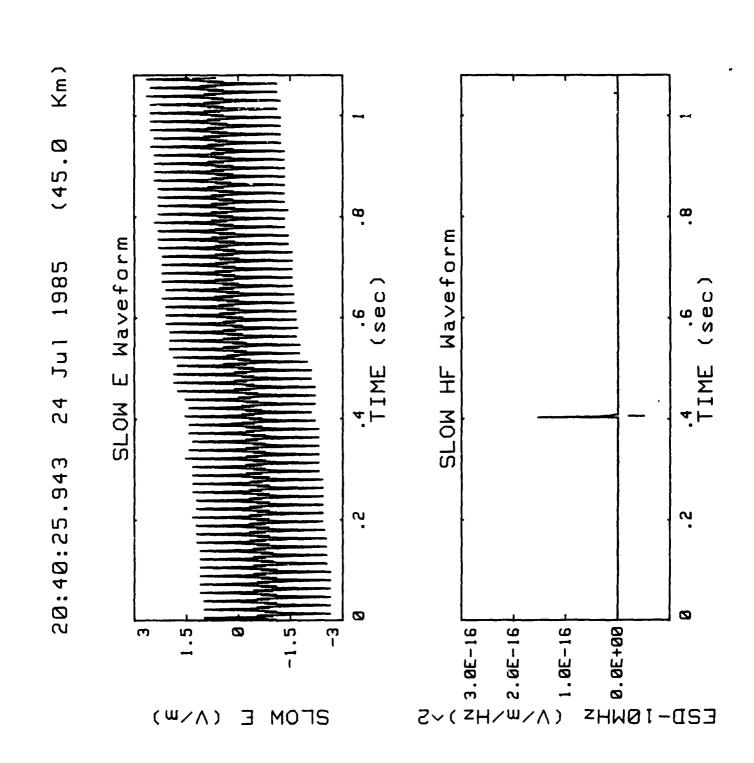


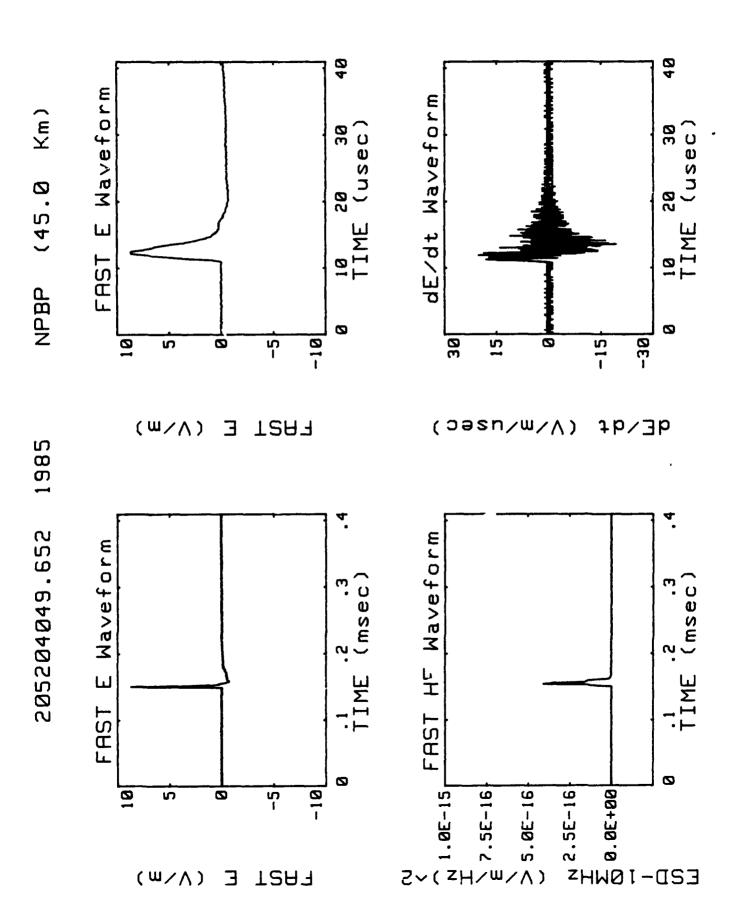


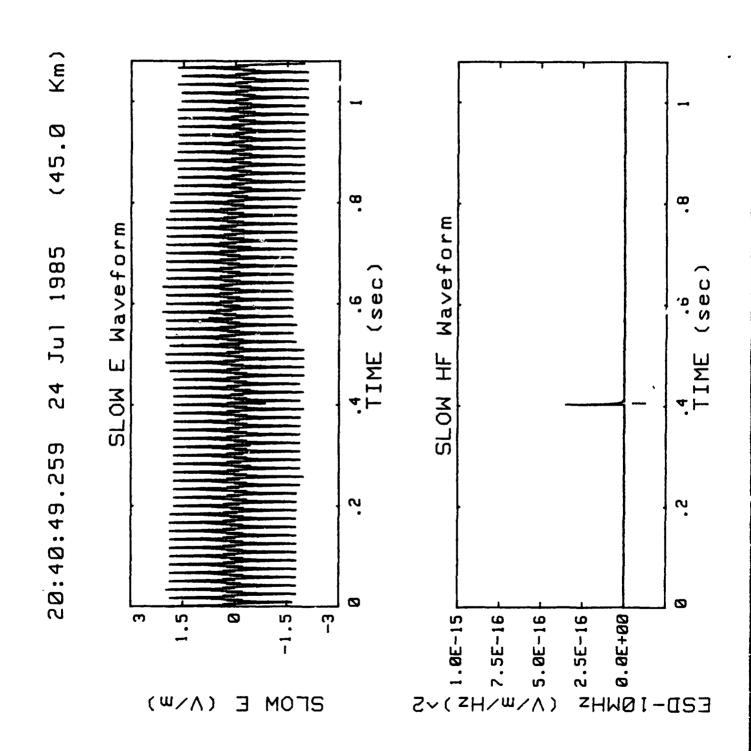


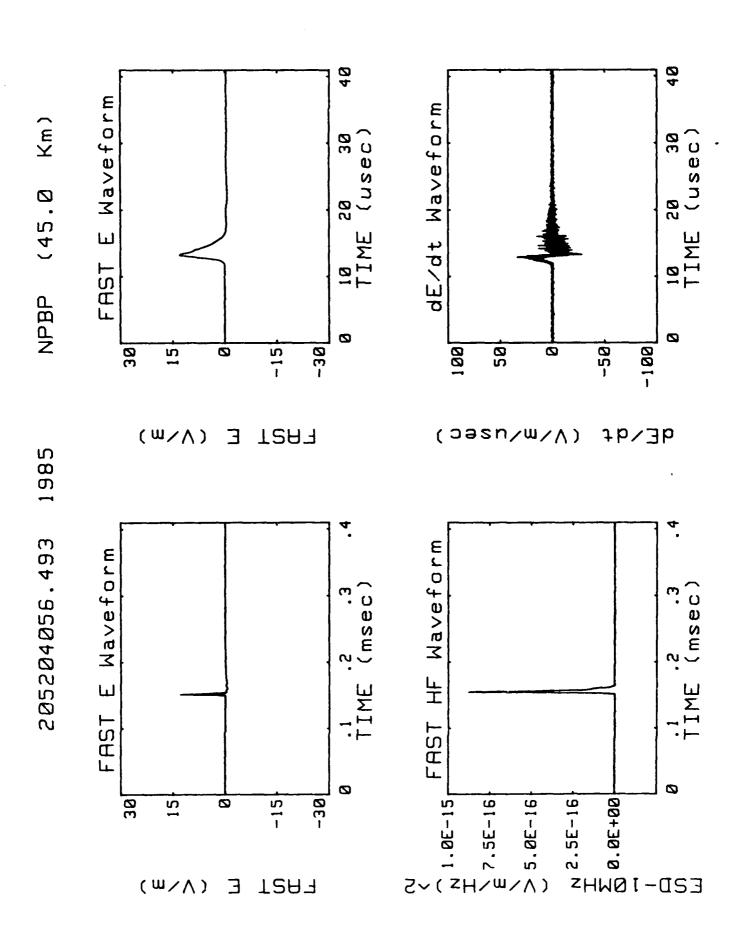


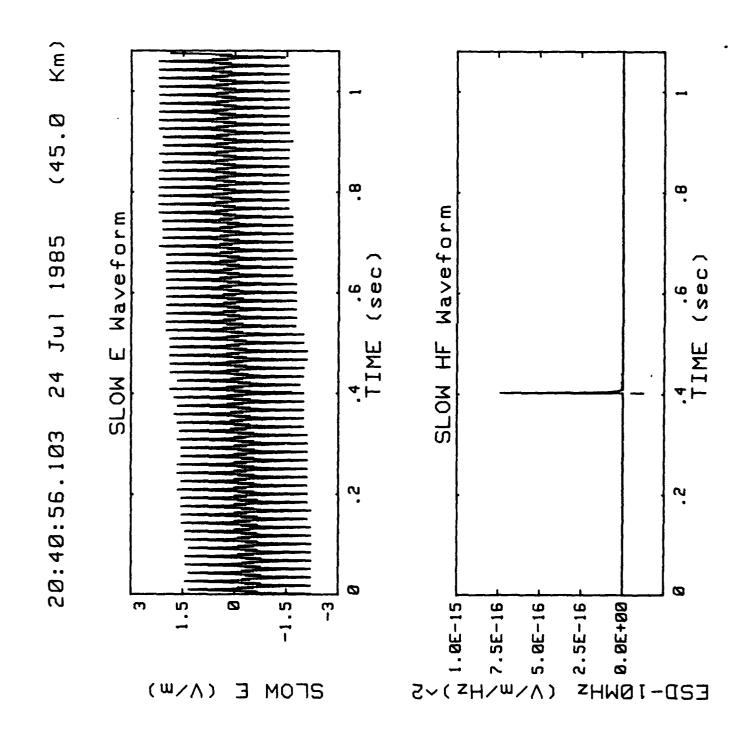


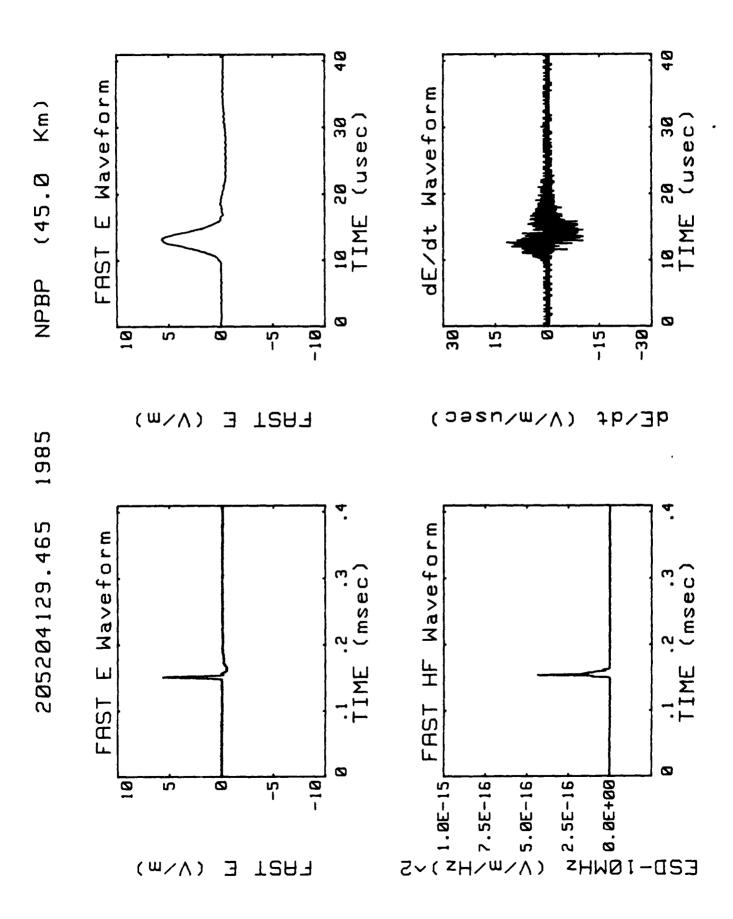


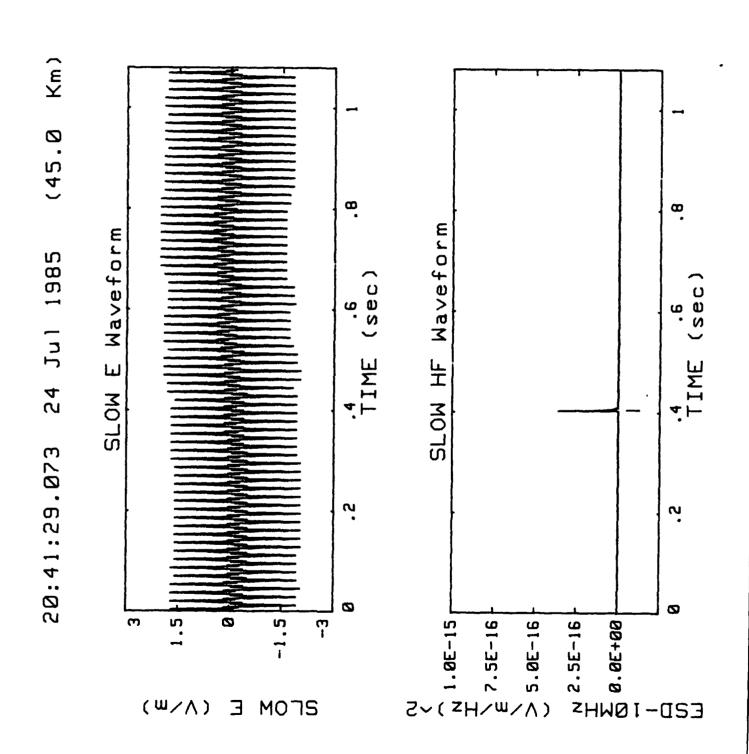


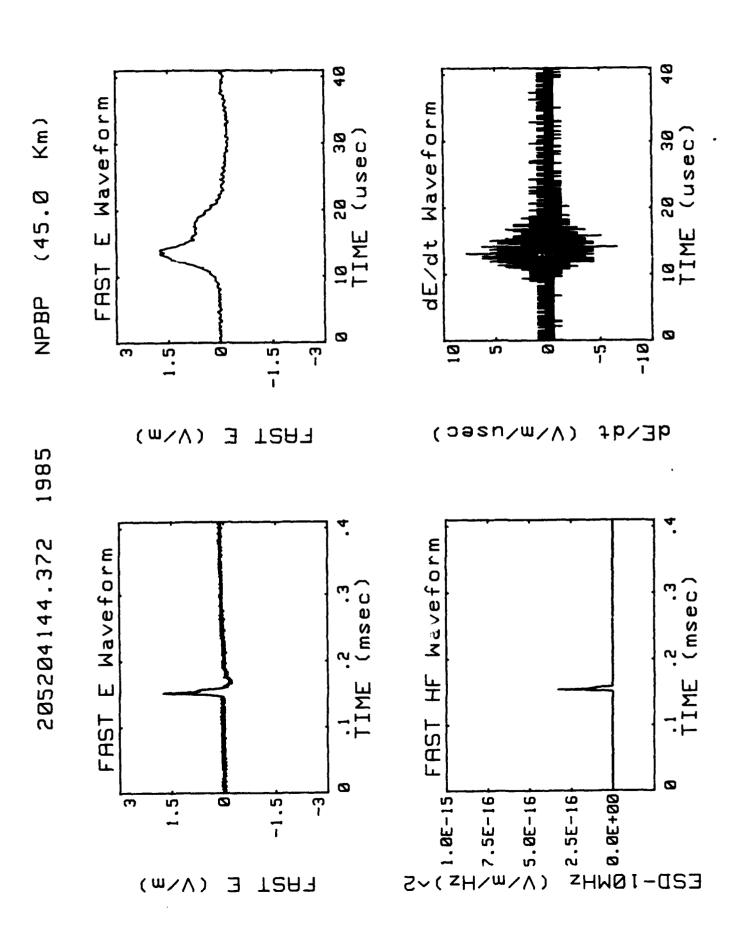


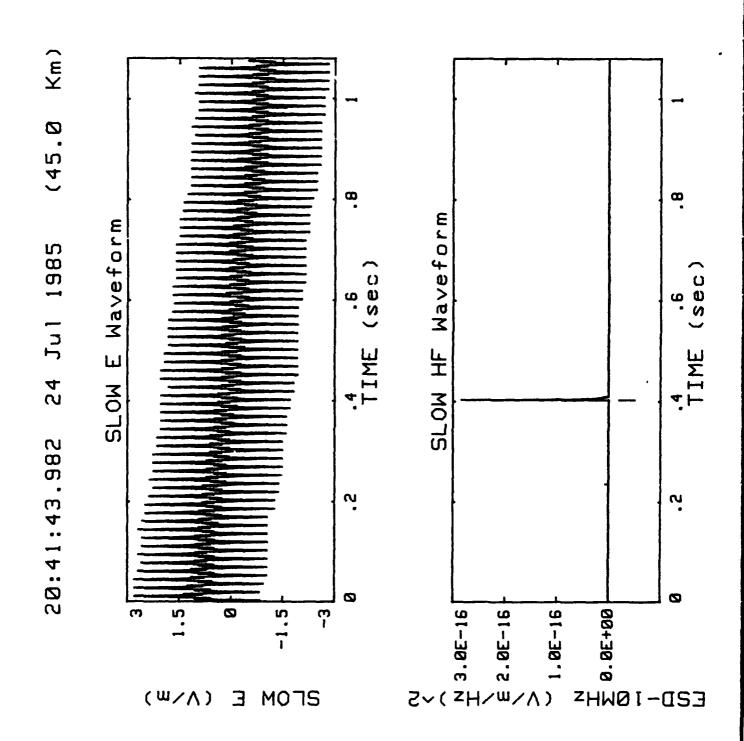


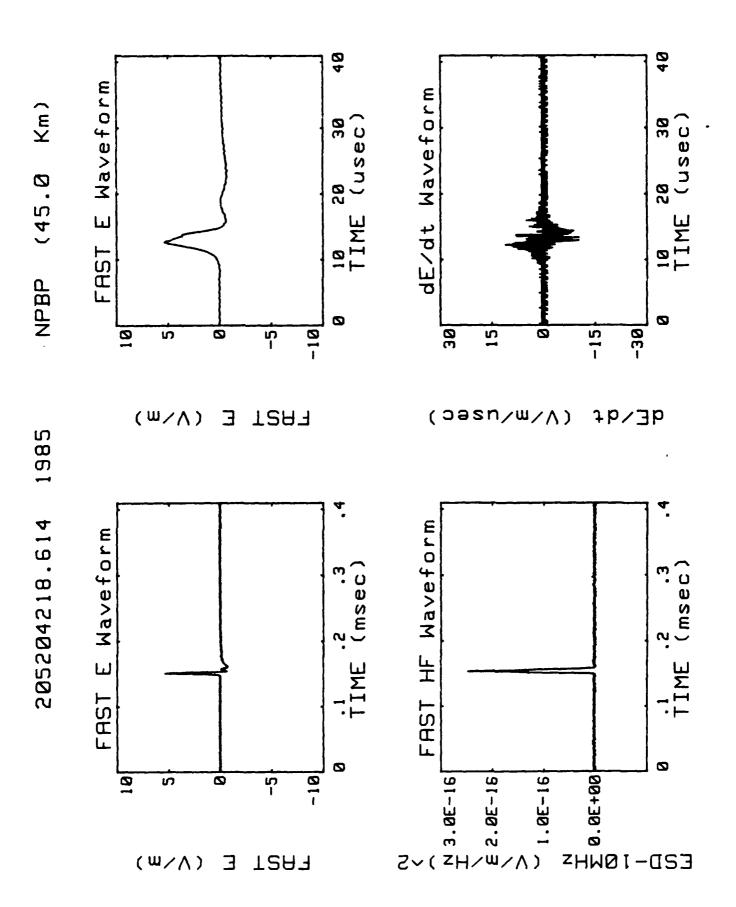


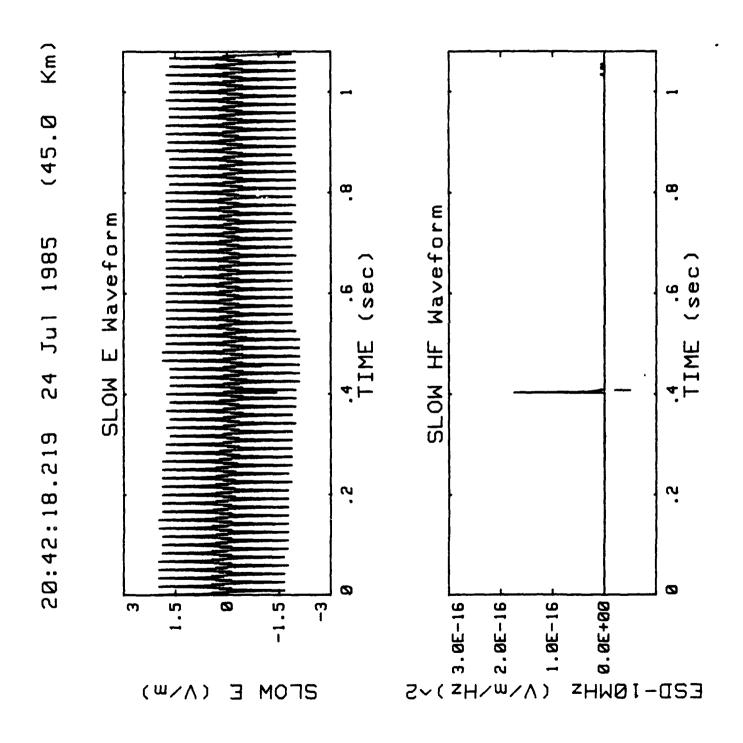


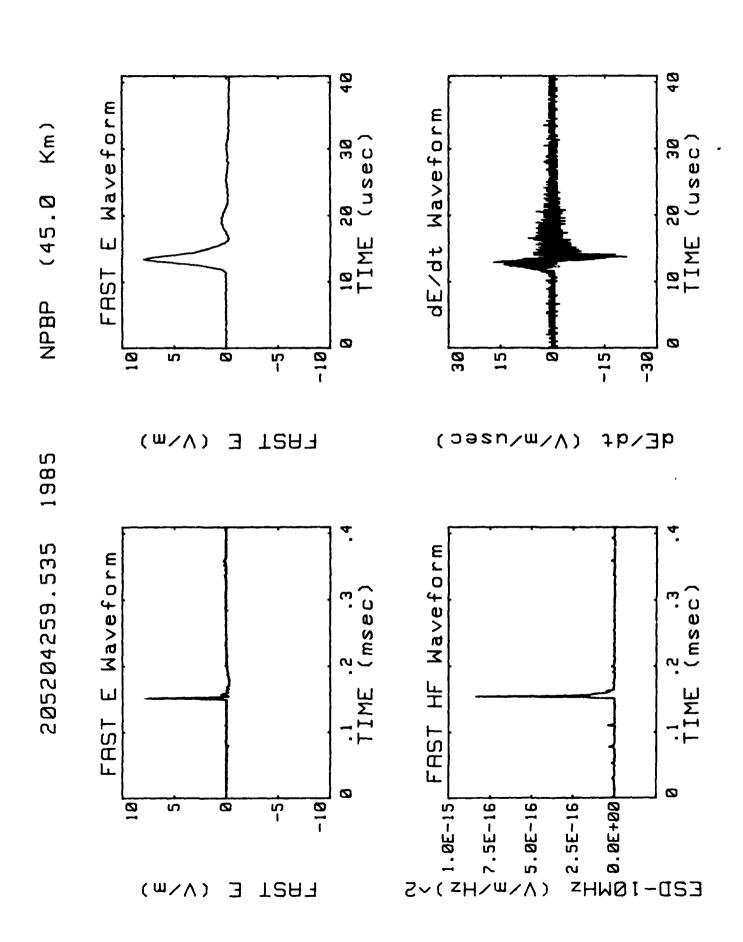


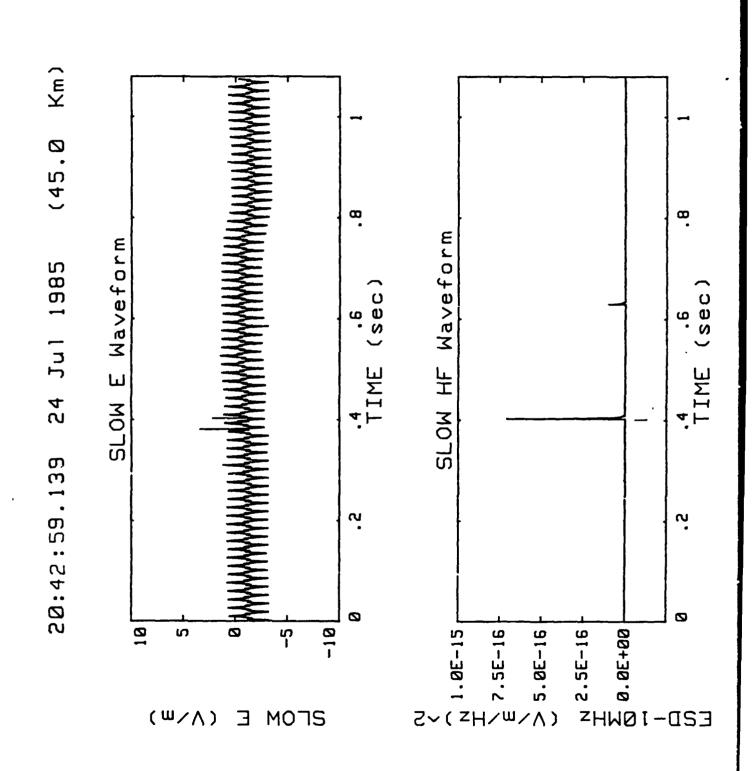


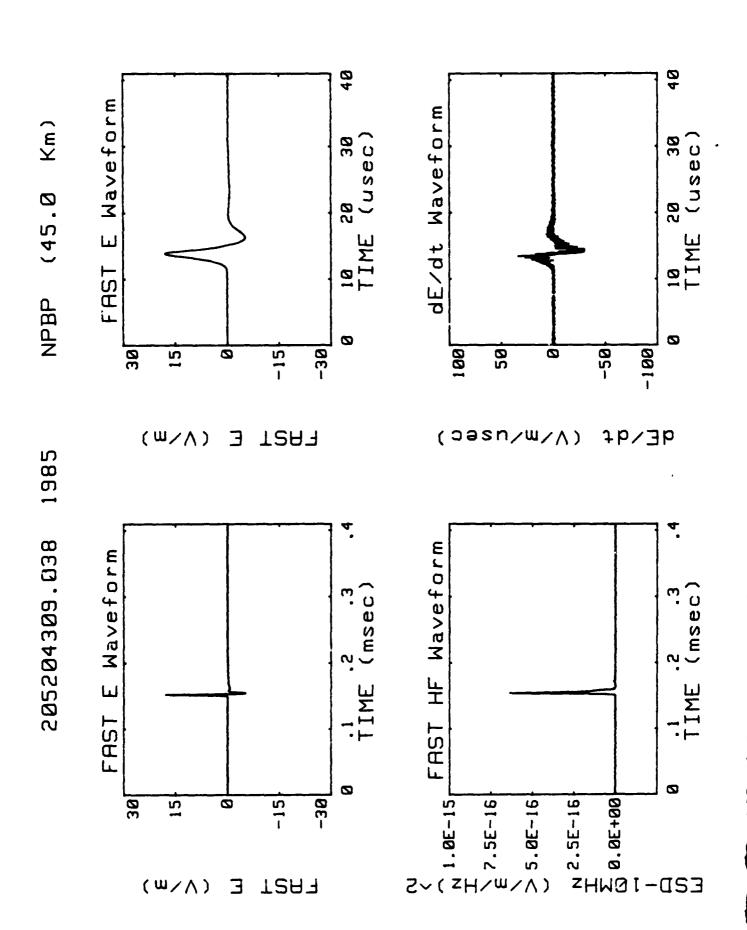


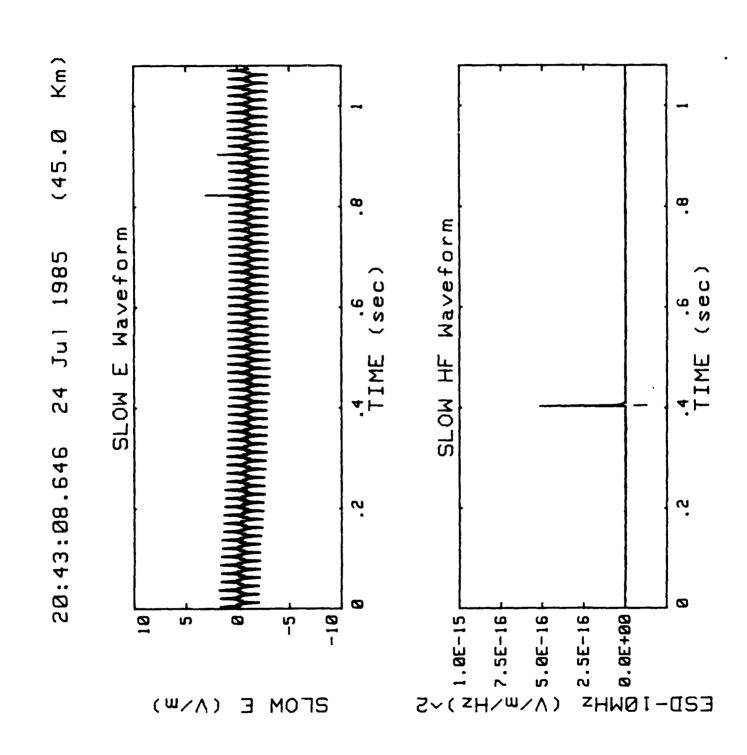


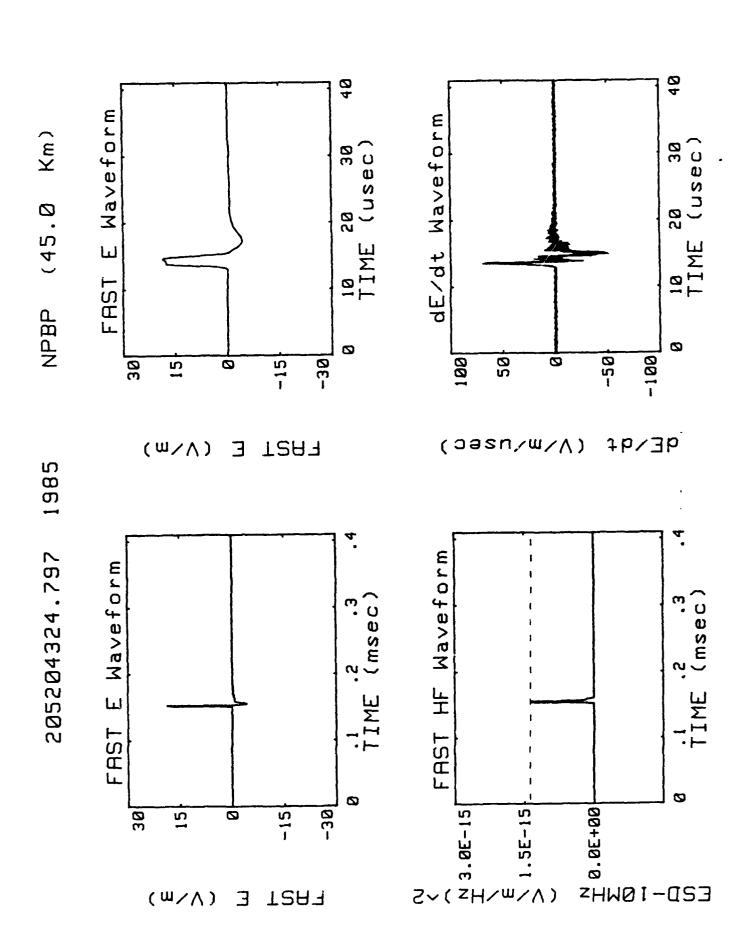


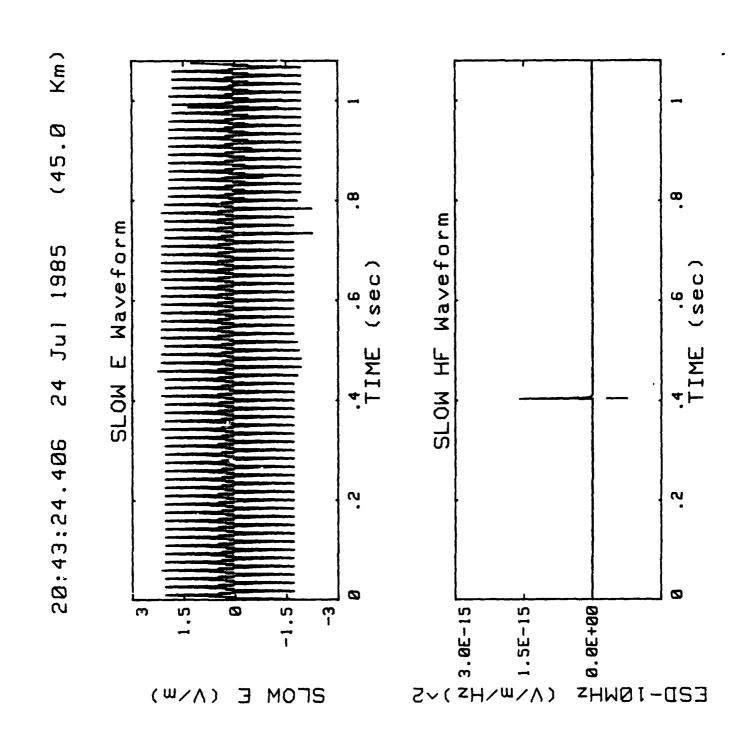


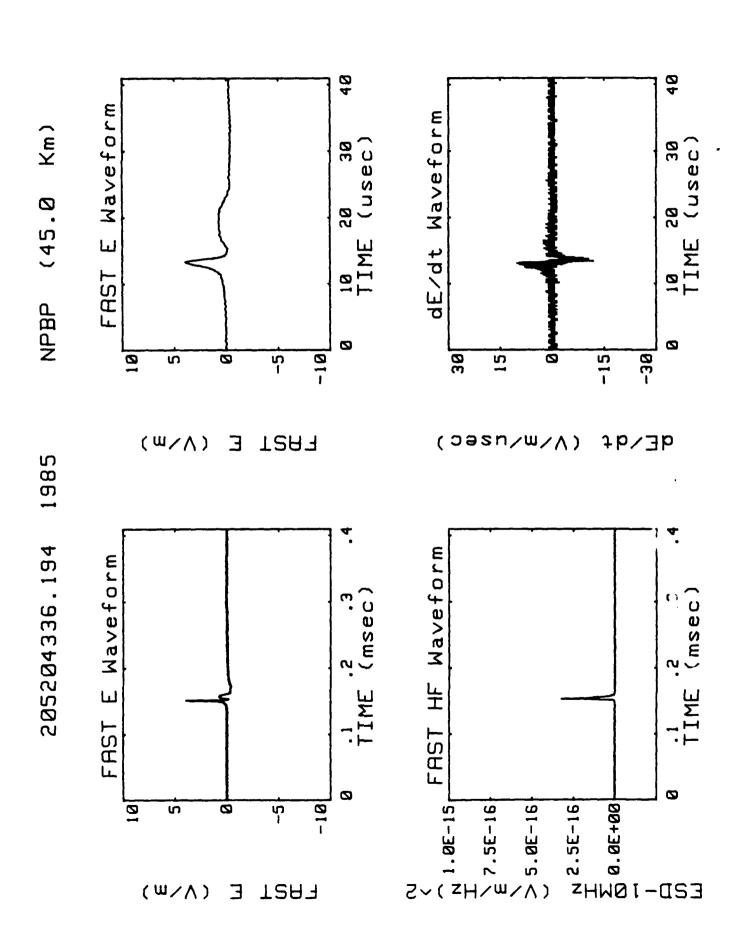


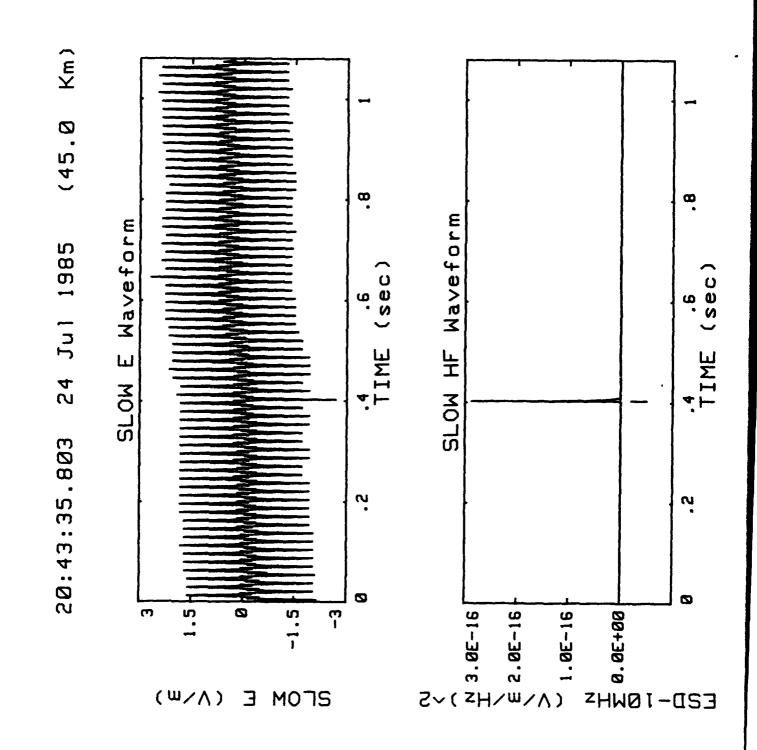


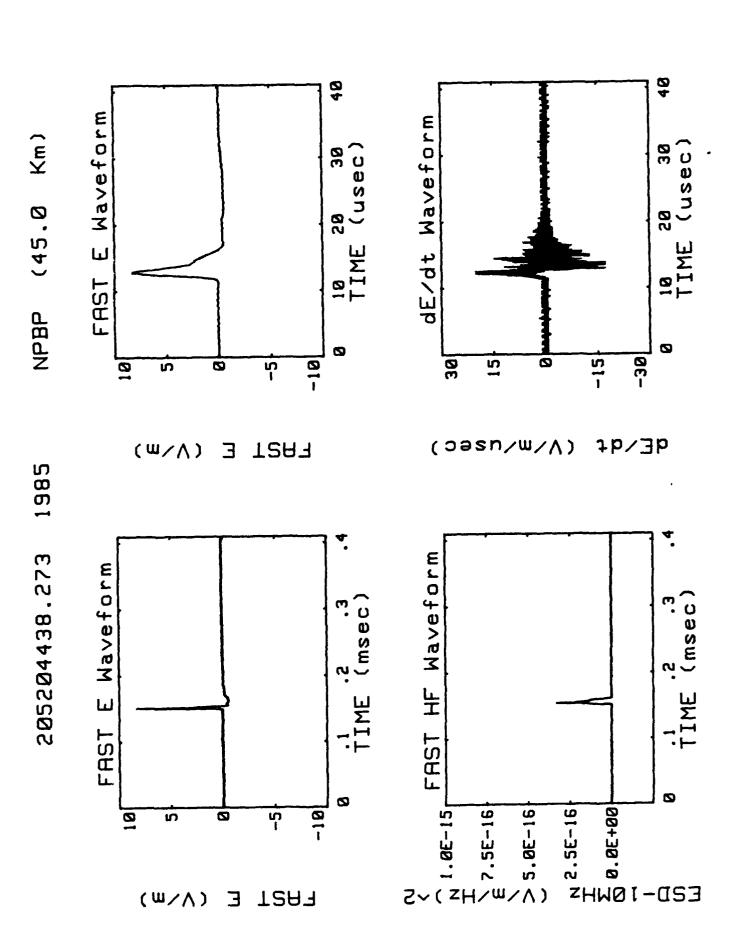


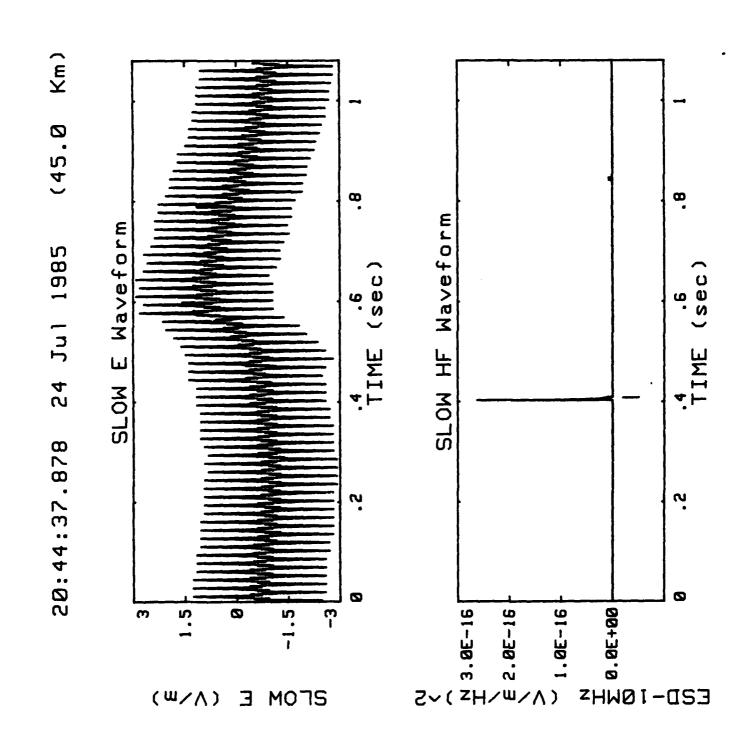


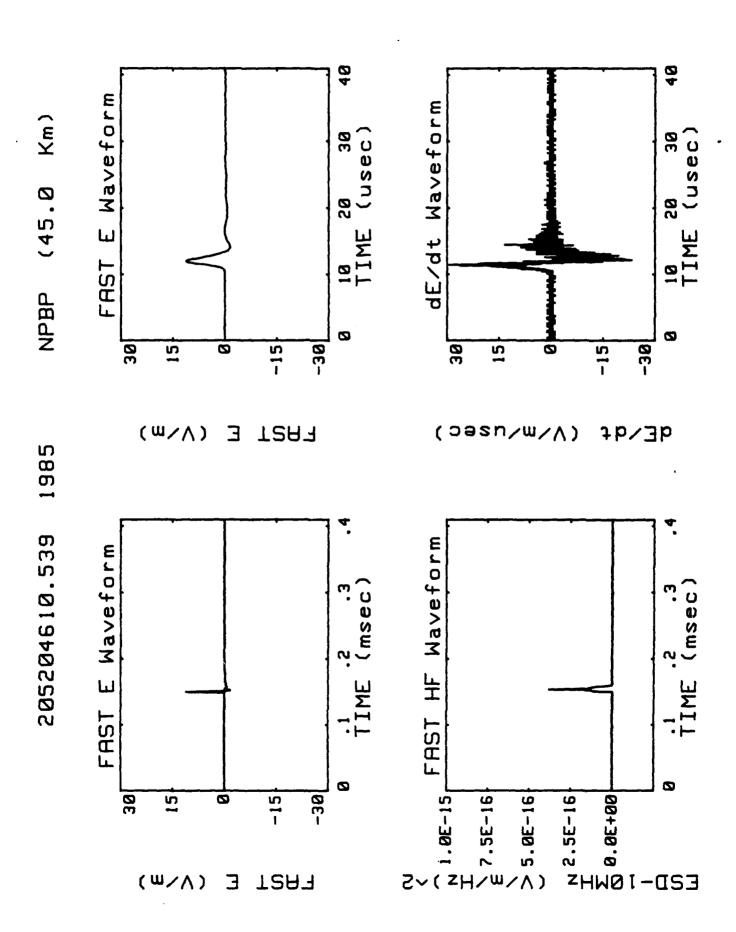


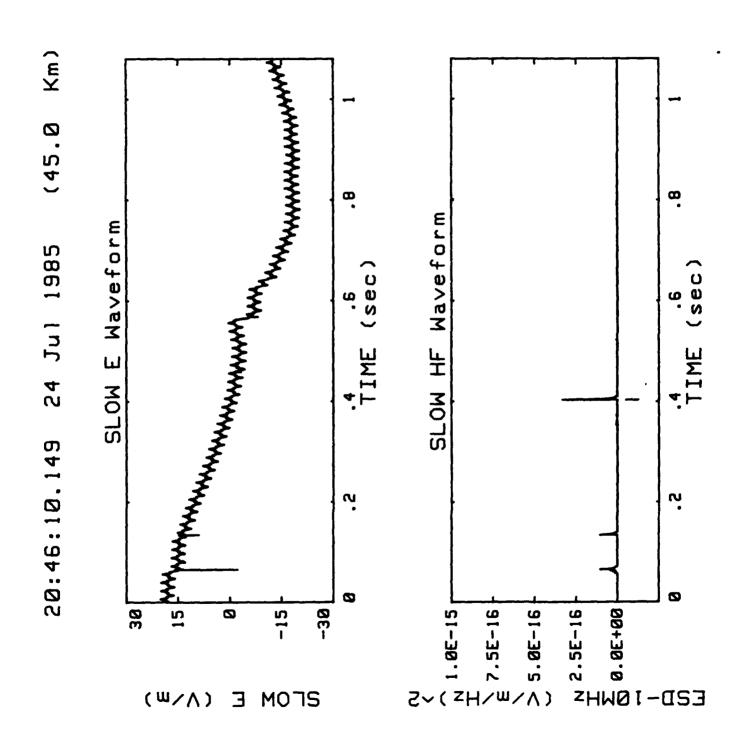


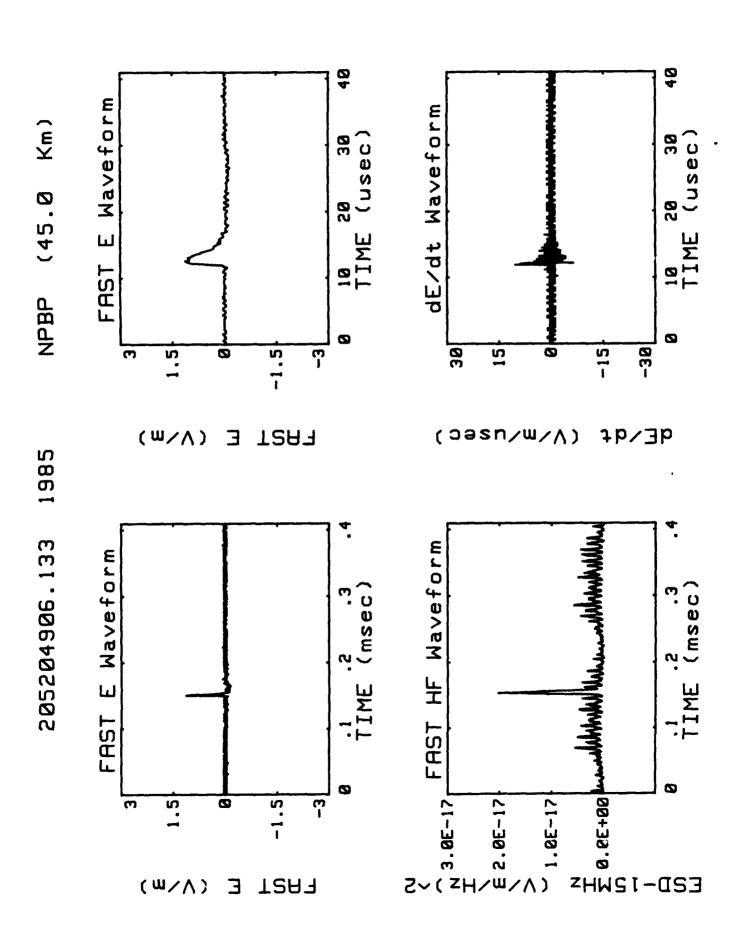


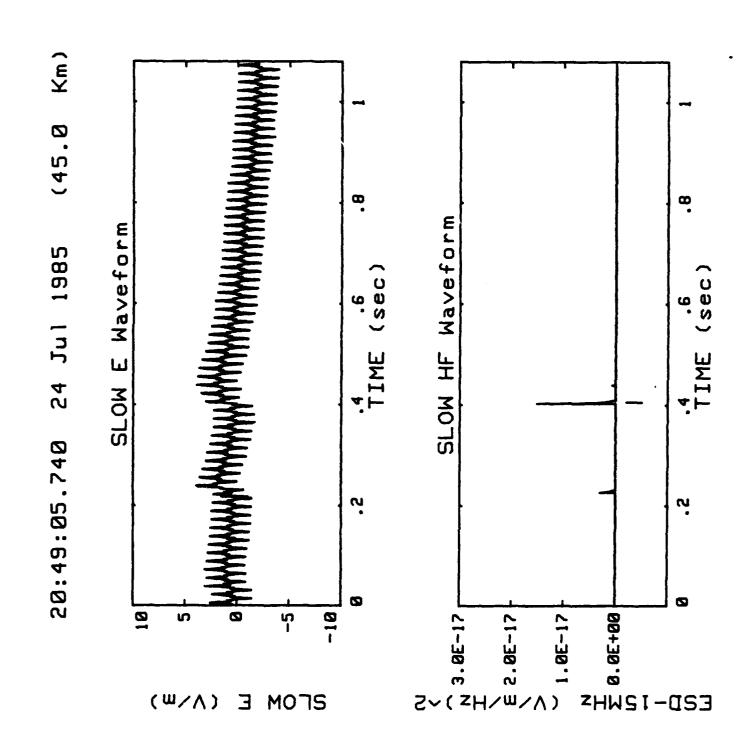




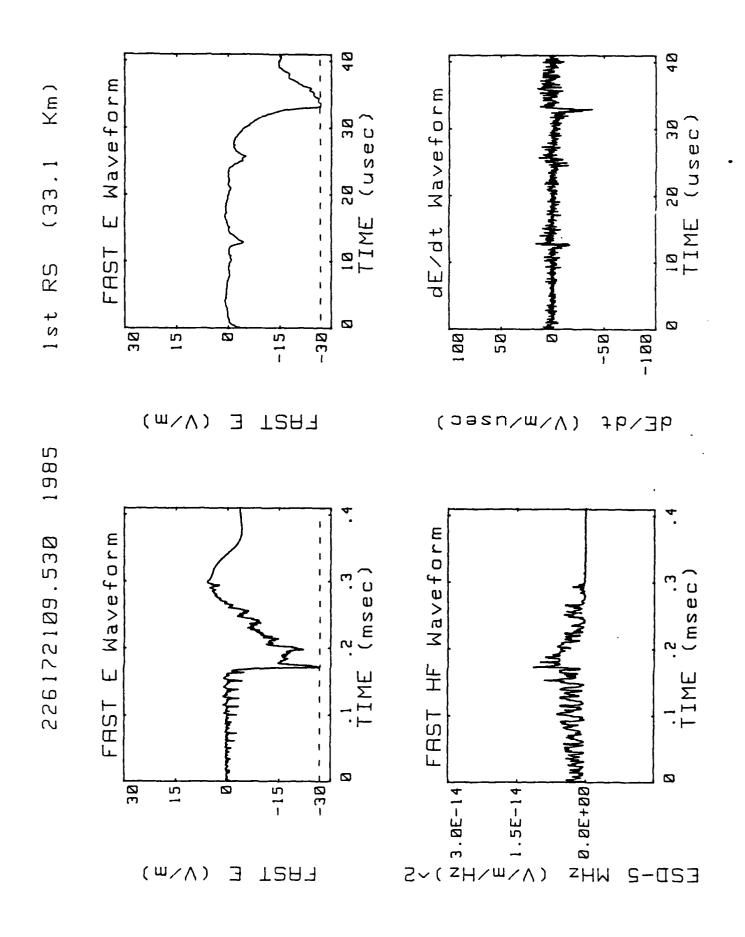


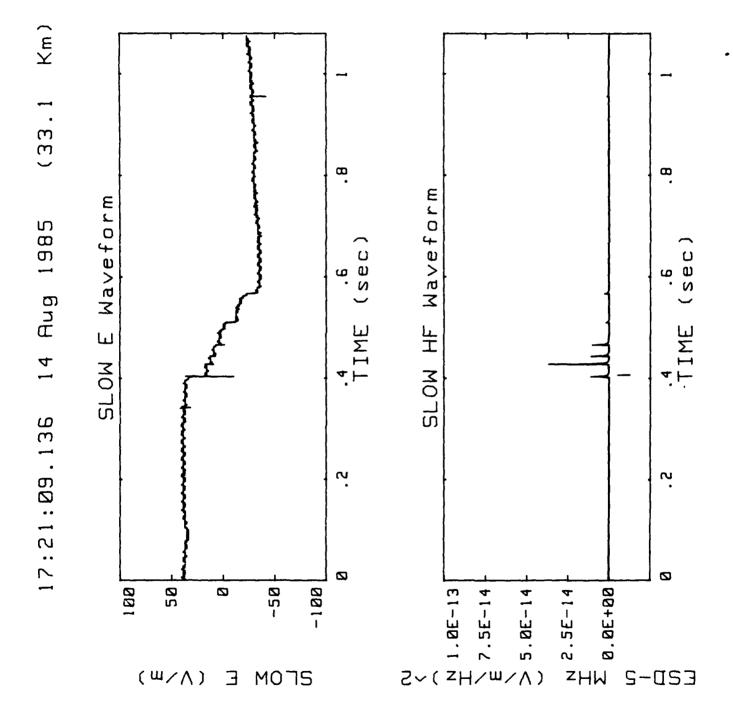


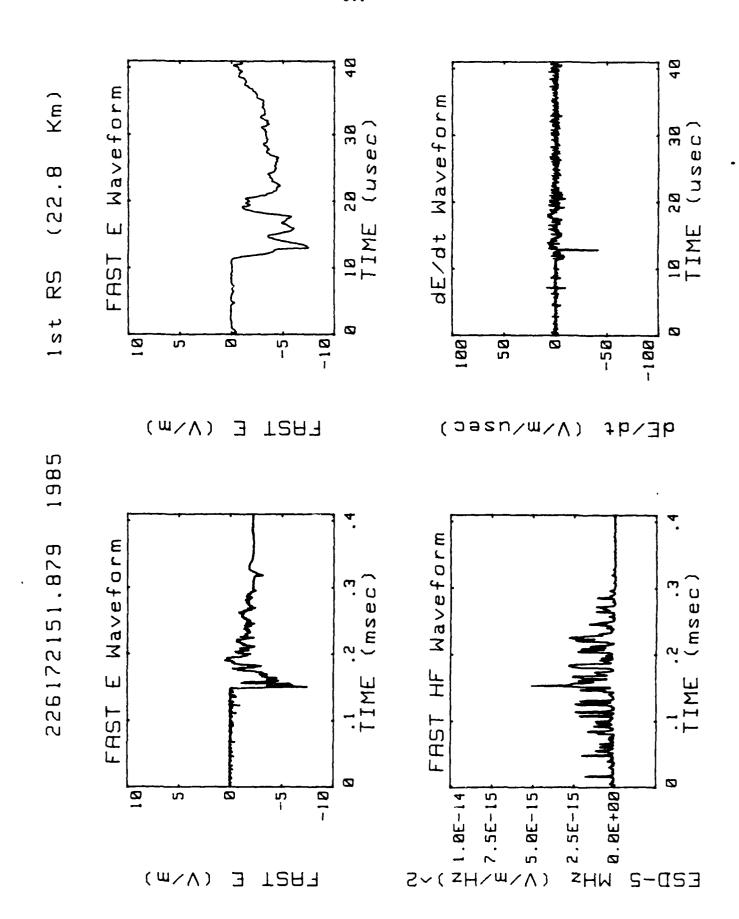


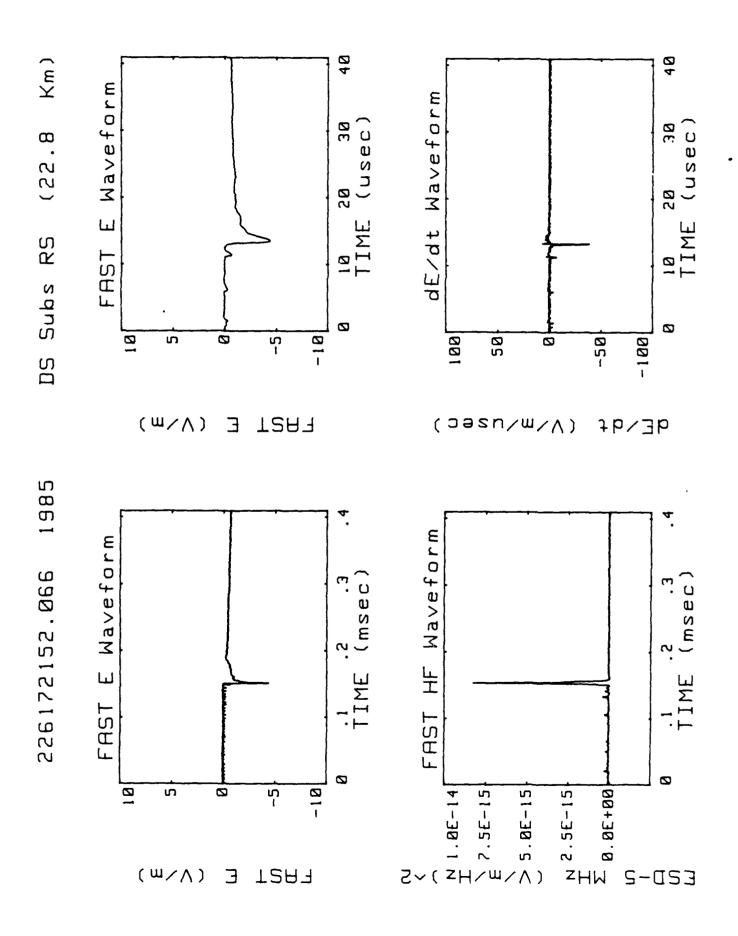


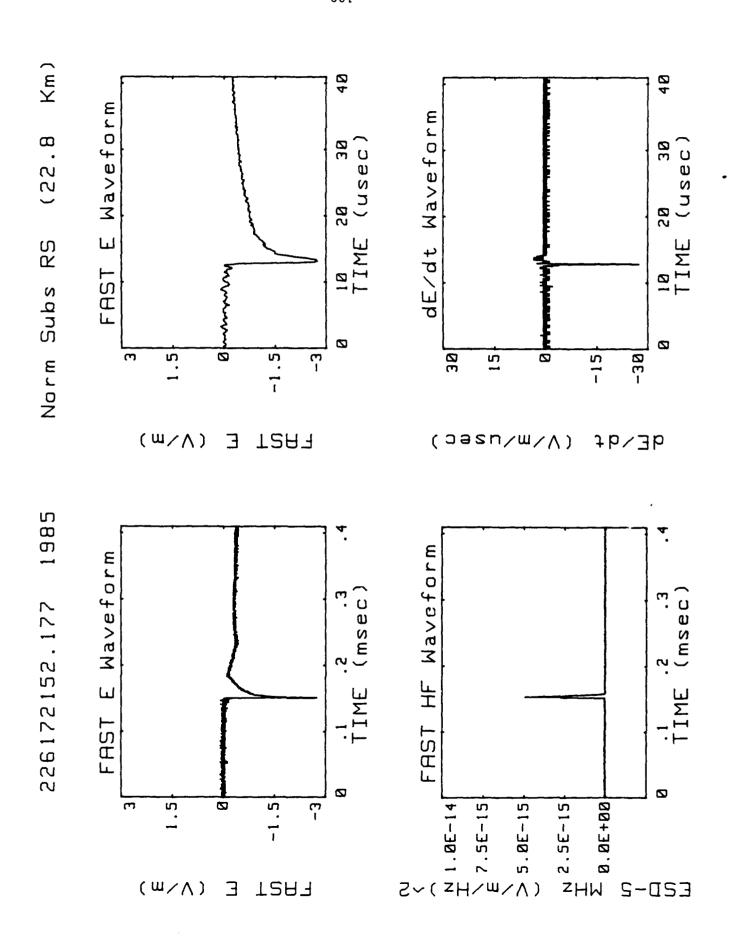
APPENDIX B
August 14 Data Plots

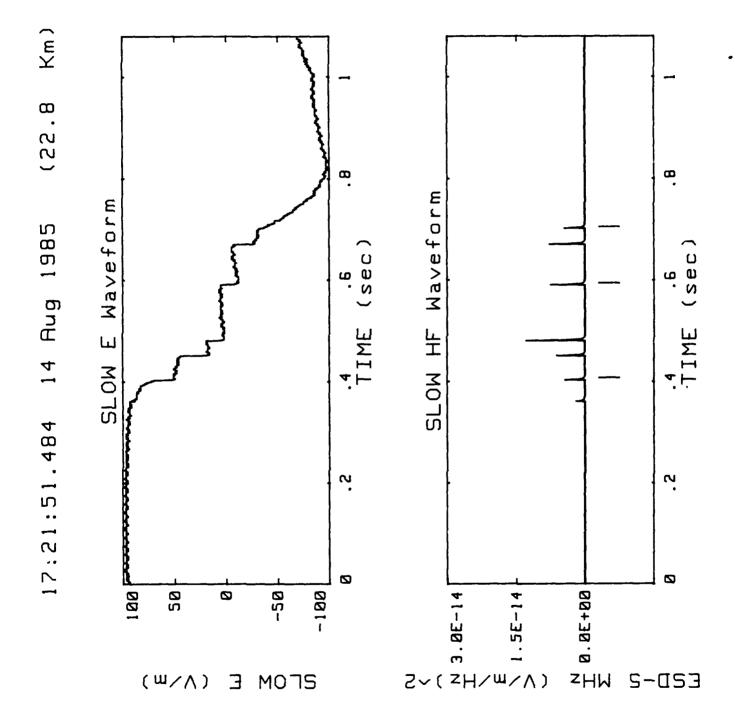


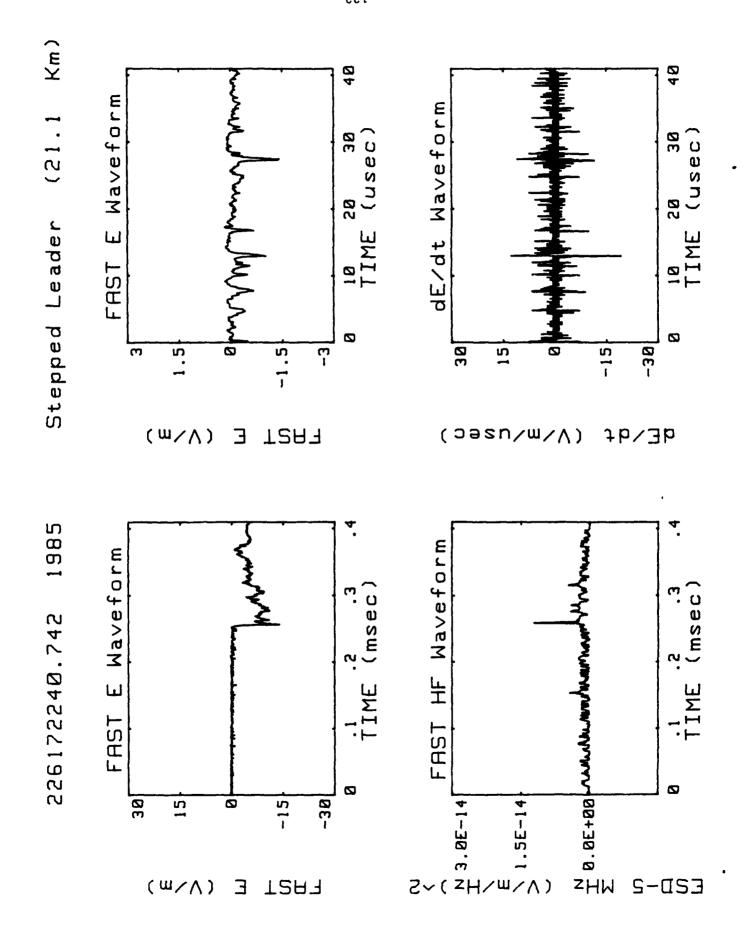


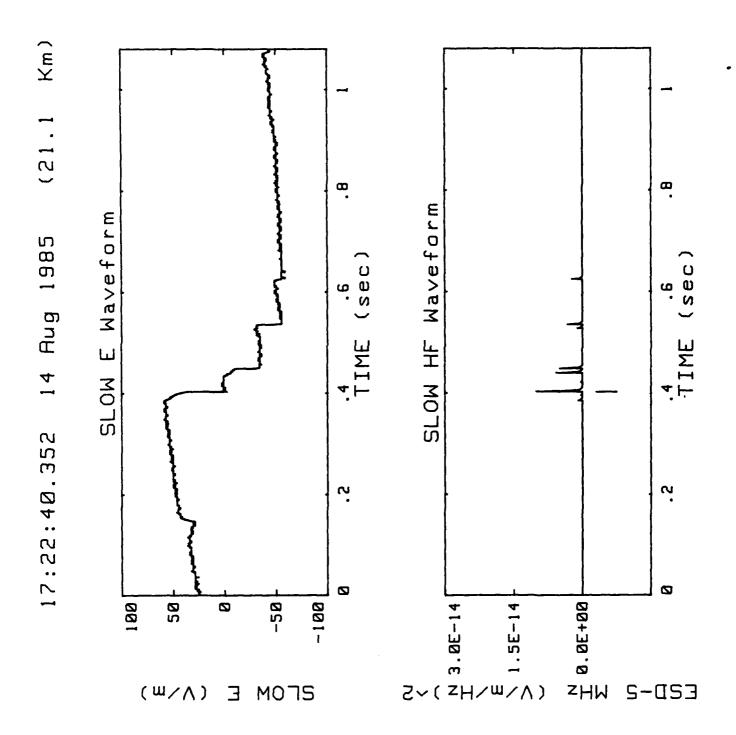


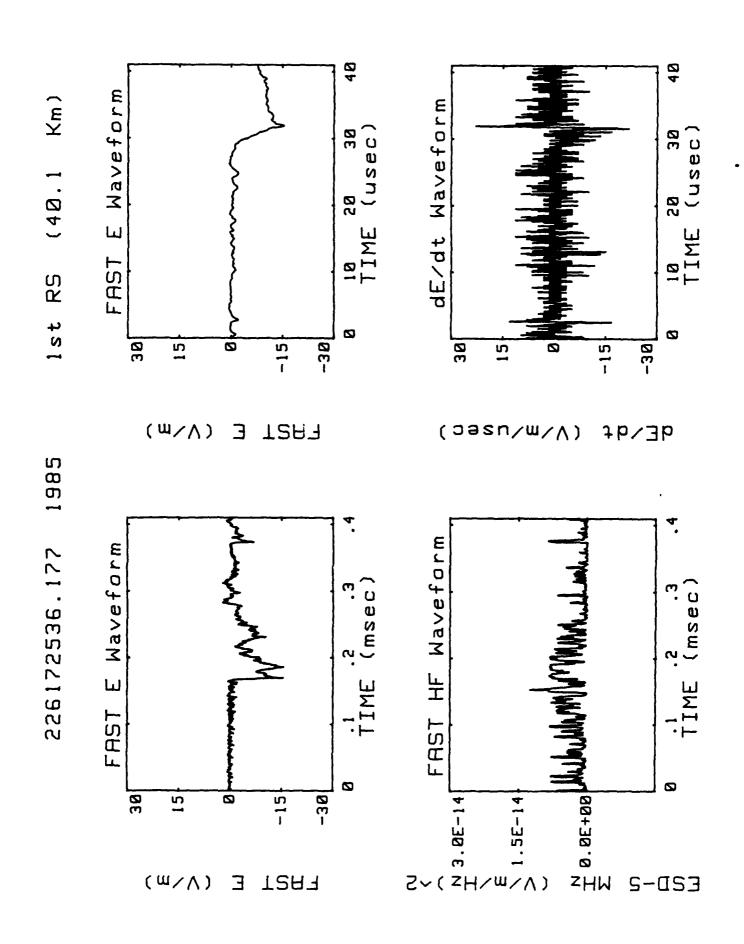


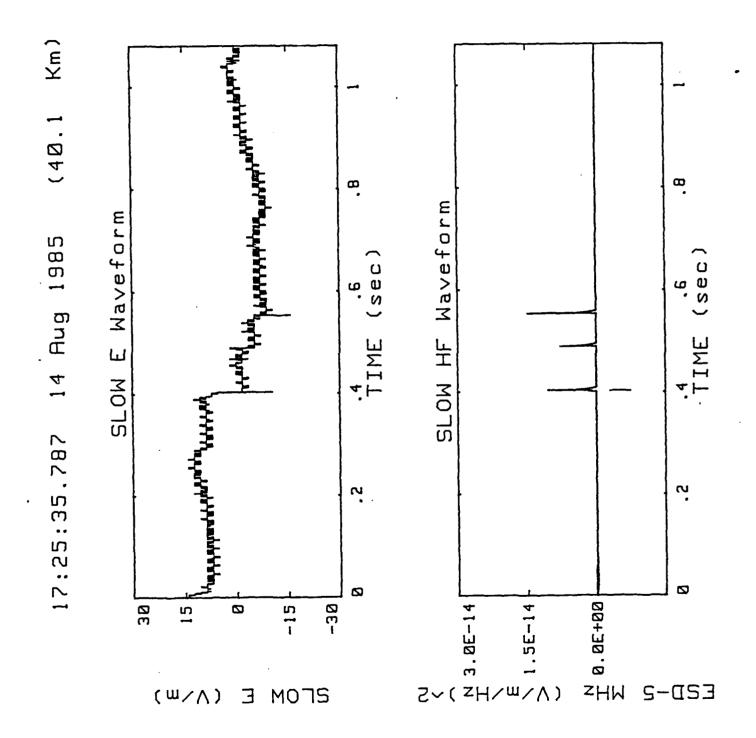


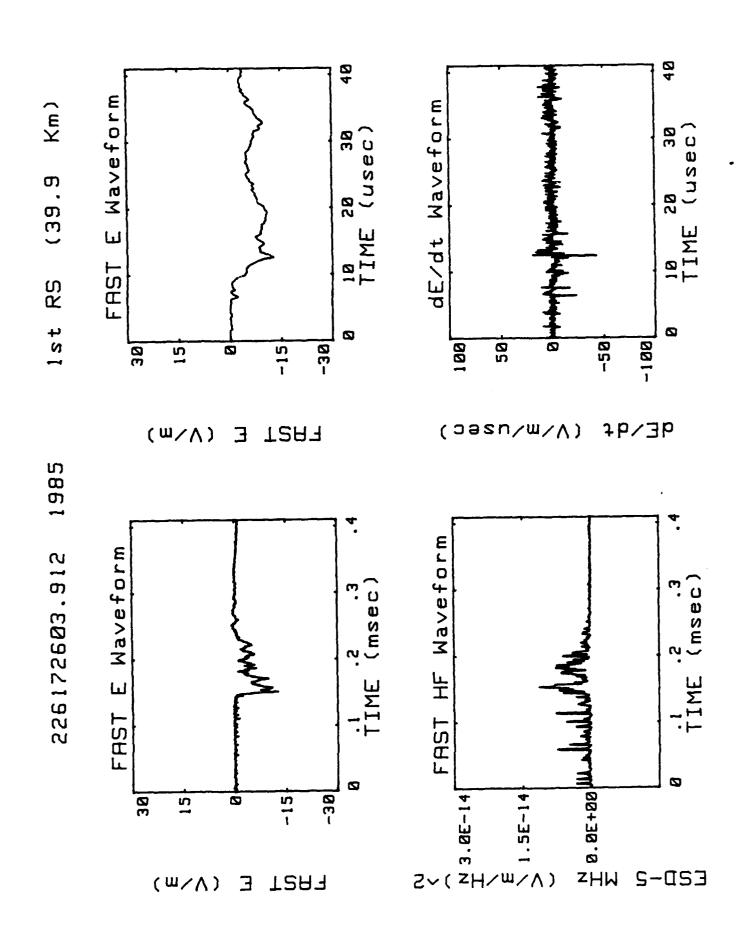


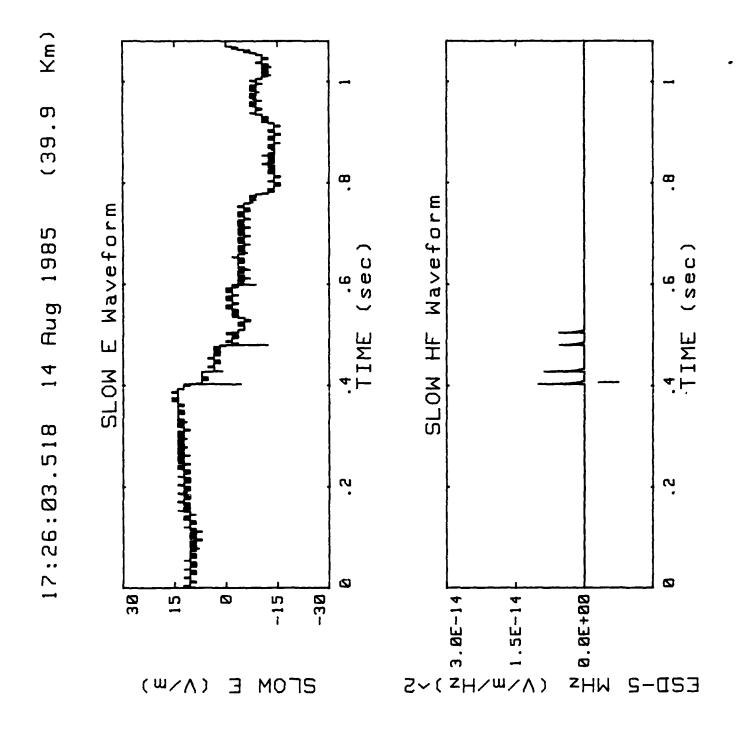


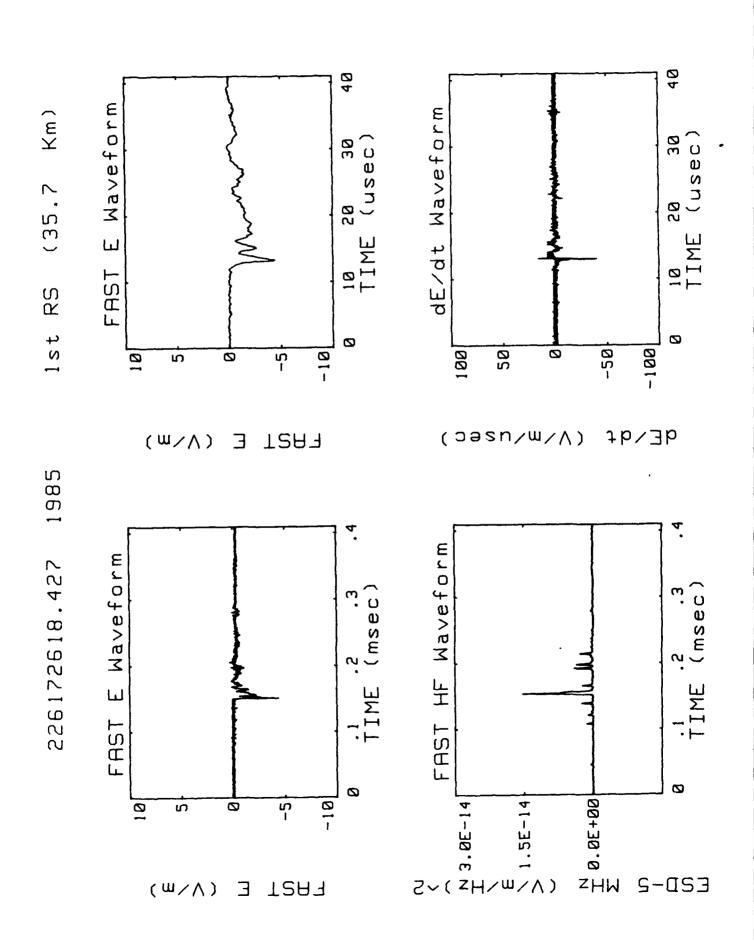


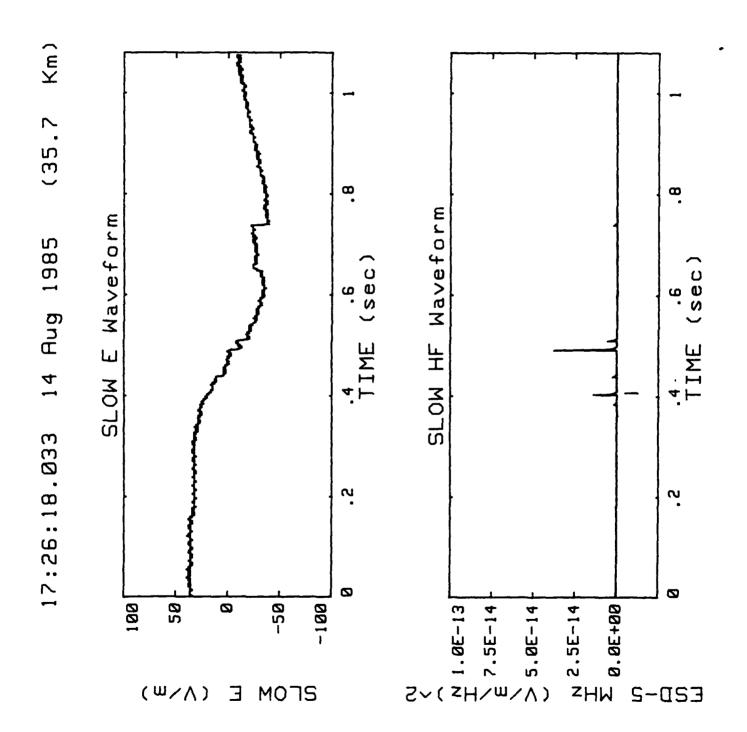


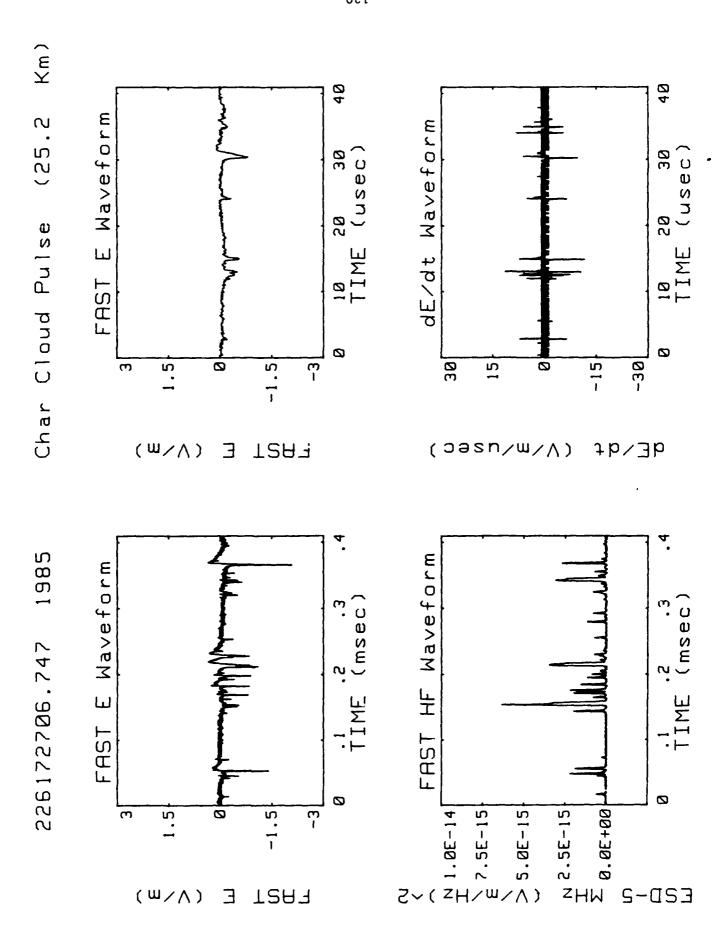


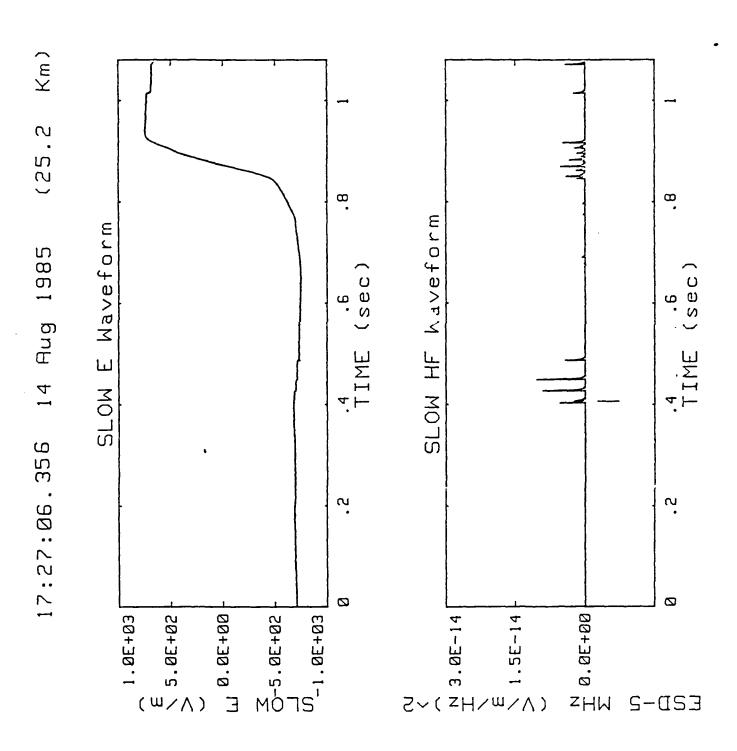


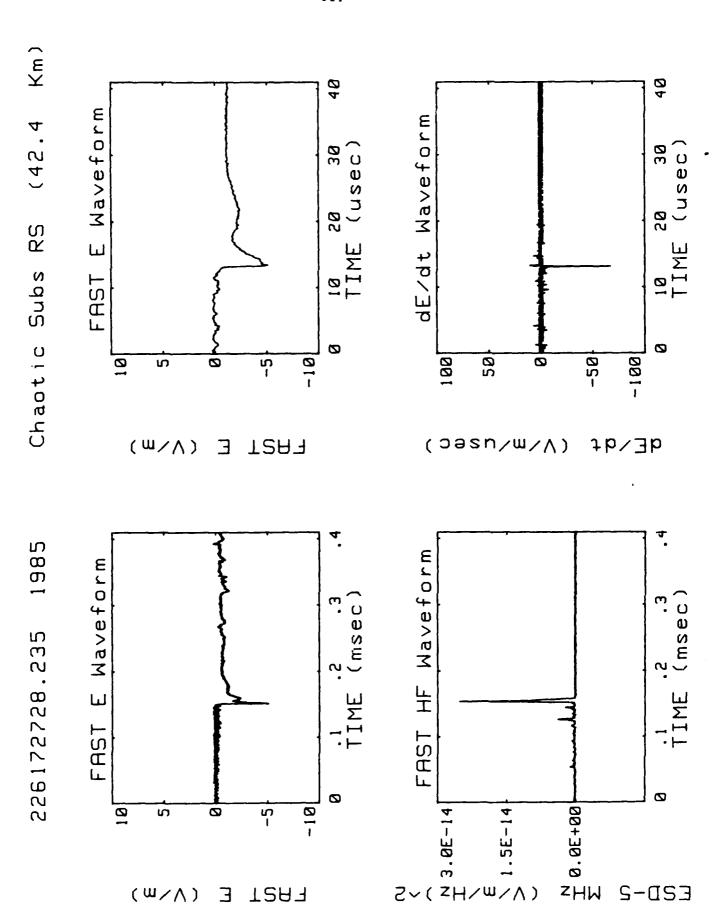


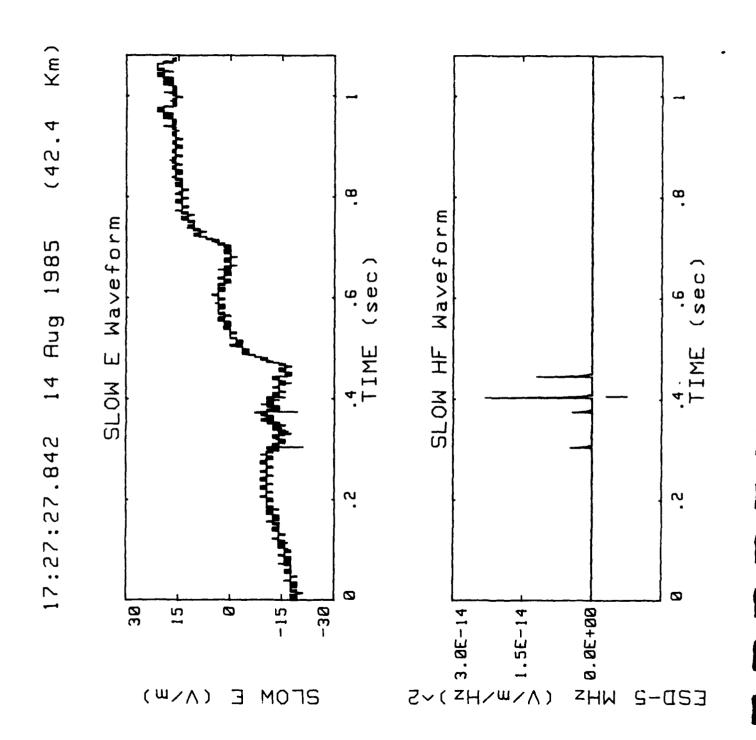


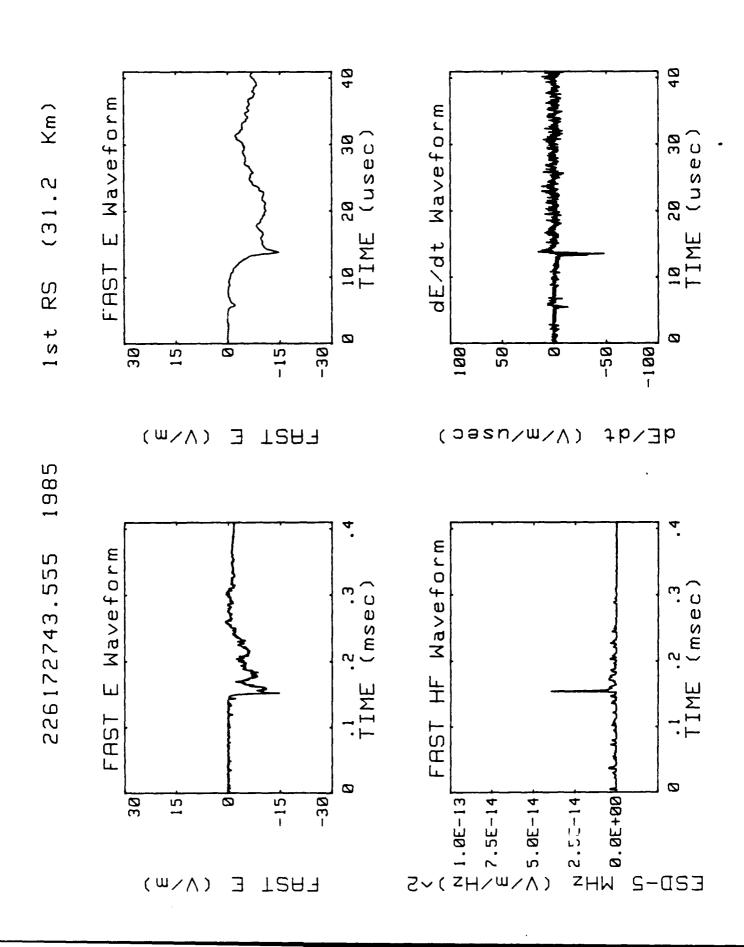


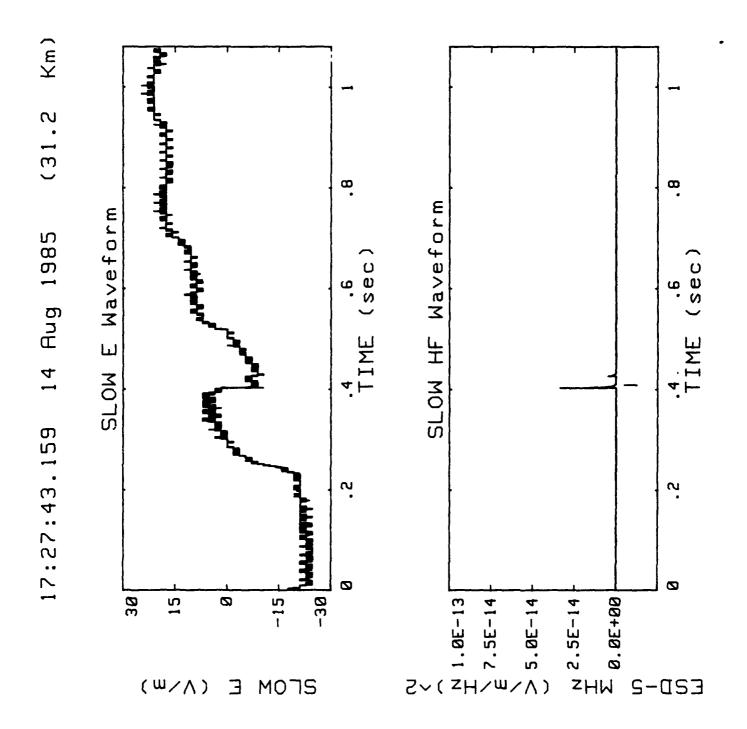


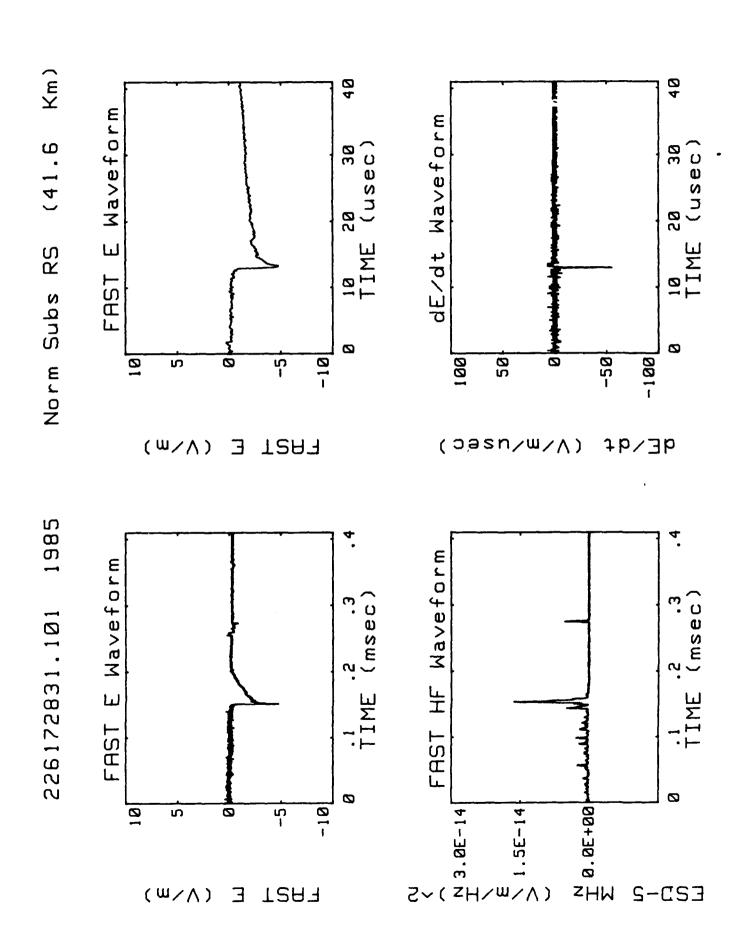


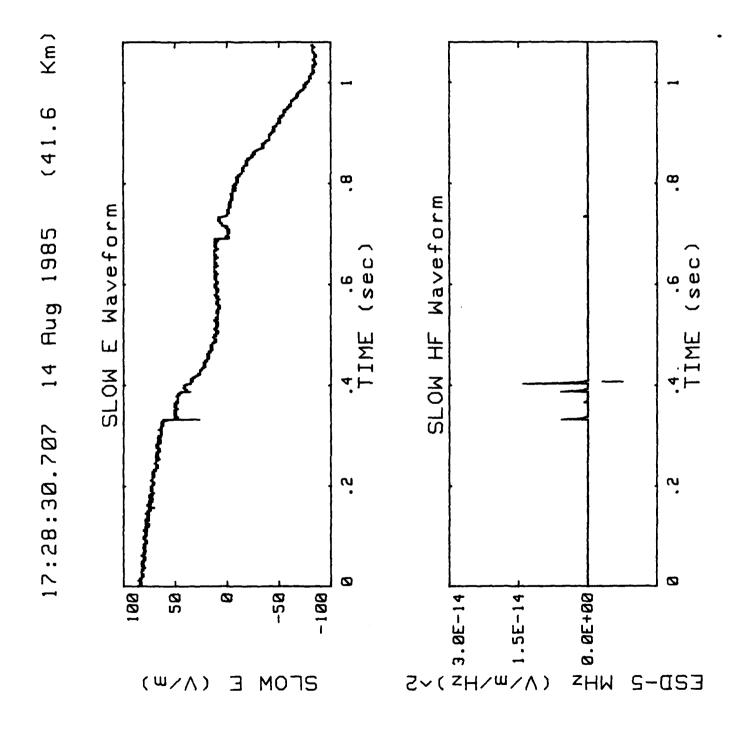


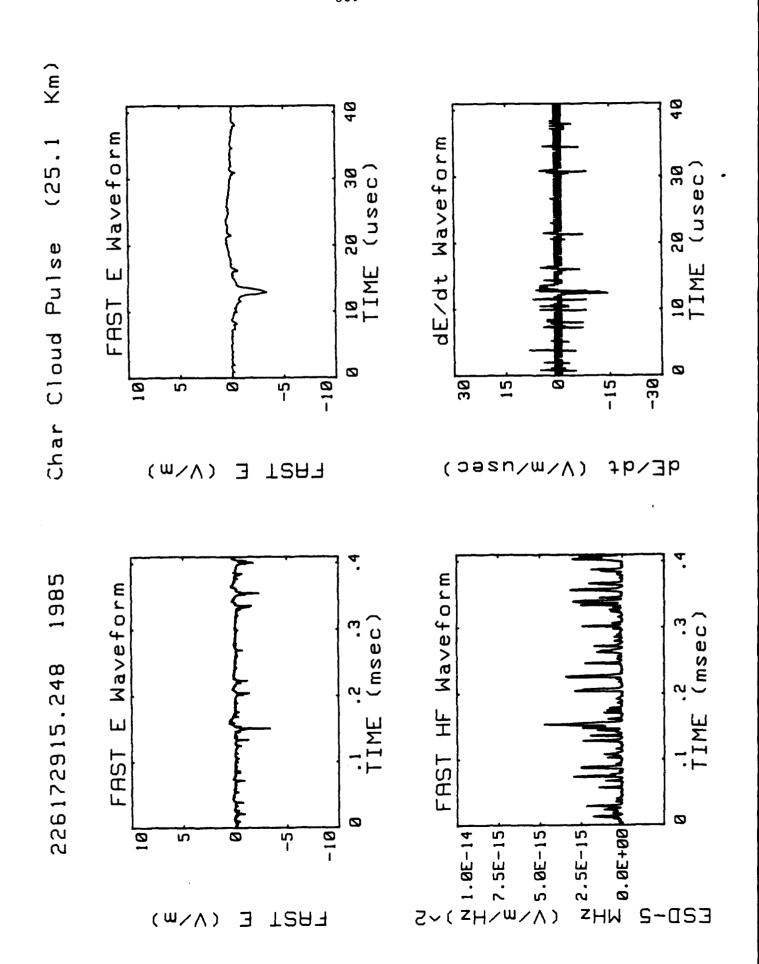


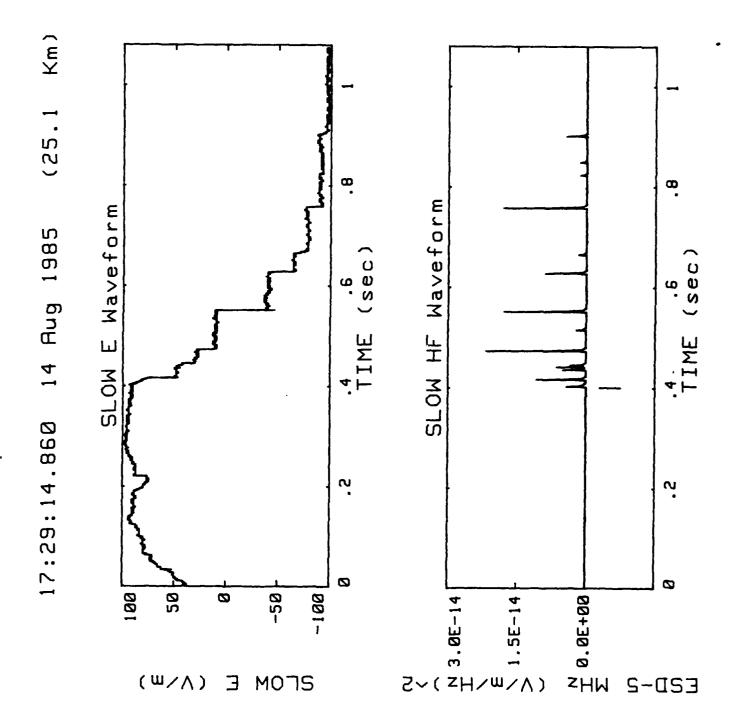


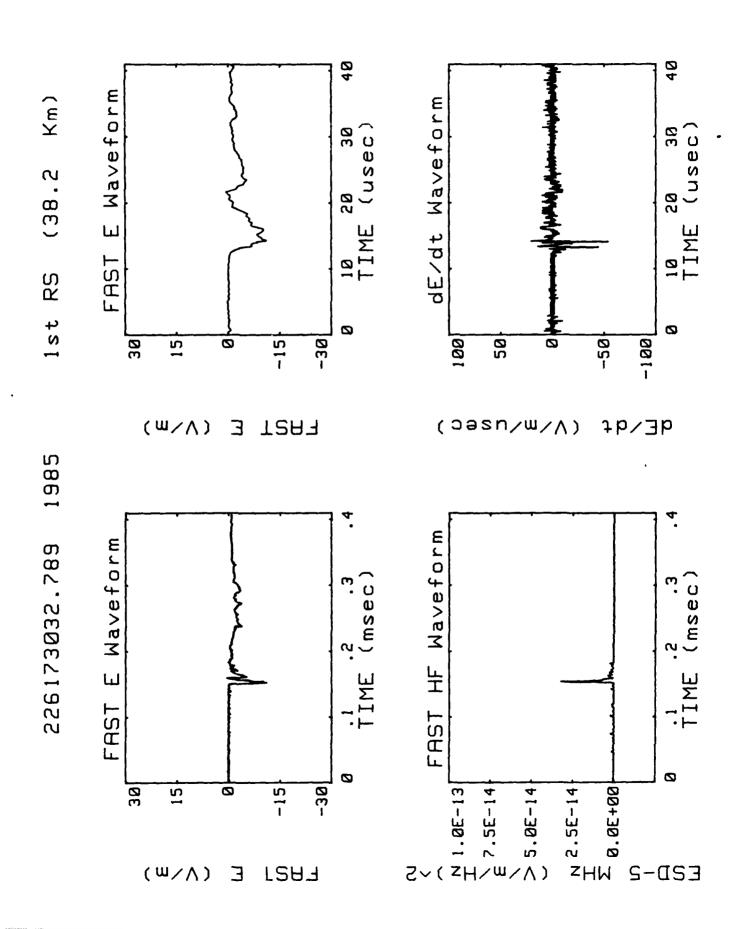


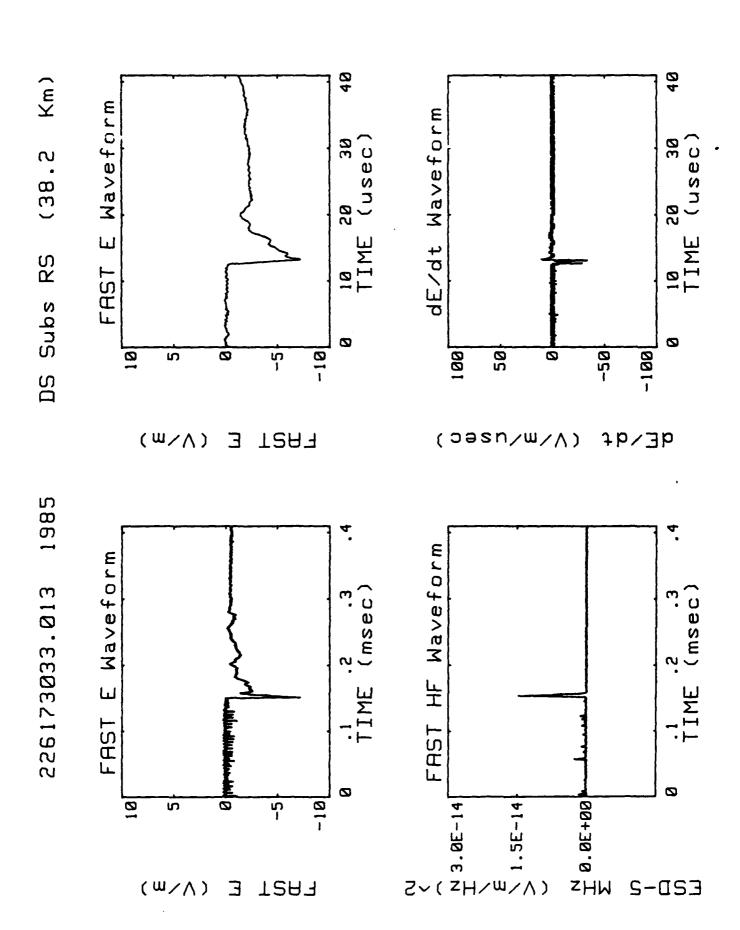


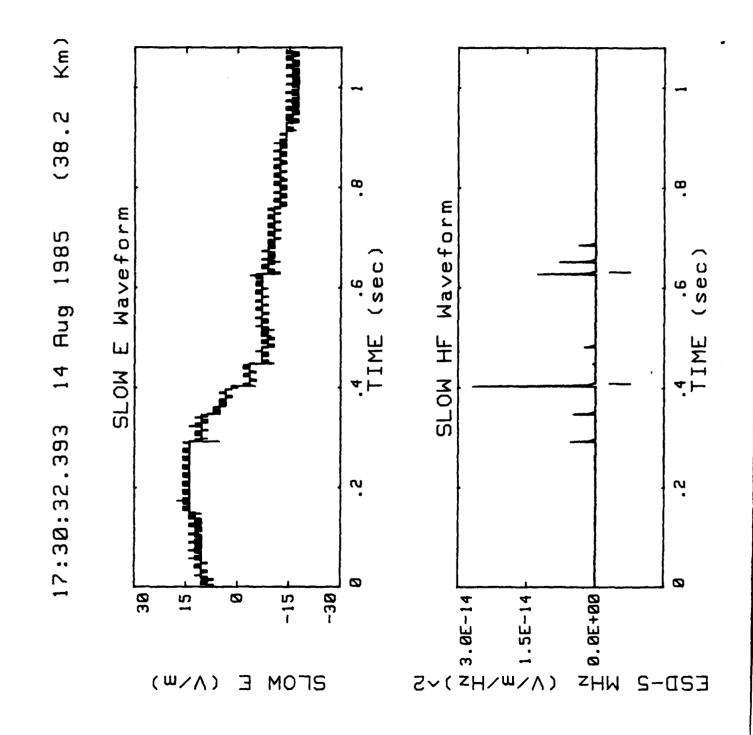


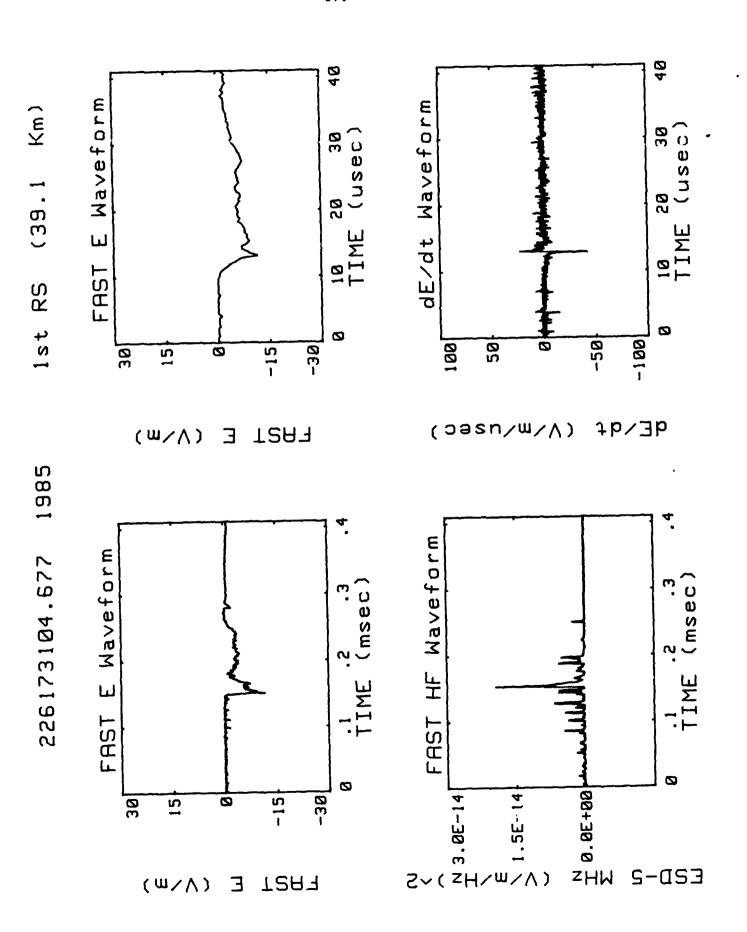


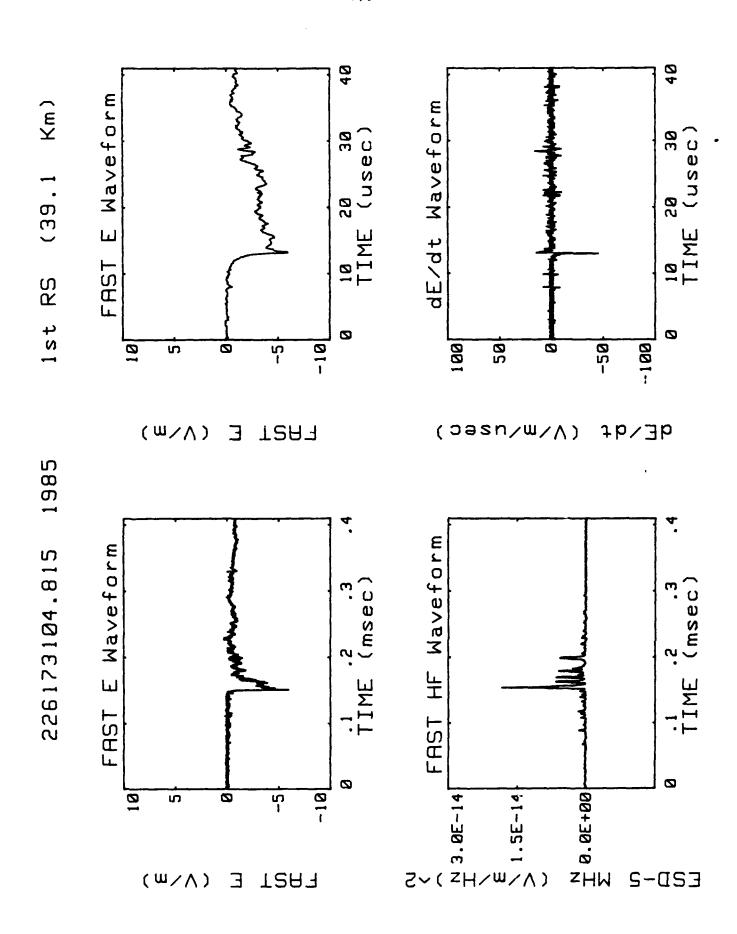


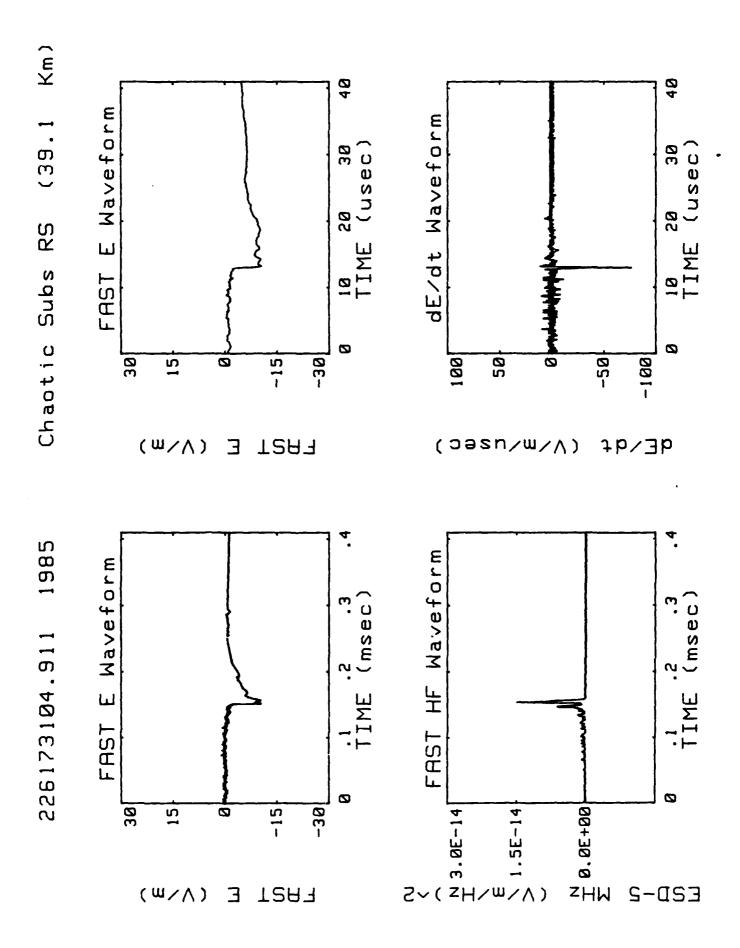


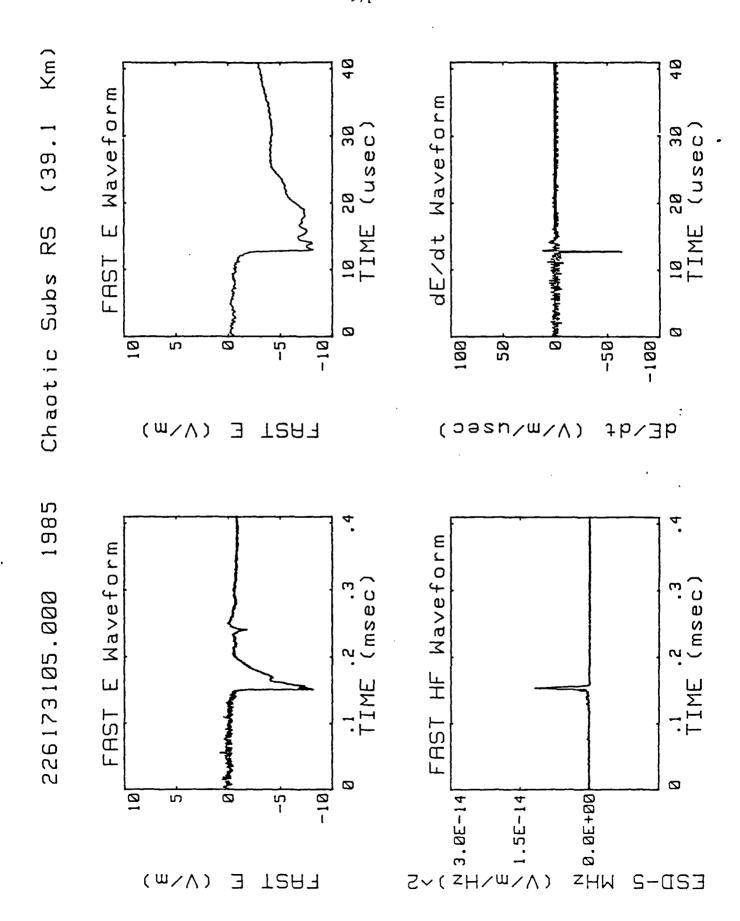


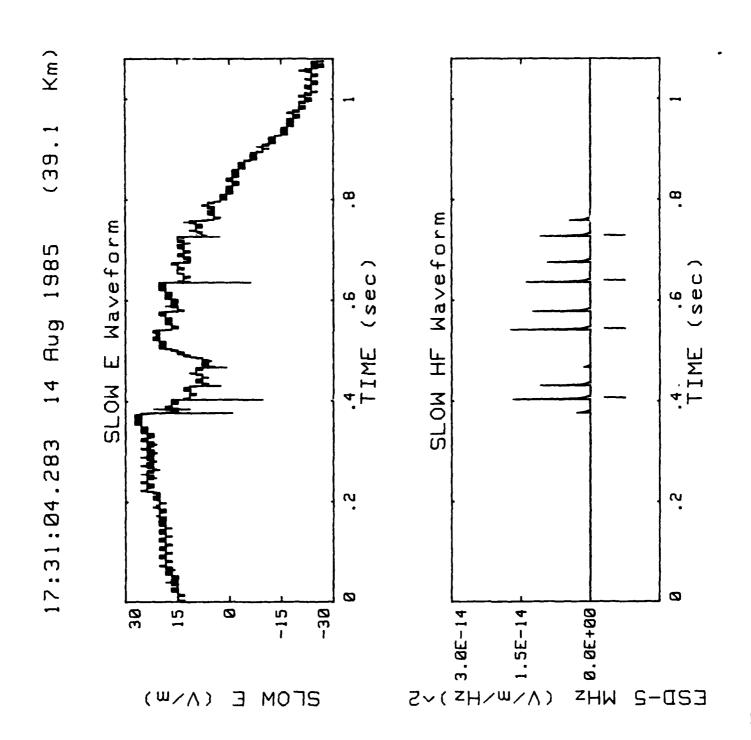


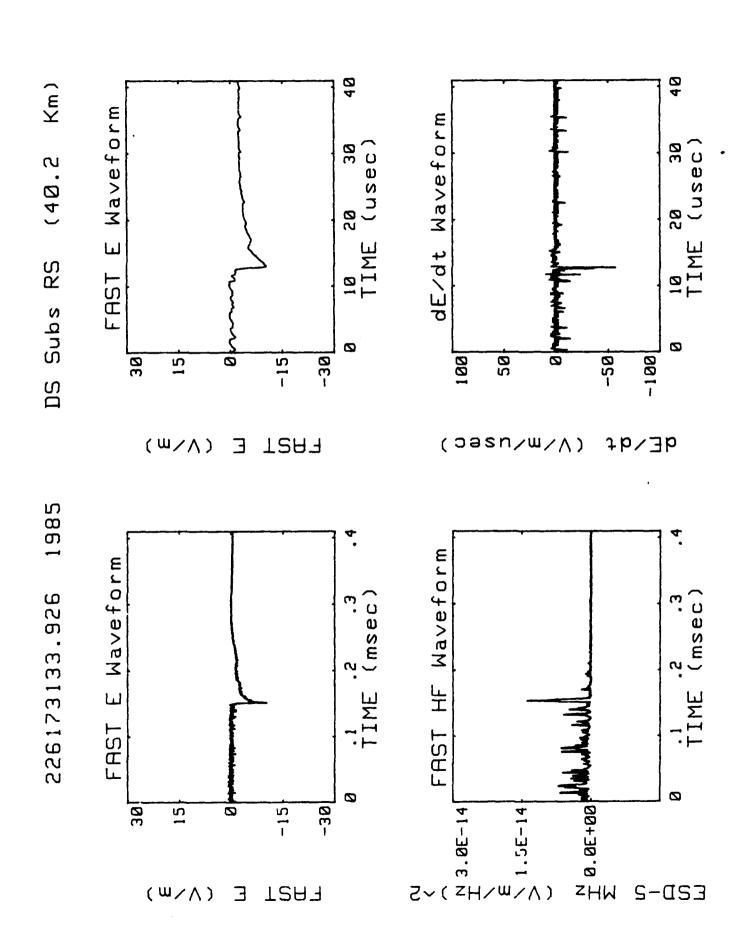


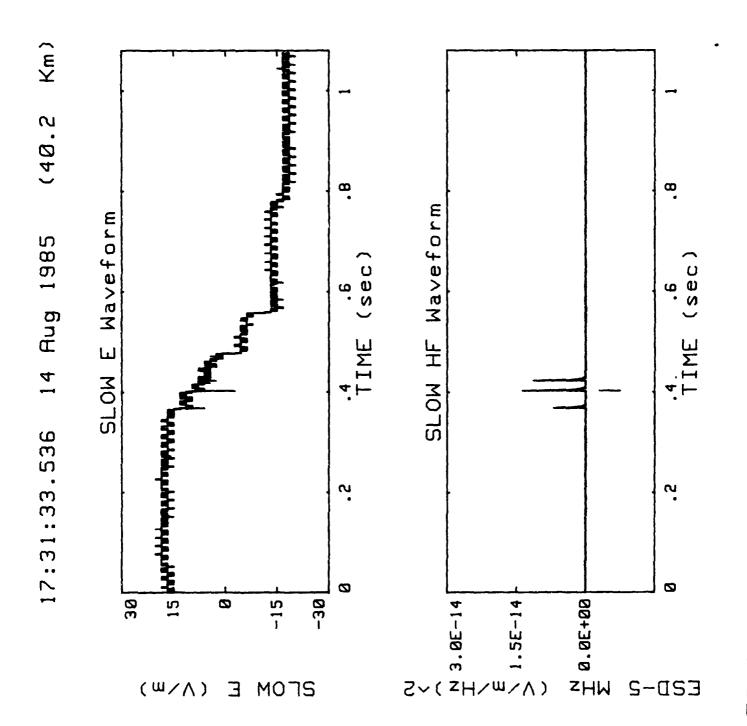


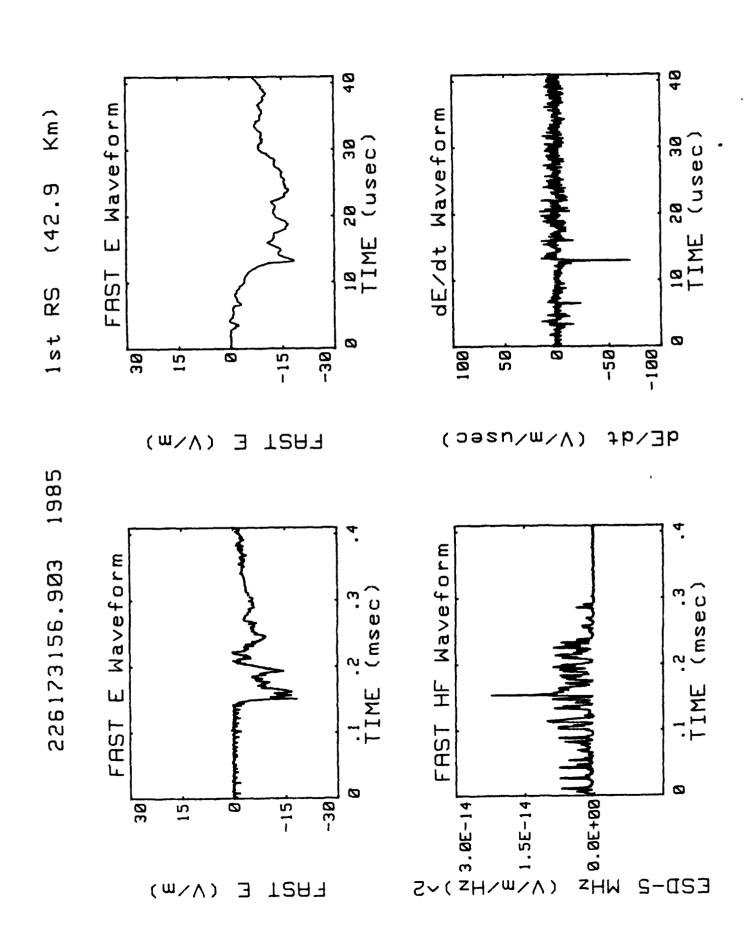


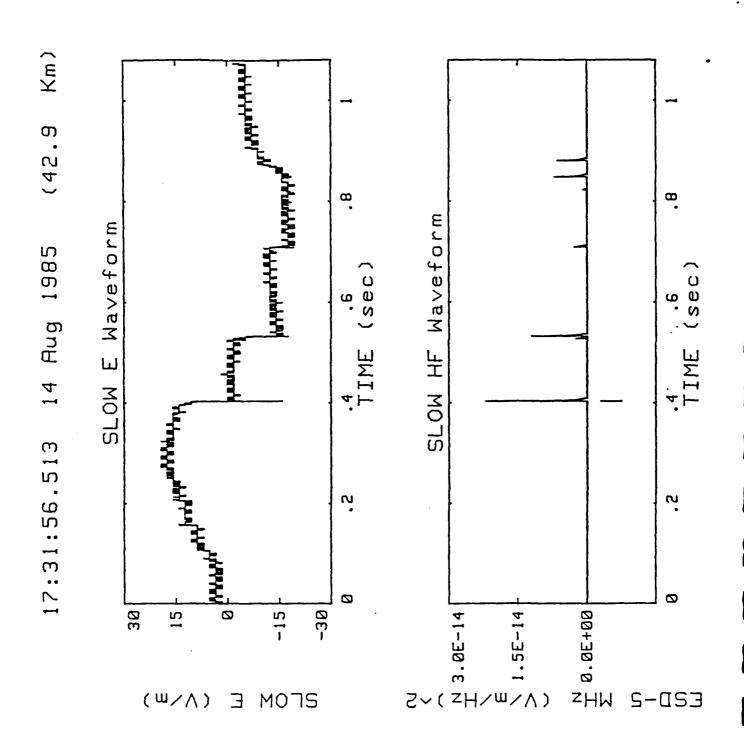


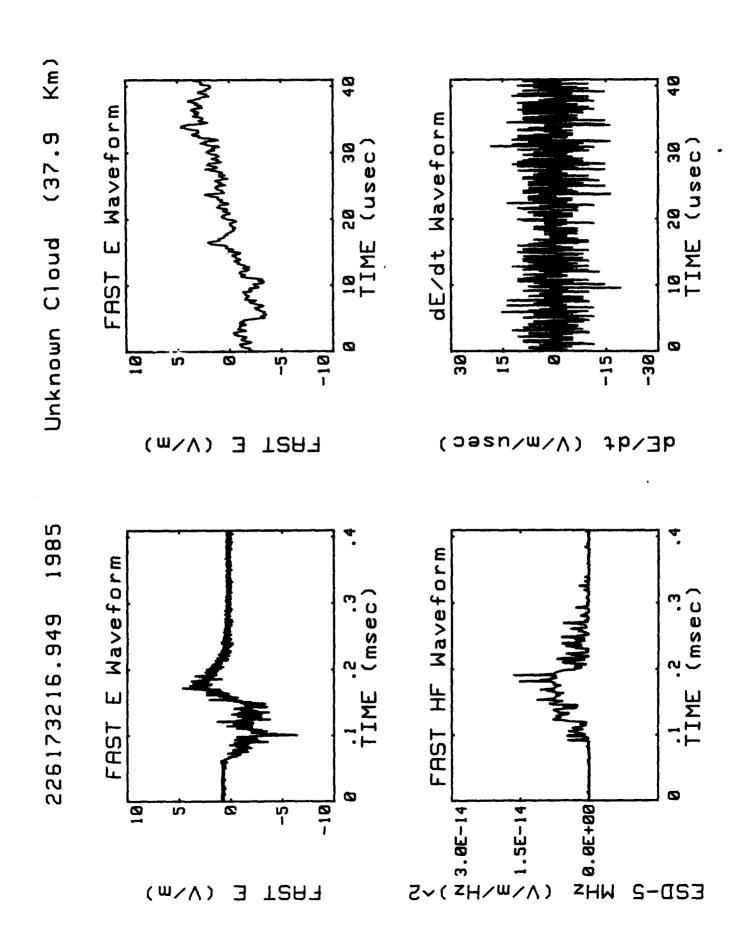


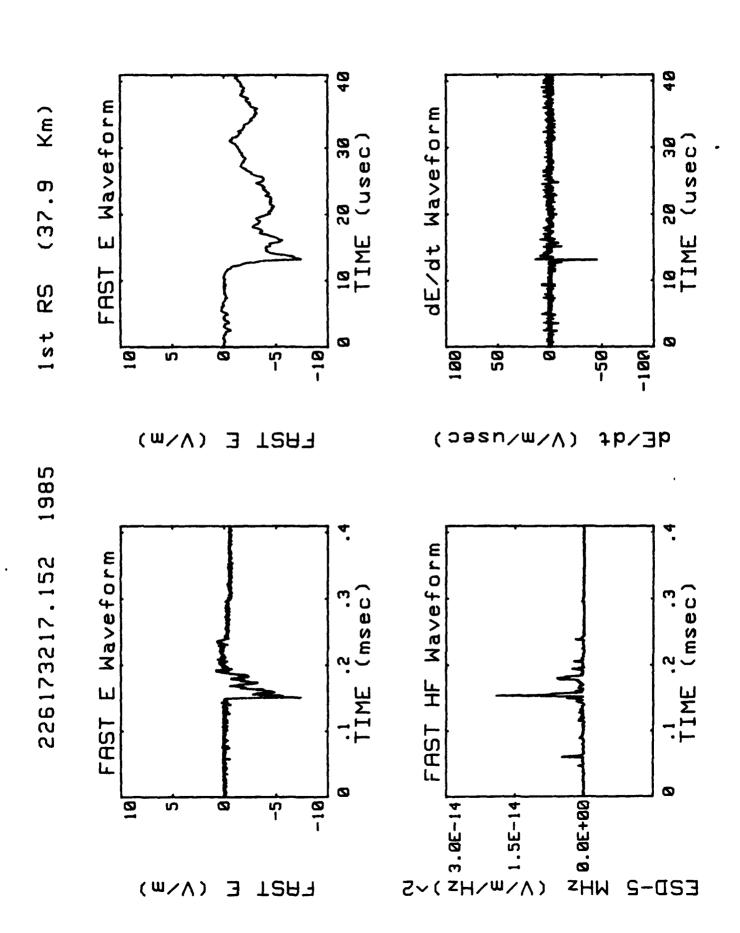


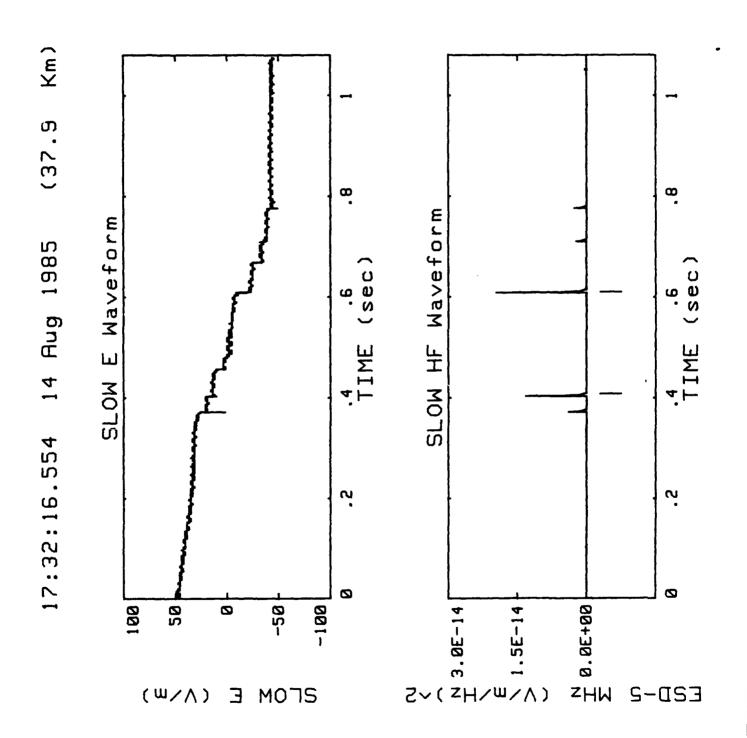


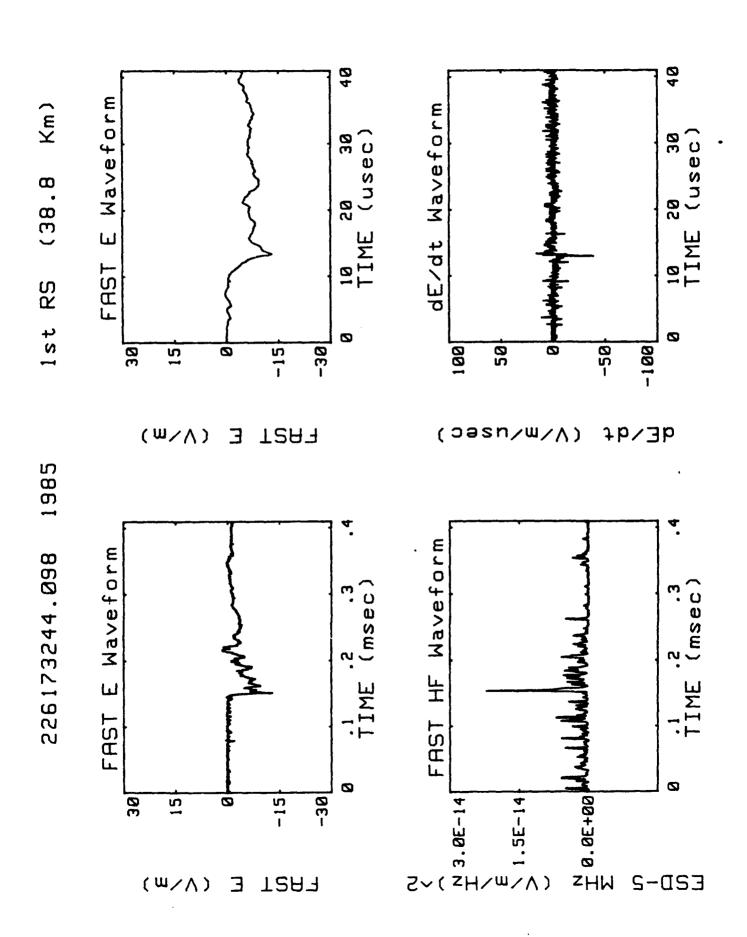


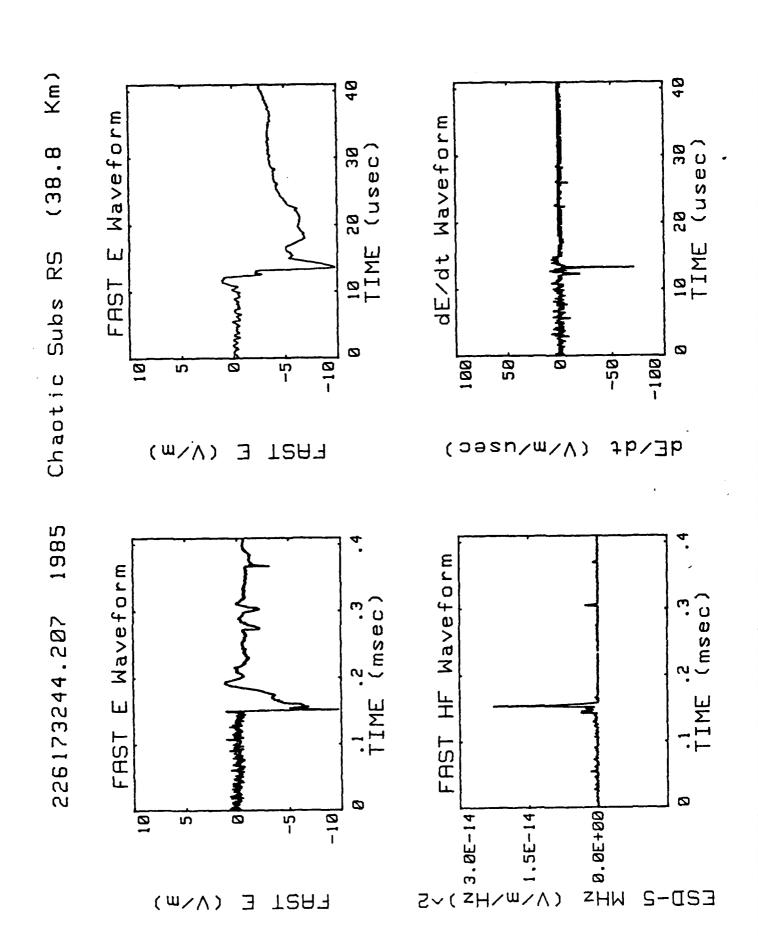


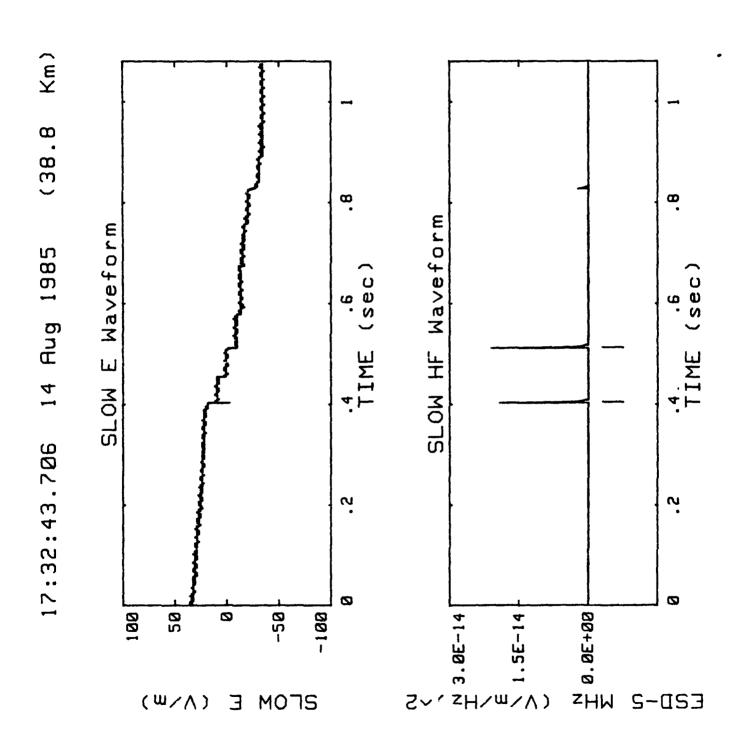


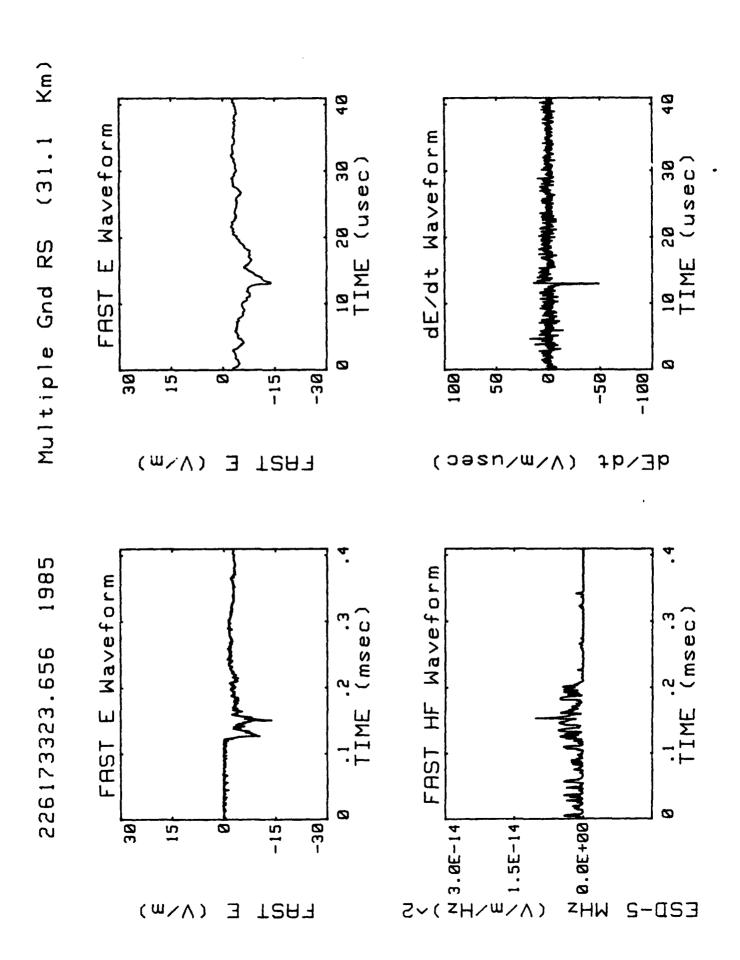


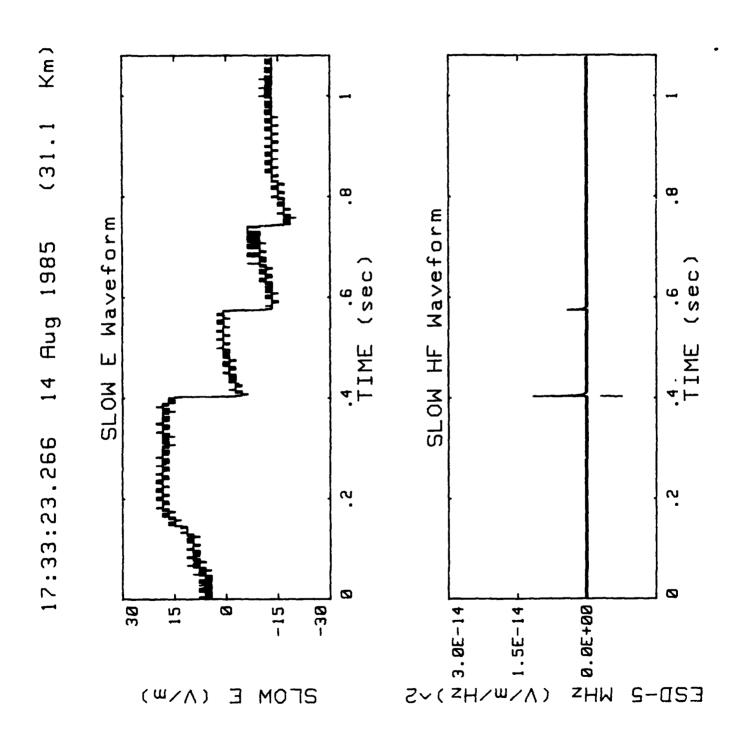


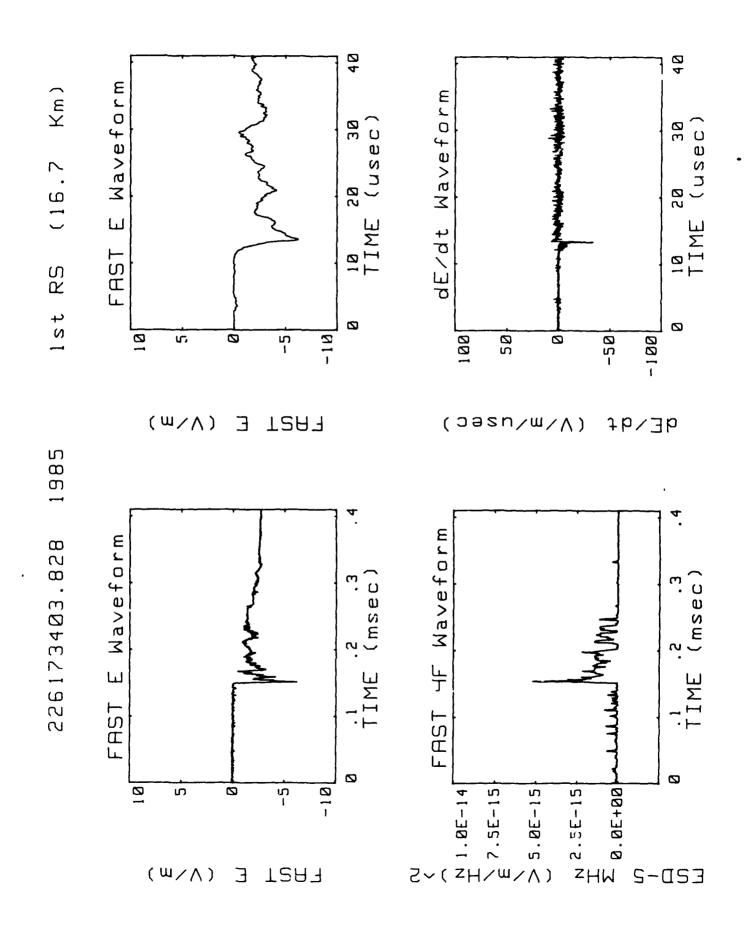


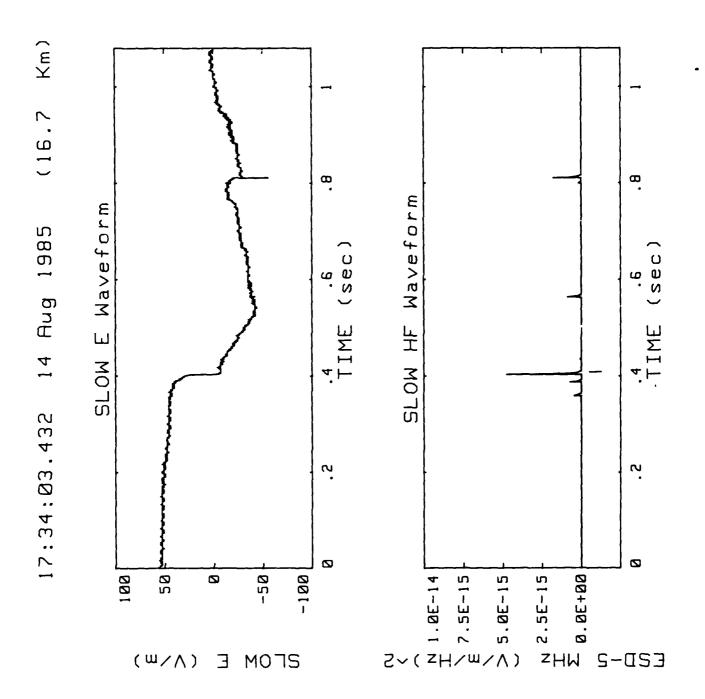


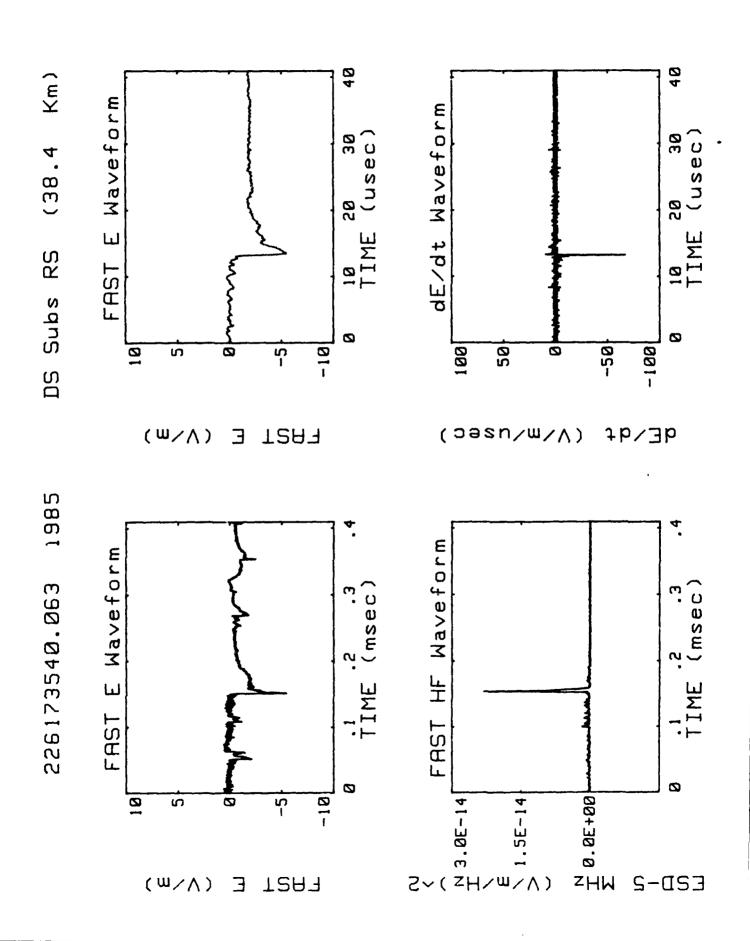


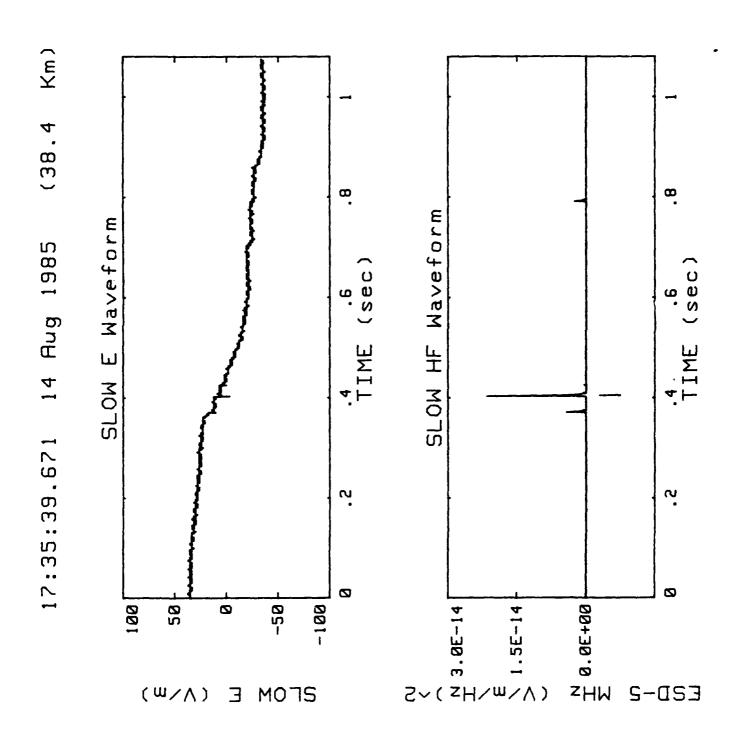


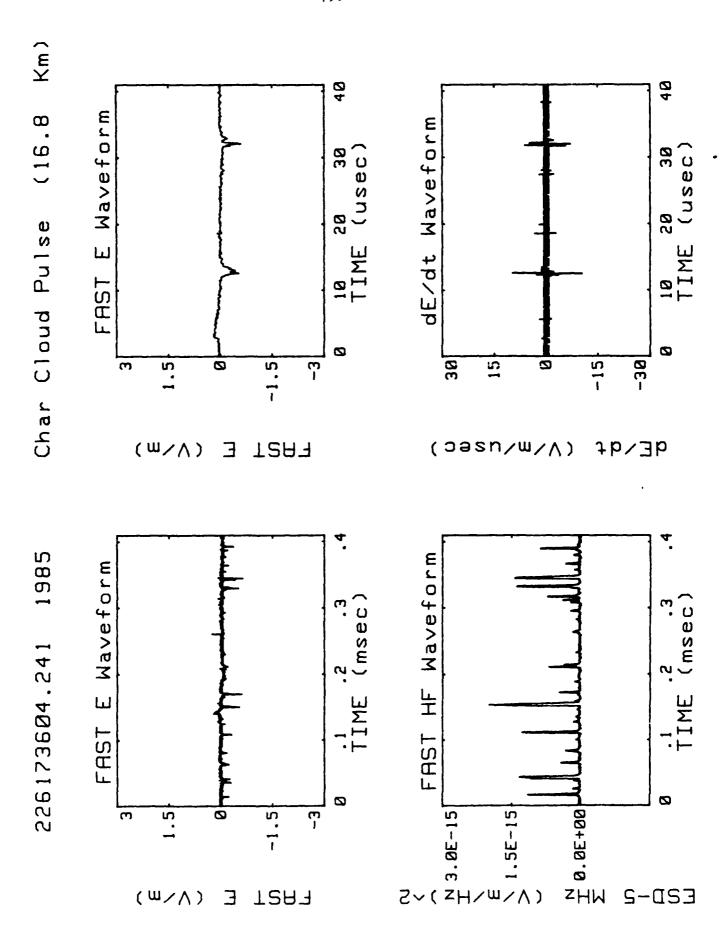


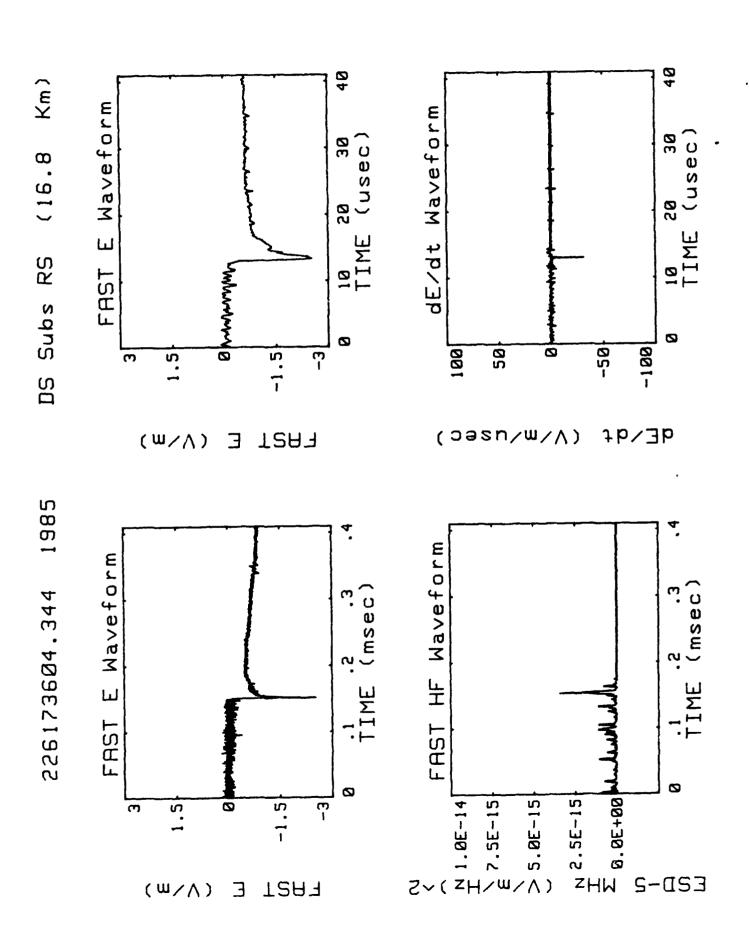


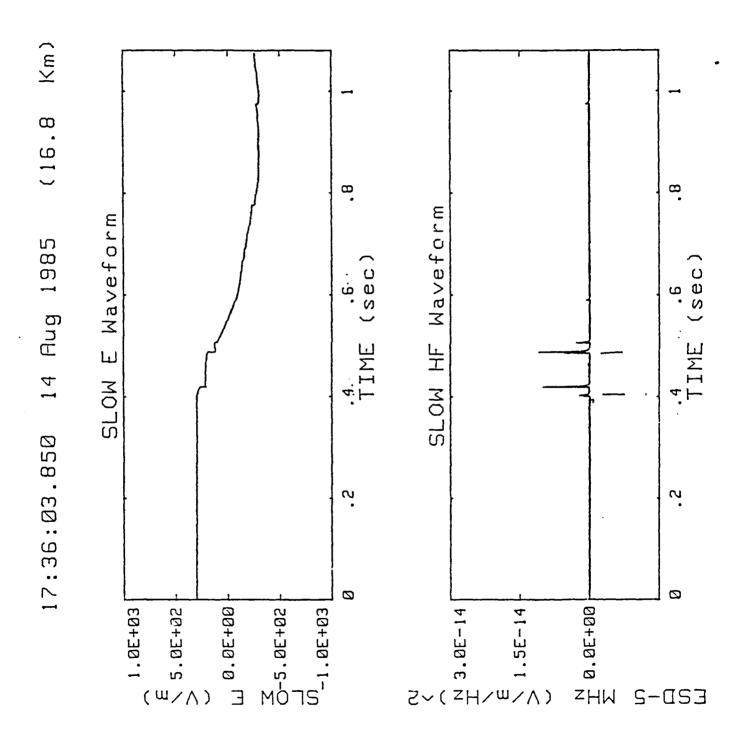


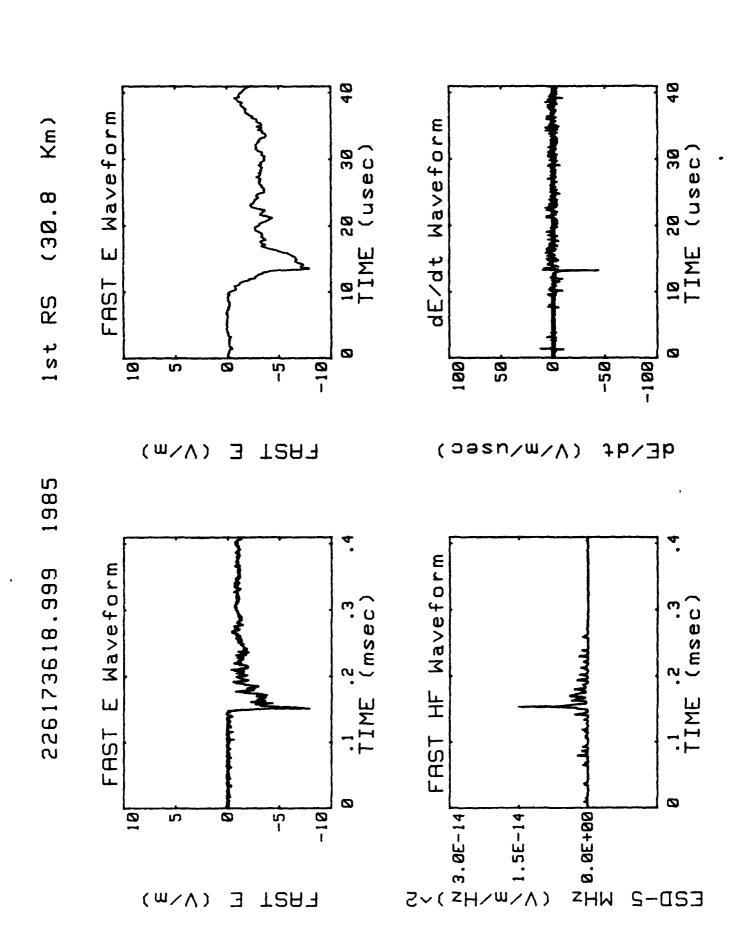


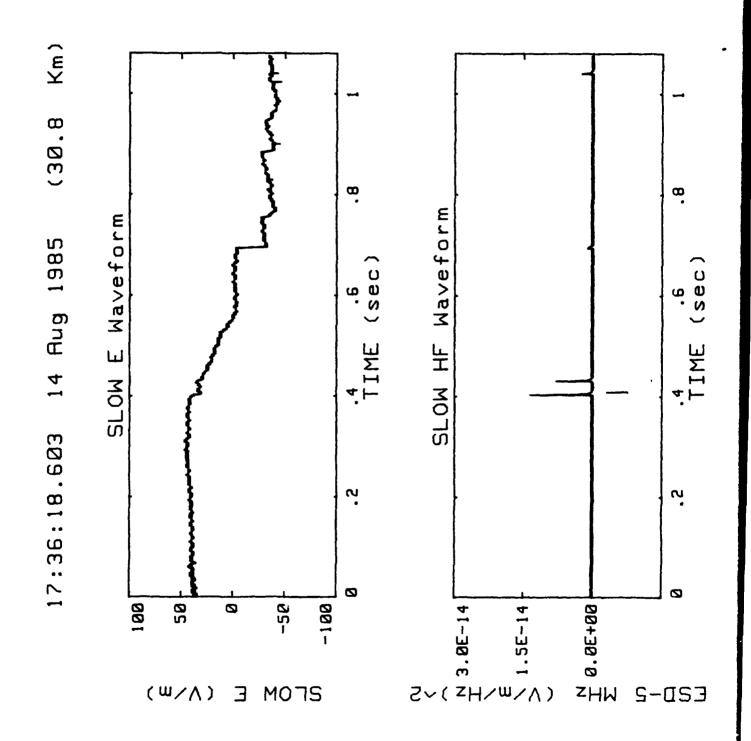


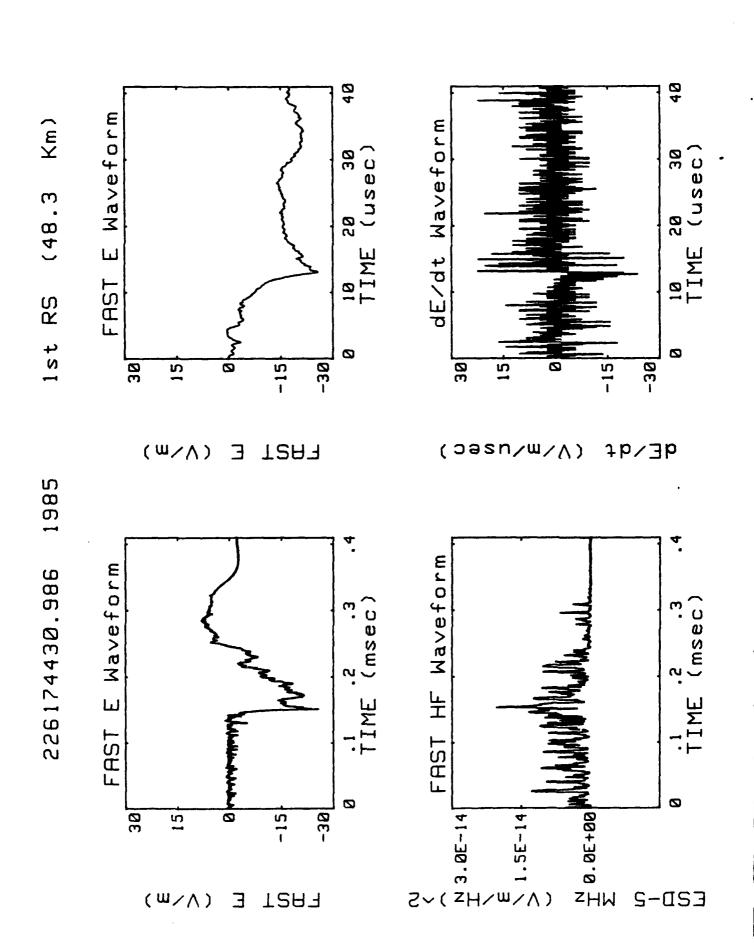


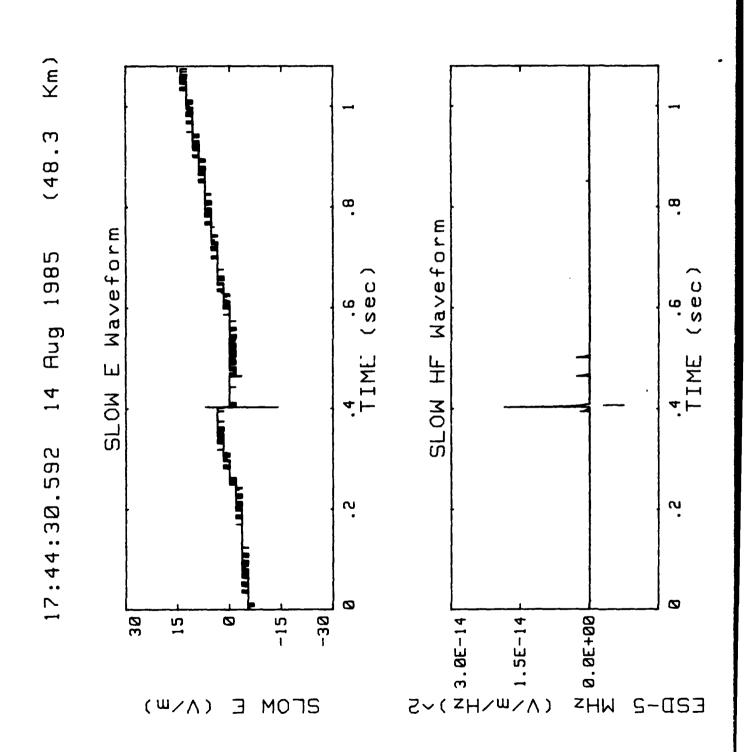


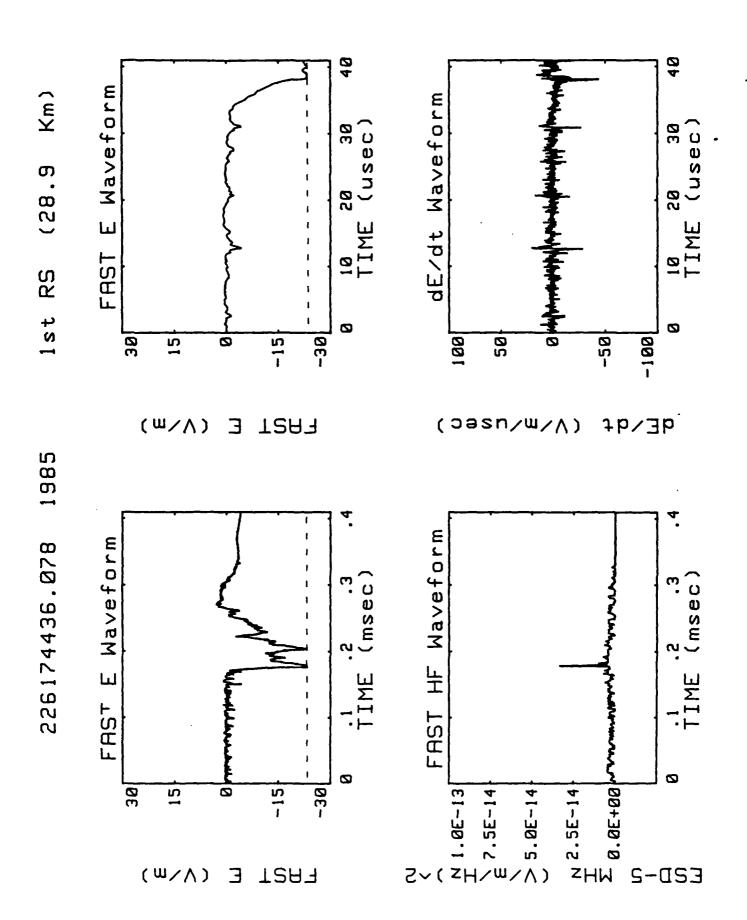


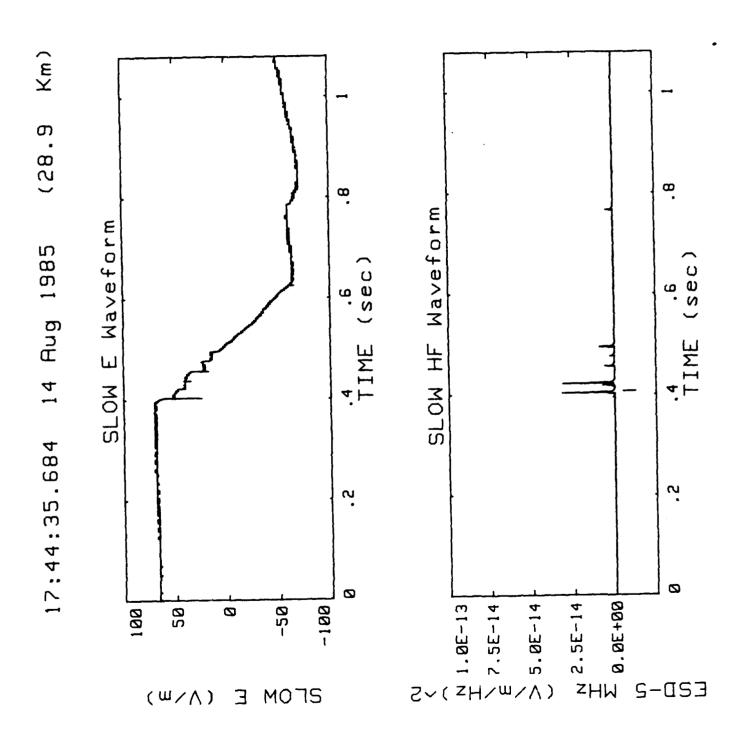












X 40 (17.5)Waveform Waveform 10 20 30 TIME (usec) 10 20 30 TIME (usec) Char Cloud Pulse Ш dE/dt FAST 0 -30 -1.5 1.5 -3 30 15 ന 0 -15 Э FAST (W//) (\\m\user\m\\u) dE/dt 1985 Waveform Waveform .2 .3 (msec) .1 .2 .3 TIME (msec) 226174504.511 TIME 上 ليا FAST FAST N 0.0E+00 -1.5 -3 c 1.5 FAST (M/N)

S-OS3

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